
CMS ECAL Cosmic Calibration

Jason Haupt

University Of Minnesota

On behalf of the CMS ECAL collaboration



Compact Muon Solenoid (CMS)



UNIVERSITY OF MINNESOTA

SUPERCONDUCTING COIL

CALORIMETERS

ECAL $e \gamma$
Scintillating
 $PbWO_4$ Crystals

HCAL
Plastic scintillator/
brass sampling

IRON YOKE

TRACKER
Silicon Micro Strips
Pixels

MUON BARREL

Drift Tube Chambers (**DT**) Resistive Plate Chambers (**RPC**)

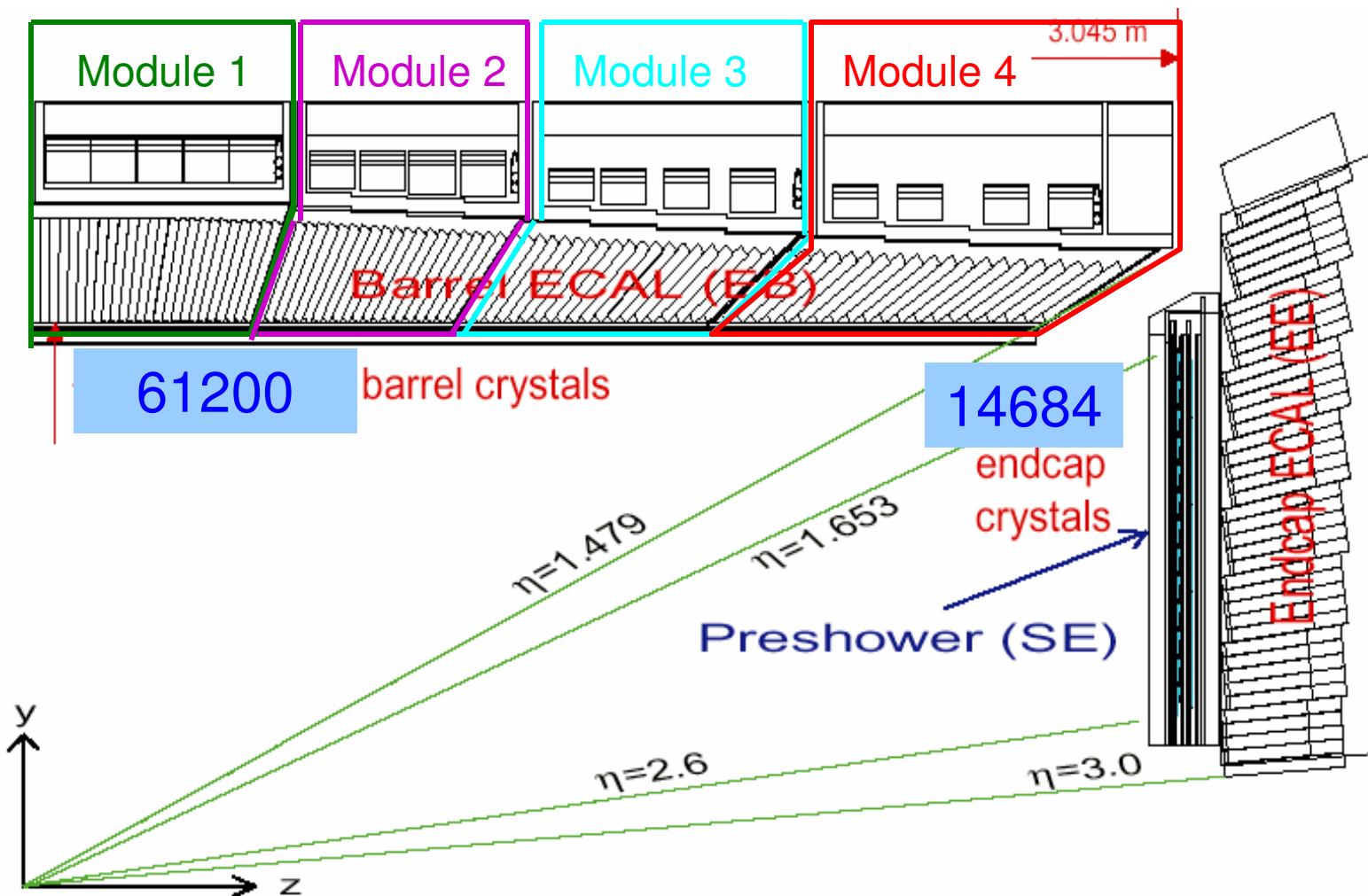
4 T magnetic Field

Total Weight 12,500 t

Overall Diameter 15m

Overall Length 21.6m

ECAL





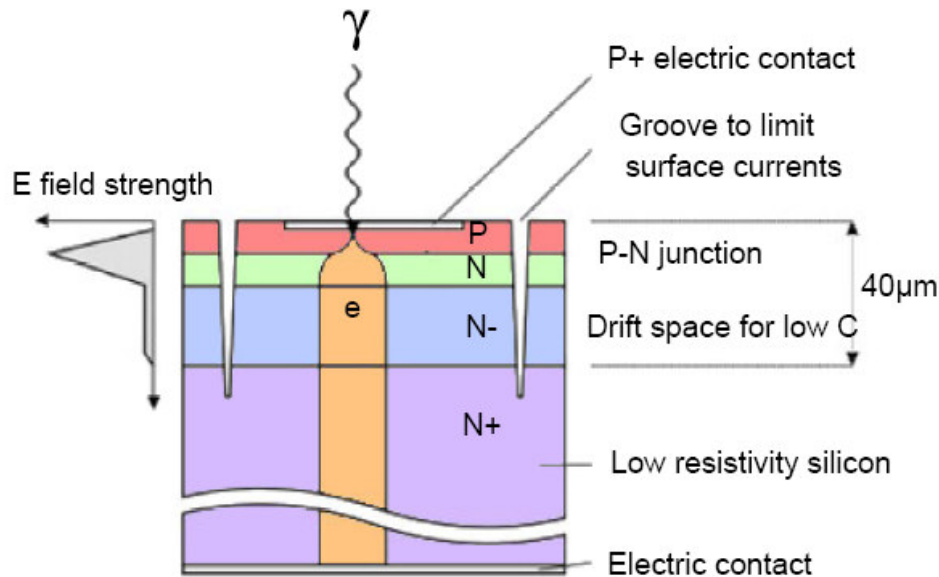
PbWO₄ Crystals



UNIVERSITY
OF MINNESOTA

- Lead Tungstate Crystals
 - Molière radius 2.2cm
 - Radiation Length 0.89cm
 - Scintillation decay time 80% at 35ns
 - Been shown to be radiation resistant
 - Low LY compared to other commonly used crystals (CsI BGO ...)
 - -1.9%/°C temp dependence
- Lead Tungstate Crystals in CMS (Barrel)
 - “Average size”, 2.4x2.4cm² and 23cm in length
 - 34 Different crystal shapes
 - 25.8 X₀
 - 36 Supermodules make up the barrel calorimeter
 - Each Supermodule has 1700 crystals

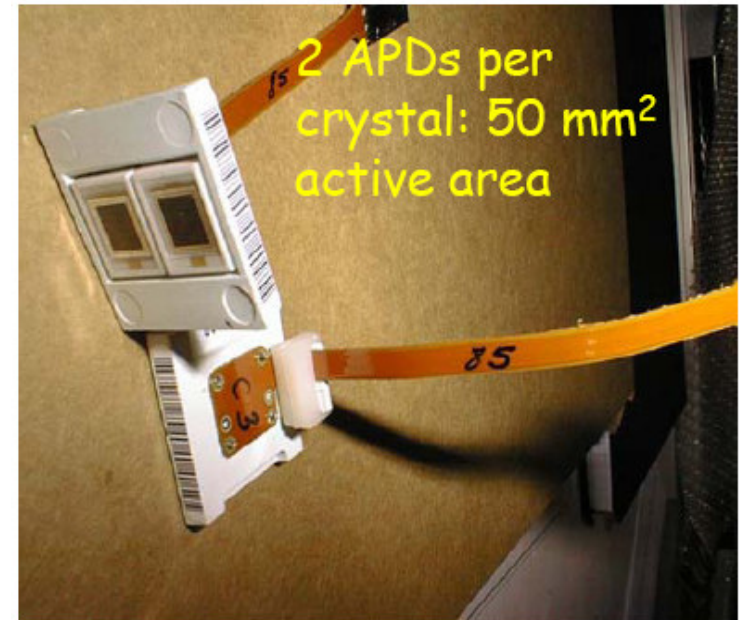
ECAL Barrel Optical Readout



≈ 4.5 photo-electrons/MeV

122400 Total APD's

Very Linear Devices



- ❑ Two 5x5 mm² APD's/crystal
- ❑ Gain – 50
- ❑ QE – 75% @ 420 nm
- ❑ Temp sensitivity – -2.4%/°C



ECAL Inter-calibration Goals

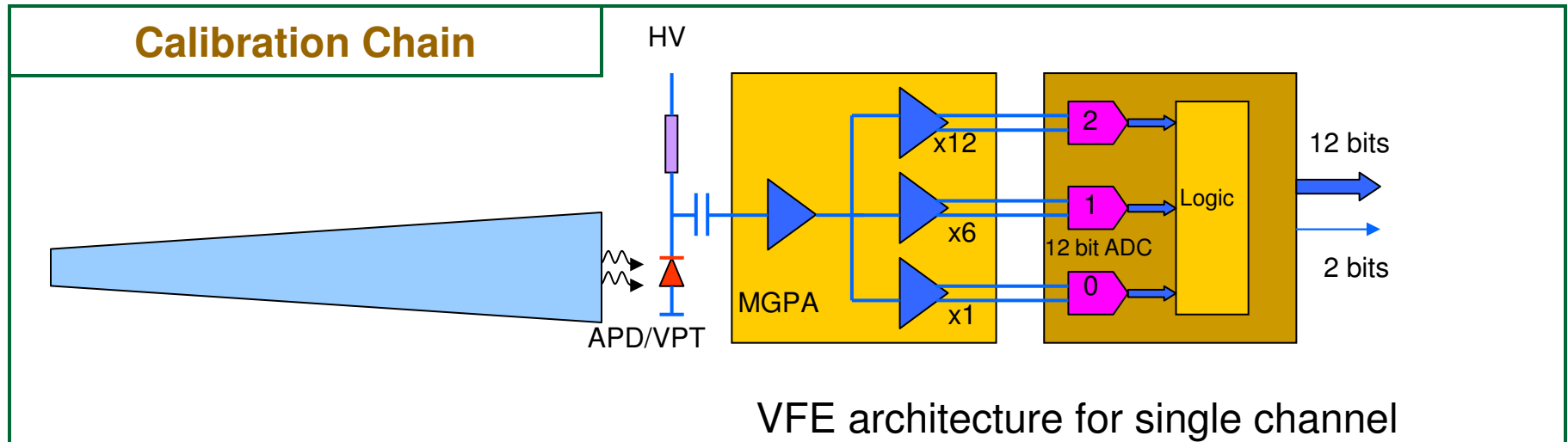


UNIVERSITY
OF MINNESOTA

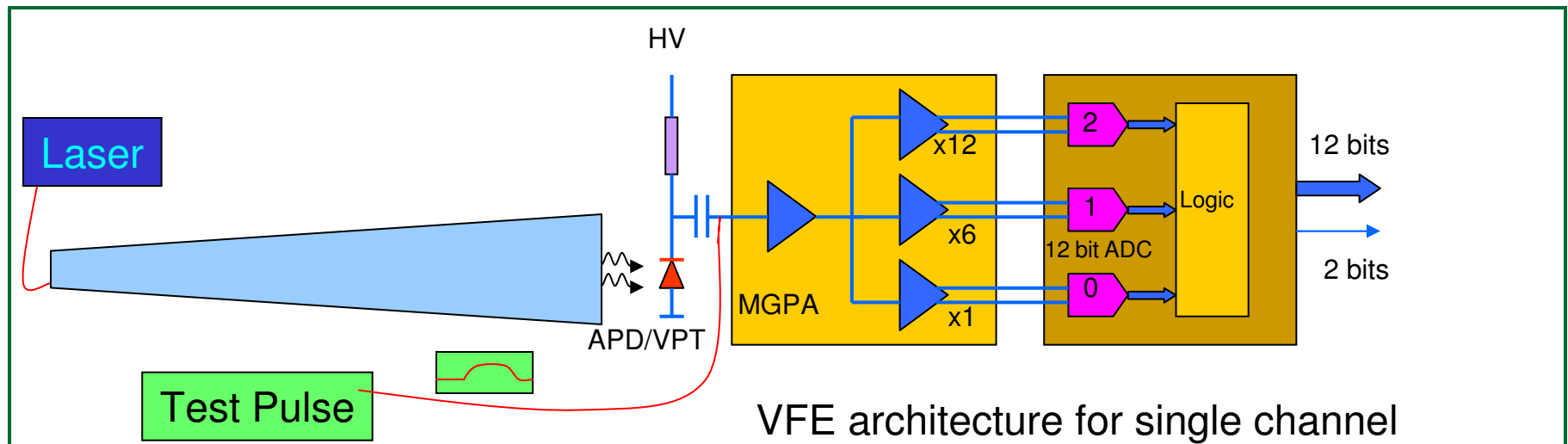
- Energy Resolution $(\sigma_E / E)^2 = (a / \sqrt{E})^2 + (b / E)^2 + c^2$
- Goal : constant term “c” < 0.5% → $\sigma/E < 0.5\%$ (For High Energies)
- *In-situ* Calibrations
 - $Z \rightarrow e^+e^-$ ~1 day 1% (with ϕ ring inter-calibration)
 - $W^\pm \rightarrow e^\pm\nu$ ~2 months E/p from Tracker
 - $\pi^0 \rightarrow \gamma\gamma$, $\eta^0 \rightarrow \gamma\gamma$, etc.
- Initial inter-calibrations
 - LY ~4%
 - Cosmics ~2%
 - Test Beam ~0.3% (Not available for all SM's)
- Reason for pre-calibration
 - Uniform detector response at startup

Calibration Chain

- Crystal Energy → ADC count
 - Crystal optical response
 - APD Gain
 - Amplifier Gain
 - ADC 12bit Out



- Cross Checks
 - Test Pulse (after APD)
 - Compared to previous test pulses
 - Laser allows for self referencing
 - Compare one laser run to another





Laboratory Inter-Calibration



UNIVERSITY
OF MINNESOTA

- Two current methods to LY measurements (Basically Quality Checks) automated
 - 1. Direct LY along crystal ^{60}Co
 - ~ 1.2 MeV source
 - 2. Transmission through crystals longitudinally at 360nm
- Combined Laboratory constants
 - Laboratory measurements are combined; LY, APD gain, the preamp.
 - Result of a $\sim 4.0\%$ agreement compared to testbeam calibration constants
 - Comparing ~ 1.2 MeV Source to 120 GeV testbeam!



Cosmic Ray Muons

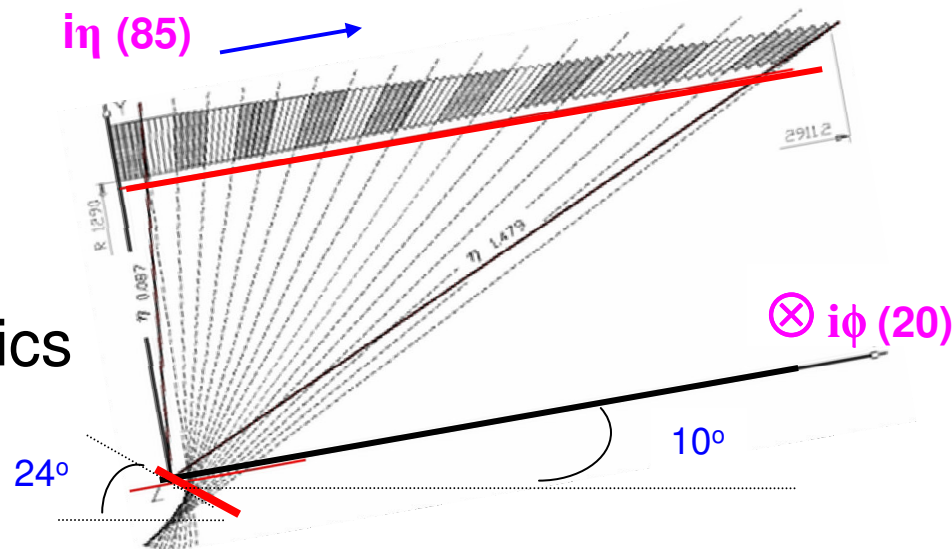


UNIVERSITY
OF MINNESOTA

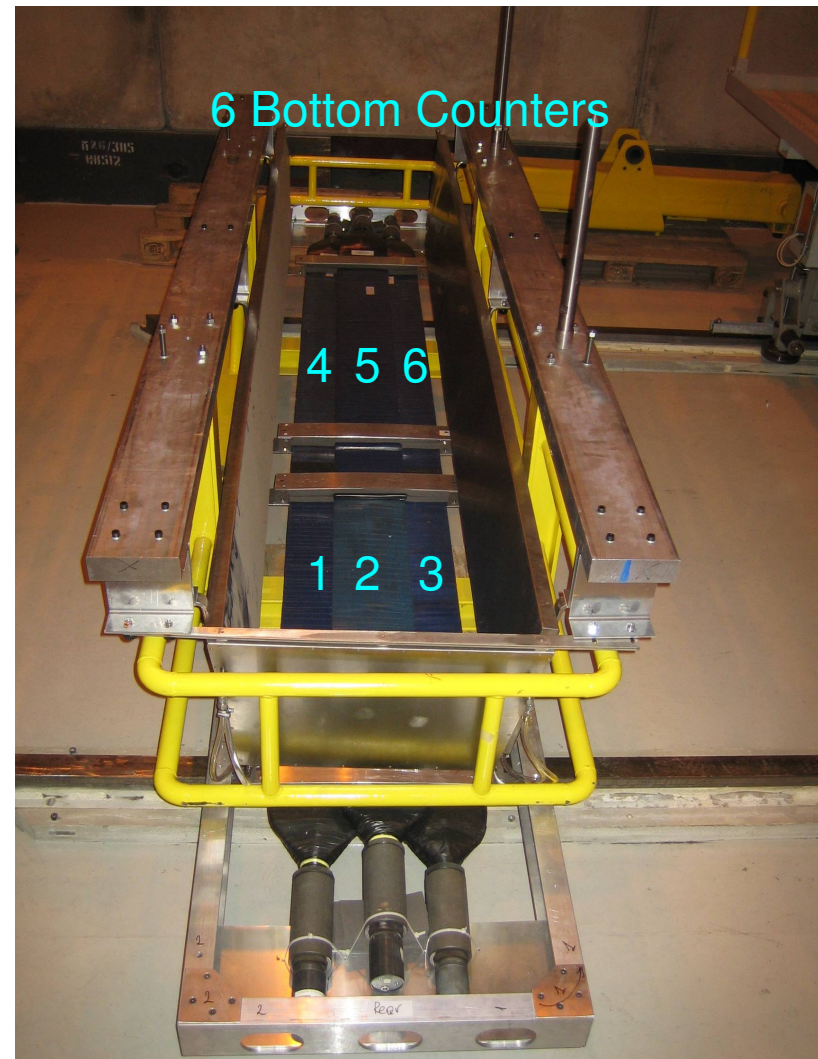
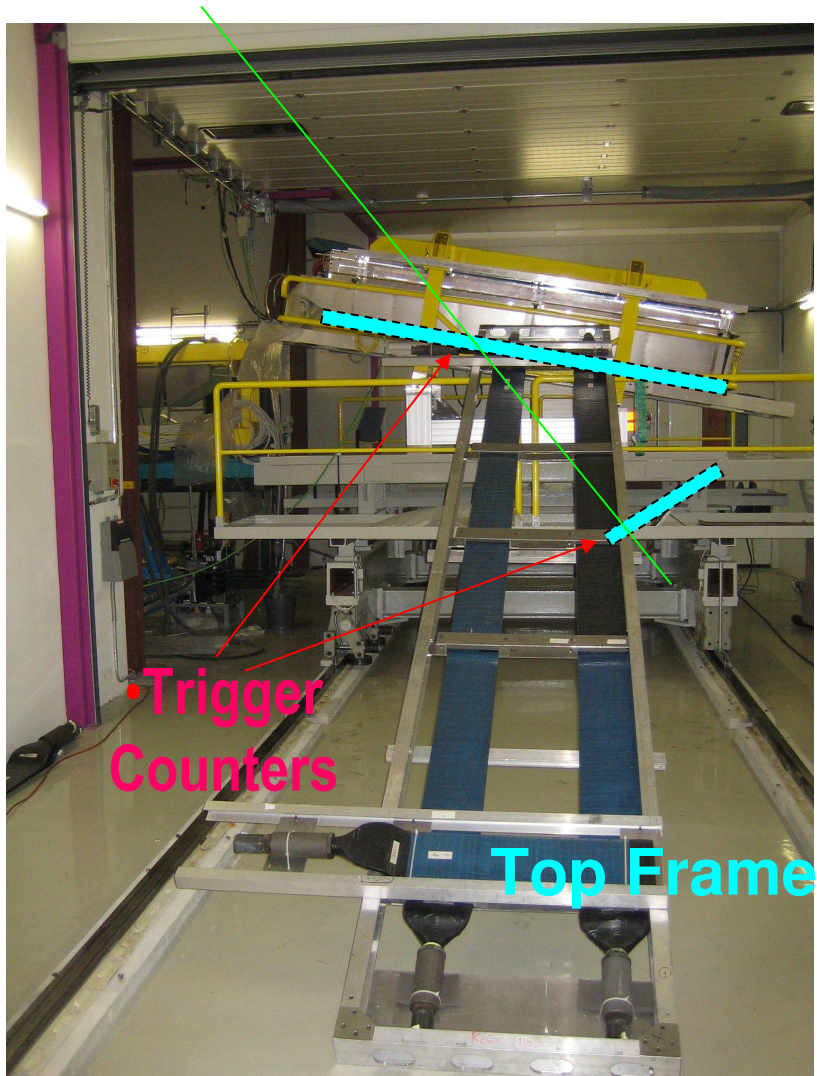
- Cosmic ray muons deposit $\sim 11 \text{ MeV/cm}$ in PbWO_4 crystals
- A through going muon will deposit $\sim 250 \text{ MeV}$ in a 23cm crystal
- APD gain was increased factor of 4 to get away from the $\sim 38 \text{ MeV}$ electronics noise. Actual gain ratio is found with laser. Also to improve the neighbor veto.
- 1 ADC count $\sim 9 \text{ MeV}$
- Goal
 - Improve measurement for all 61,200 barrel crystals with full readout chain
 - Excellent way to run each SM ~ 10 days as a final “burn-in” step

Cosmic Muon Trigger

- The trigger is a coincidence of one of 6 plastic scintillator counters spread over the bottom of the Super-Module with another counter placed near the interaction point.
- Trigger is designed as to select muons that are through-going. In this situation the amount of deposited energy is most well determined.
- 85x20 in \times i ϕ grid
- One SM has more than 34 Million Triggers
- Most have 4-7 M
- 23 SM's have been calibrated with cosmics

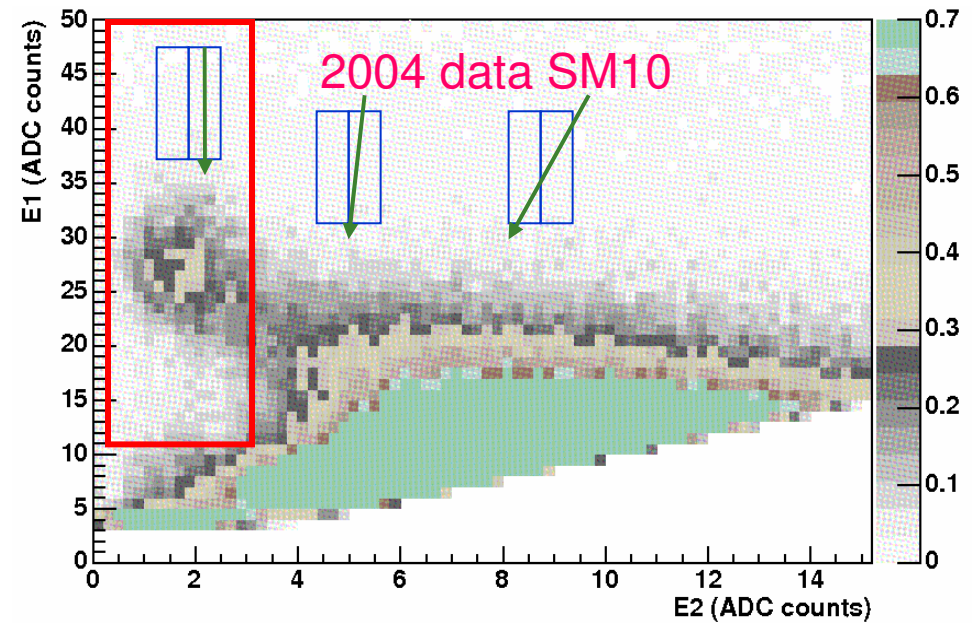
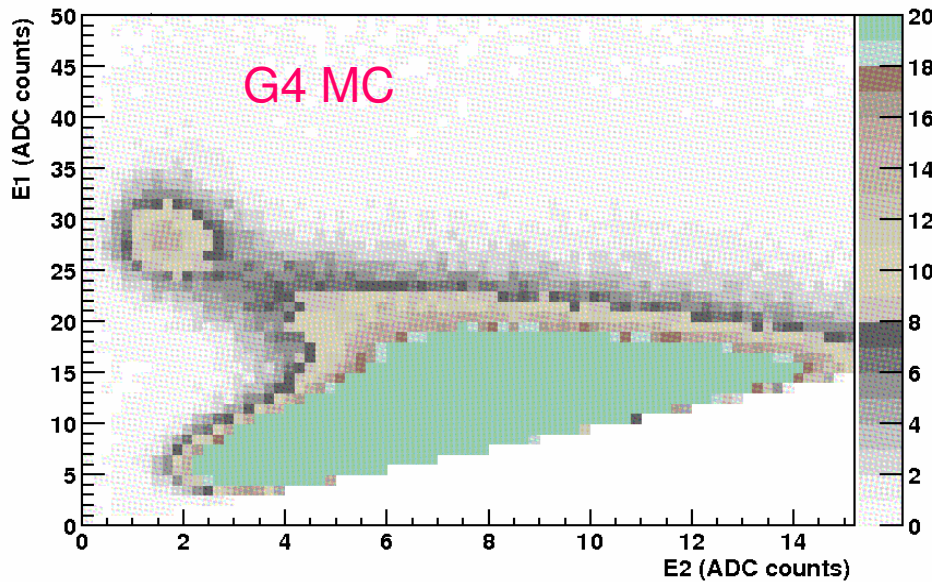
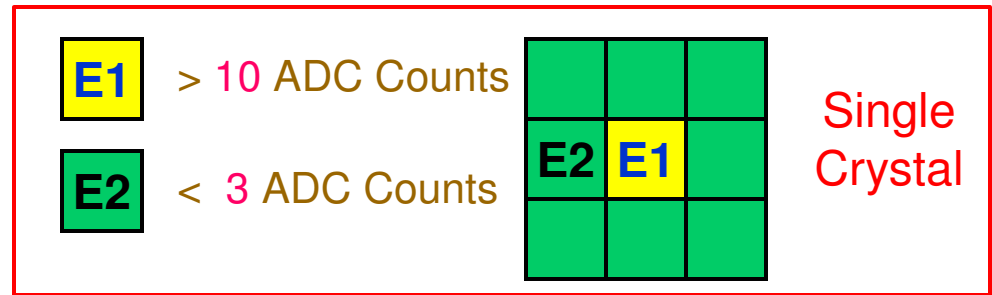


Cosmic Trigger Setup



Crystal Selection

- **E1** is the Energy of the Maximal crystal
- **E2** is the energy of the next highest crystal around the max
- **RED** marks selection



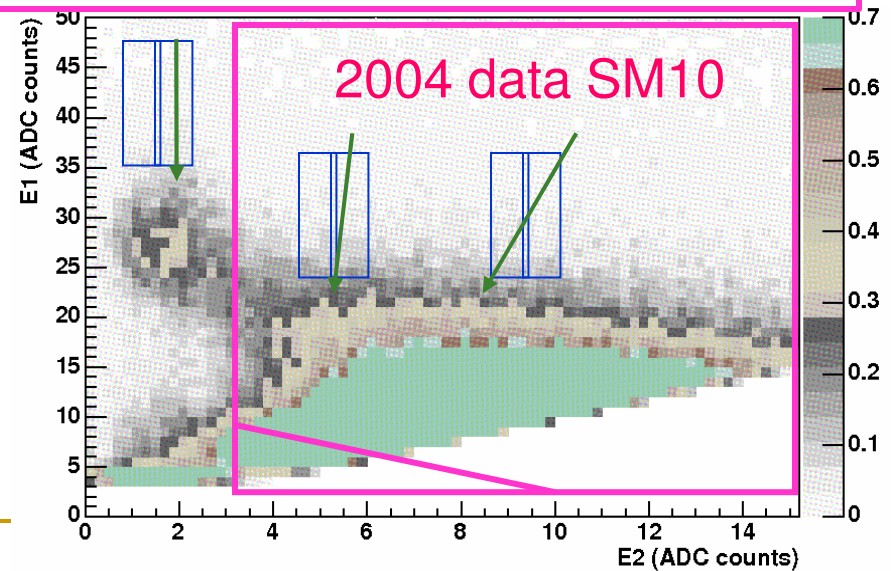
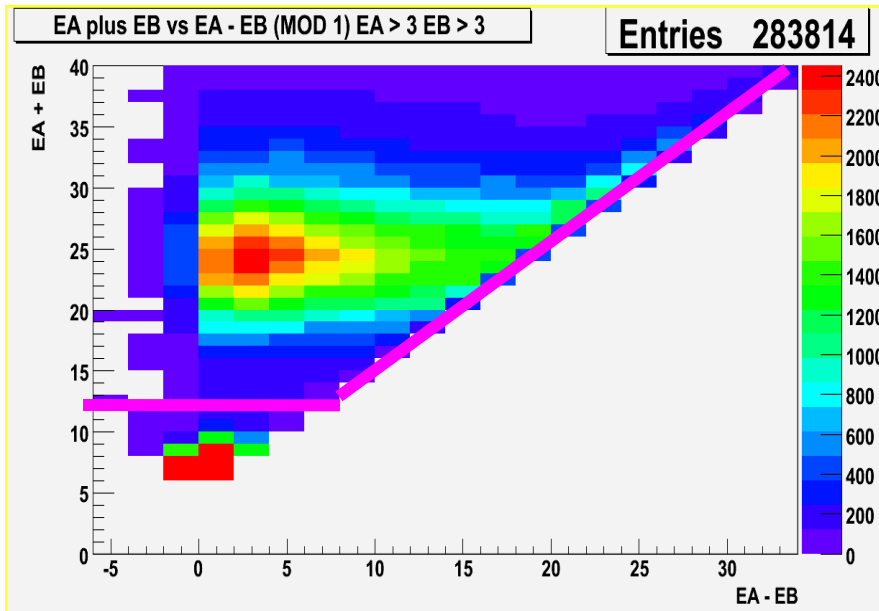
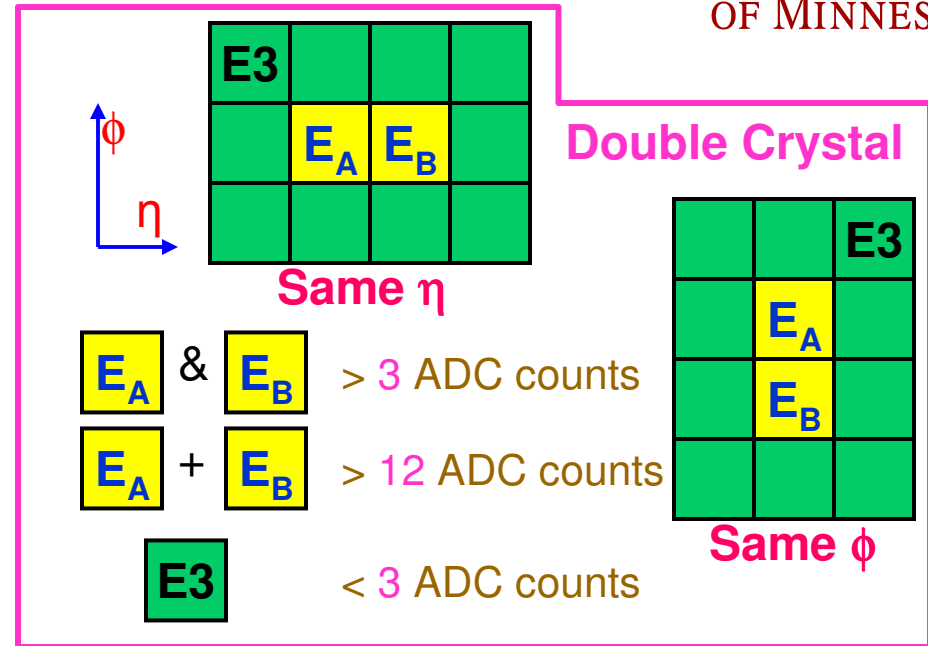


Crystal Selection



UNIVERSITY
OF MINNESOTA

- E_A is the Energy of the Maximal crystal
- E_B is the energy of the next highest crystal around the max
- E3 refers to all other surrounding crystals
- Magenta marks selection



10/28/2006



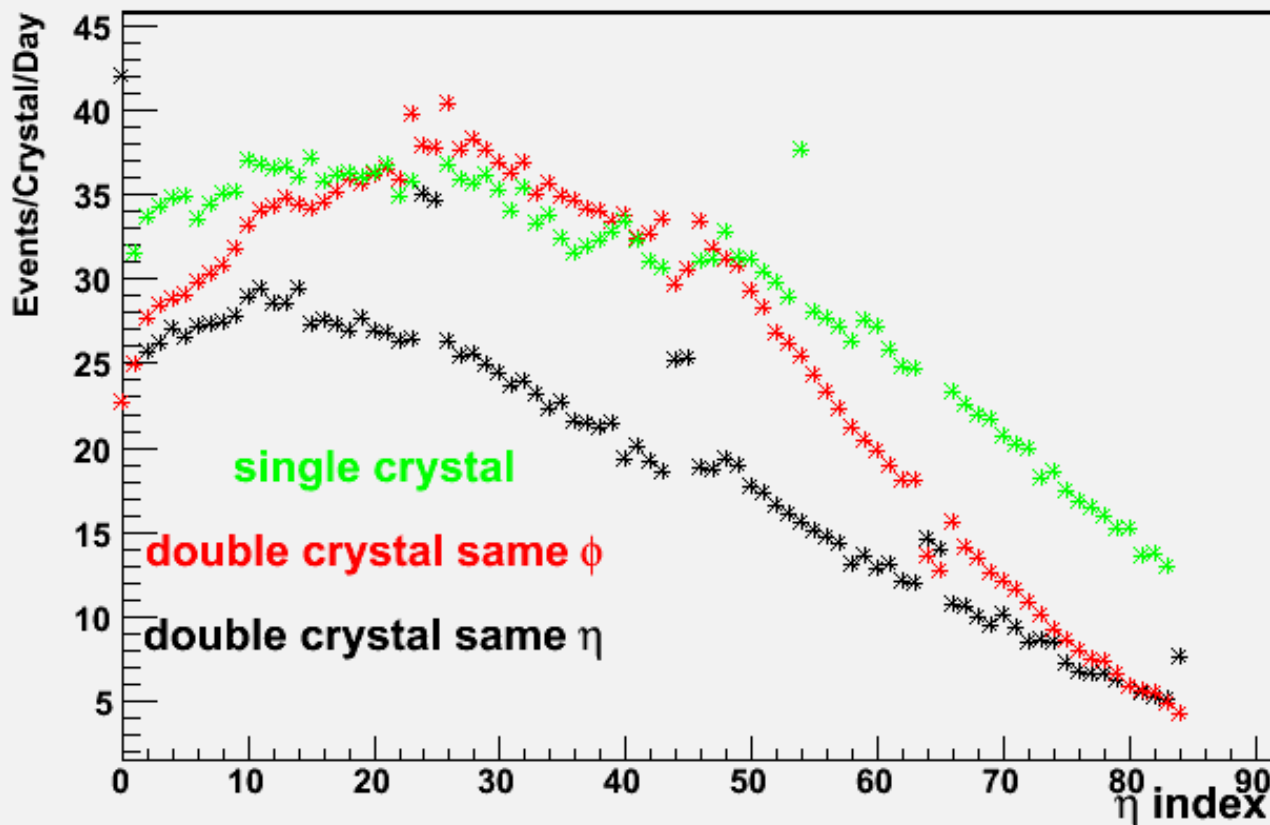
Useful Events



UNIVERSITY
OF MINNESOTA

Of the triggers
Single ~10% Same eta ~6.8% Same phi ~8.6%

Event Occupancy SM16



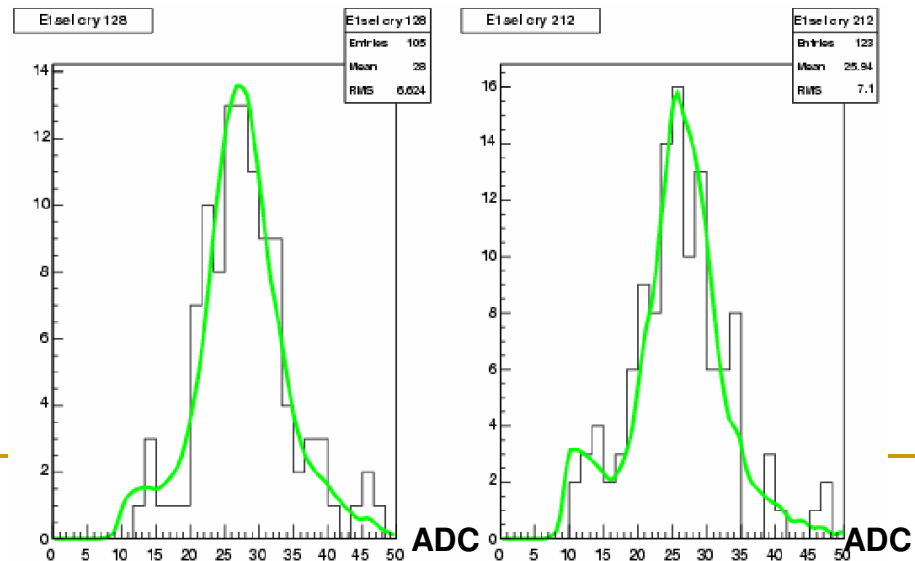
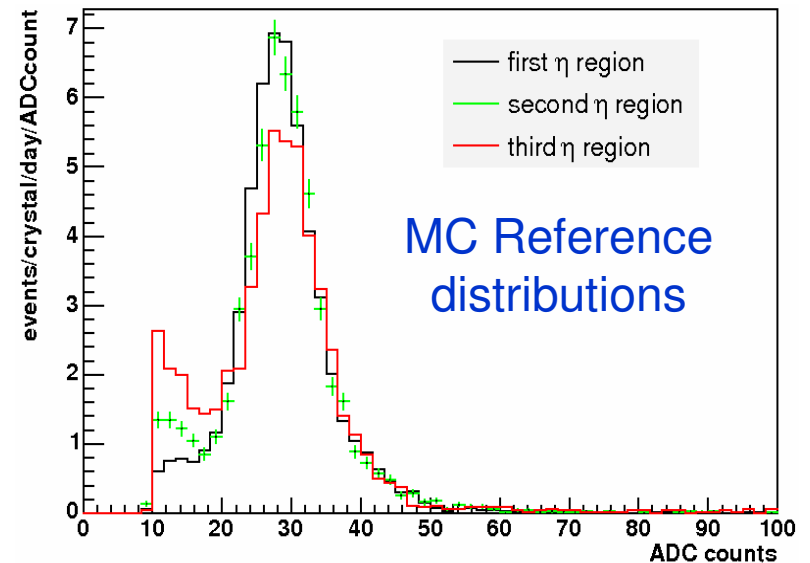
~500,000 Triggers/Day
~10 Days

~25% of the
triggers combined!

Single Crystal Method

- Pulse-height distributions of the crystal of highest energy deposited (E1) are made for each crystal
- Pulse-height distributions are also generated from MC information.
 - 17 different η dependent MC energy distributions are created.
- The constant is then found by adjusting the pulse-height to MC with an unbinned maximum likelihood method.
 - Factor: Relative Scale = Calibration constant

$$\mathcal{L} = \prod_i pdf(c \cdot E_i)$$





Two Crystal Method (Matrix Inversion)



UNIVERSITY
OF MINNESOTA

- Energy = Sum of both crystals
- Define a χ^2
- Minimize the χ^2 for each constant
- Fill Matrix event by event
- Invert
- **Only Input is the mean of the MC distributions**
- 17 Monte Carlo (MC) E1+E2 reference distributions are made
 - The mean is extracted from each within the selection range
- Reference distributions are created for special cases.
 - Module Borders
 - ϕ edges
 - η edges

$$E_{Total} = \sum c_i E_i$$

$$\chi^2 = \sum_{Events} \frac{(E_{True} - \sum c_i E_i)^2}{\sigma_E^2}$$

$$R_j = \sum_{Events} \frac{E_{True} E_j}{\sigma_E^2} \quad A_{ij} = \sum_{Events} \frac{E_i E_j}{\sigma_E^2}$$

$$R_j = \sum_i c_i A_{ij}$$

$$c = R \times A^{-1}$$



Calibration Schema

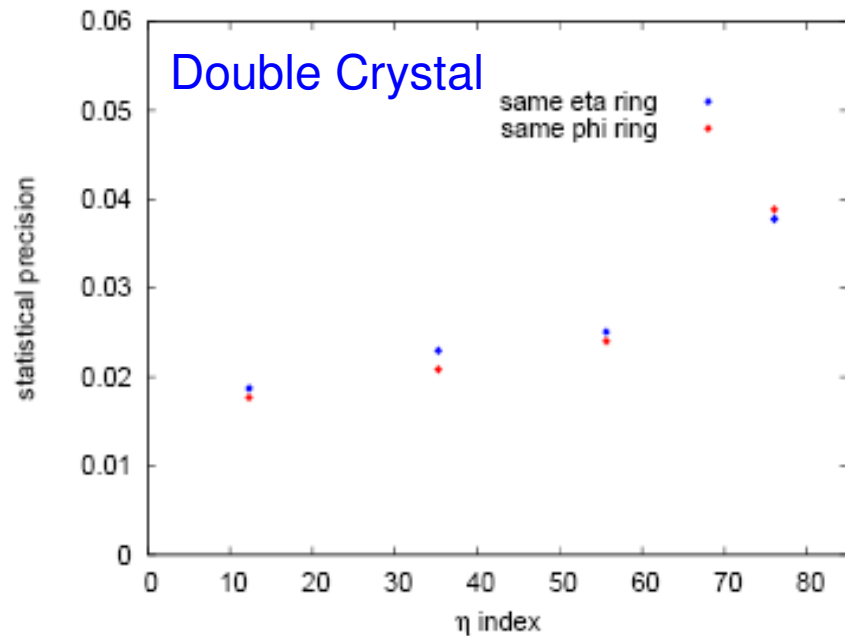
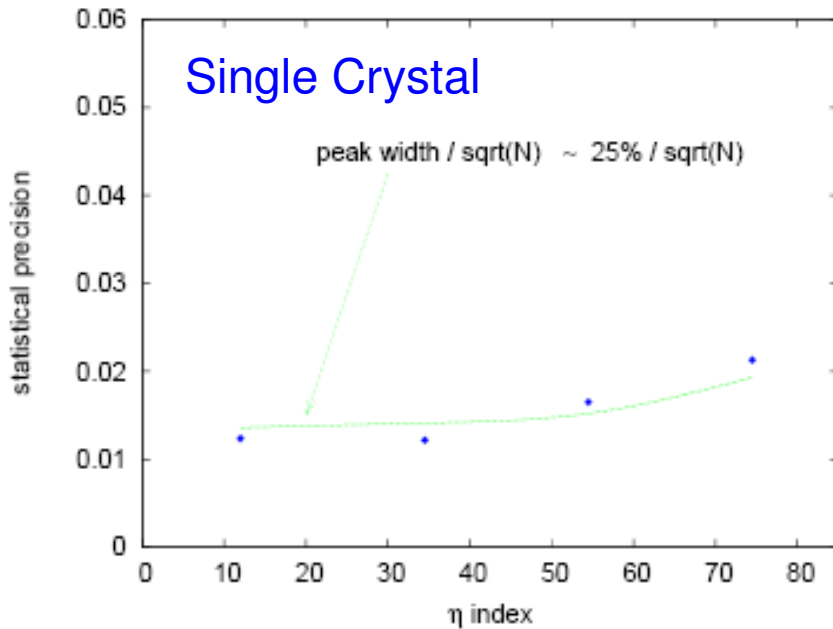
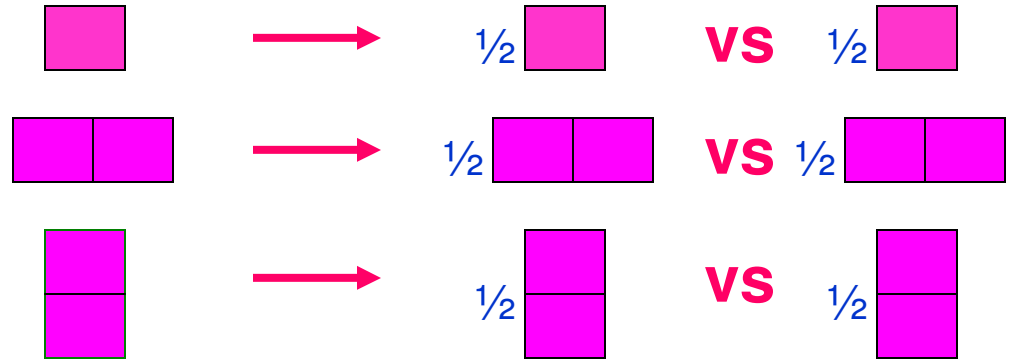


UNIVERSITY
OF MINNESOTA

- 36 SM's
- 10 will have electron testbeam calibration
- Use the beam inter-calibration constants to build reference distributions for the cosmic data
- Use these cosmic reference distributions in place of MC
- Validate method by comparing inter-calibration constants obtained with test beam and with cosmics for different SM's.
 - **Difference** → **precision**
- Then use cosmic ray data to obtain inter-calibration constants for remainder of SM's.

Statistical Precision

- Each of the 3 data sets are divided in half
- Odd vs. Even events
- $\sigma_{\text{stat}} = \sigma_{\text{odd vs. even}} / 2$





Test Beam data



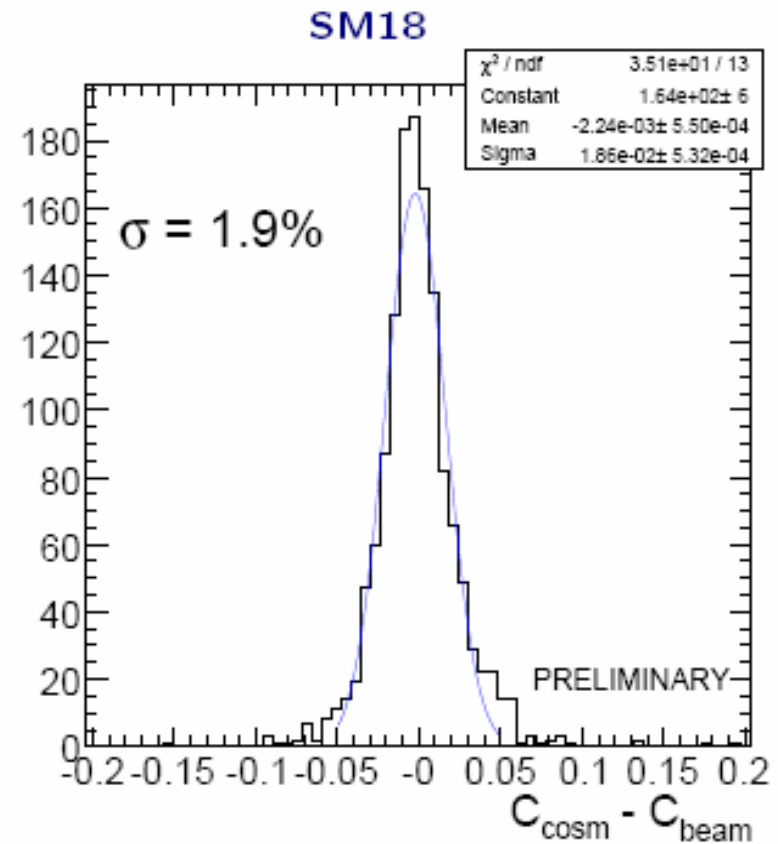
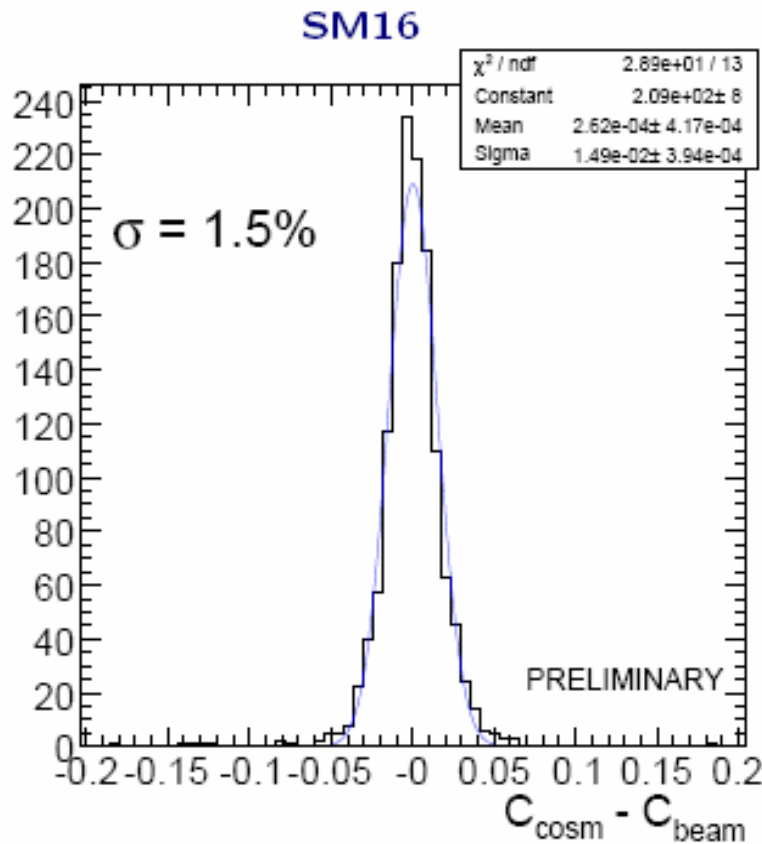
UNIVERSITY
OF MINNESOTA

- ECAL setup employs **full** CMS geometry and electronics setup
- Using electron beam with energy of 120 GeV
- Preliminaries available for two supermodules
- Used SM16 to build references
- Validate them with SM18
- **Combine** the datasets according to the statistical uncertainties

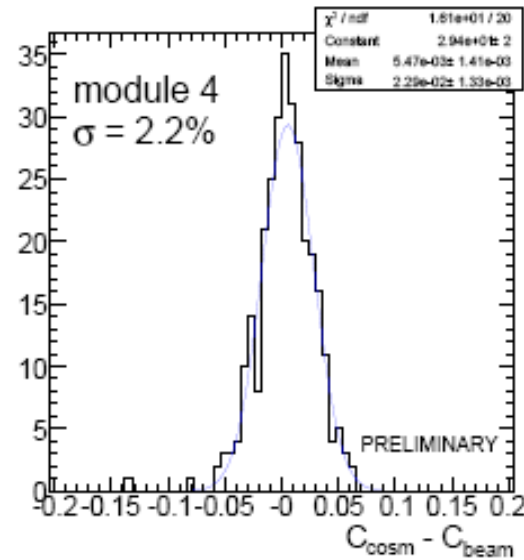
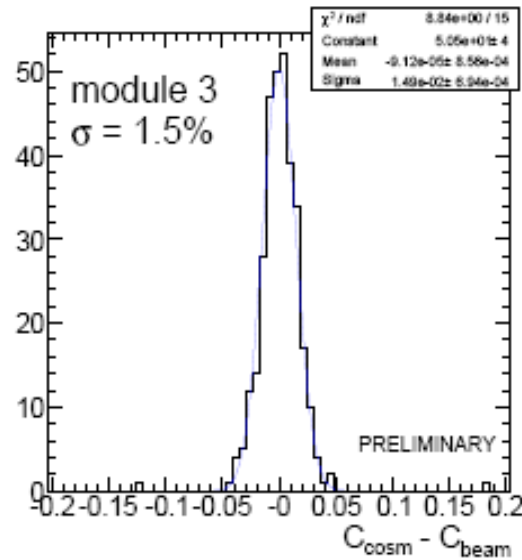
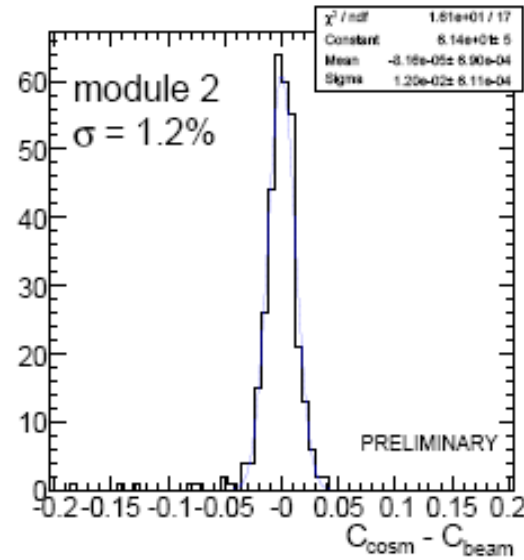
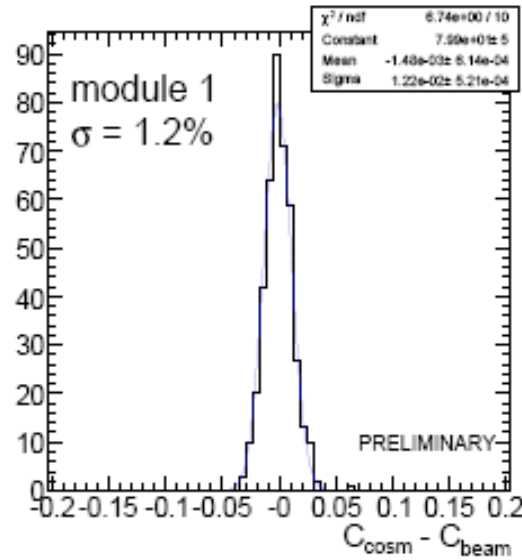
$$C_{comb}^i = \frac{\frac{C_{Single}^i}{\sigma_{Single}^2(\eta)} + \frac{C_{i\eta}^i}{\sigma_{i\eta}^2(\eta)} + \frac{C_{i\phi}^i}{\sigma_{i\phi}^2(\eta)}}{\frac{1}{\sigma_{Single}^2(\eta)} + \frac{1}{\sigma_{i\eta}^2(\eta)} + \frac{1}{\sigma_{i\phi}^2(\eta)}}$$

Test-Beam Comparison

- Good to 1.5% in SM16 and 1.9% in SM18
- Ignoring edges and module borders



SM16 TB Comparison by module





Future



UNIVERSITY
OF MINNESOTA

- Combined results with all beam calibrated SM's
 - Evaluate systematic uncertainties.
- Including beam data to check for differences between **muons** and **electrons** signals
- More than **23 SM's** have collected cosmic data. All are now expected to collect cosmic data.
- SM installation starting soon



Conclusion



UNIVERSITY
OF MINNESOTA

- 5 million triggers (10 days), provides a statistically accuracy of 2% or better
- Testbeam comparisons show an average agreement over an entire SM of less than 2%
- Cosmic ray muons will provide the most accurate inter-calibrations that will be available to all barrel crystals at CMS startup
- Using the ~250 MeV cosmic signal we can predict to 2% the calibration constants found at 120 GeV.
 - Uniform energy response between all crystals



References



UNIVERSITY
OF MINNESOTA

■ LY

- L. M. Barone *et al.* “Correlation between Light Yield and Longitudinal Transmission in PbWO₄ crystals and impact on the precision of the crystal intercalibration.” CMS RN 2004-005
http://cmsdoc.cern.ch/documents/04/rn04_005.pdf
- F. Cavallari *et al.* “Improvement on PbWO₄ Crystal Intercalibration Precision from Light Yield Measurements at the INFN-ENEA Regional Center” CMS RN 2004-003
http://cmsdoc.cern.ch/documents/04/rn04_003.pdf
- F. Cavallari *et al.* “Relative Light Yield comparison between laboratory and testbeam data for CMS ECAL PbWO₄ crystals” CMS RN 2004-002 http://cmsdoc.cern.ch/documents/04/rn04_002.pdf
- E. Auffray *et al.* “SPECIFICATIONS FOR LEAD TUNGSTATE CRYSTALS PREPRODUCTION” CMS Note 1998-038 http://cmsdoc.cern.ch/documents/04/rn98_038.pdf

■ Cosmic

- M. Bonesini *et al.* “Inter-calibration of the CMS electromagnetic calorimeter with cosmic rays before installation” CMS NOTE 2005/023
- W. Bertl *et al.* “Feasibility of Intercalibration of CMS ECAL Supermodules with Cosmic Rays” CMS NOTE 2004/036
- K. Deiters *et al.* “Test of the Feasibility of Pre-intercalibration of ECAL Supermodules with Cosmic Rays” CMS IN 2004/023

■ General

- CMS Collaboration, *CMS Technical Proposal*, CERN/LHCC 94-38
- CMS Collaboration, *The Electromagnetic Calorimeter Project TDR*, CERN/LHCC 97-33
- Particle Data Handbook and references therein



UNIVERSITY
OF MINNESOTA

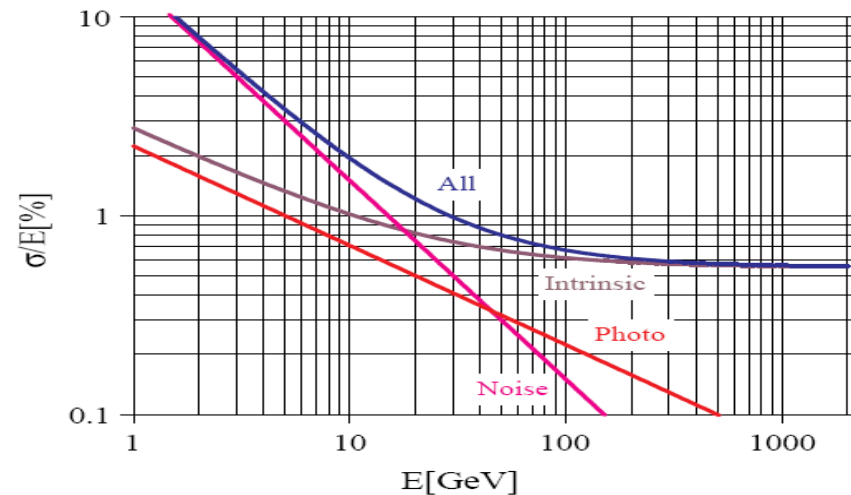
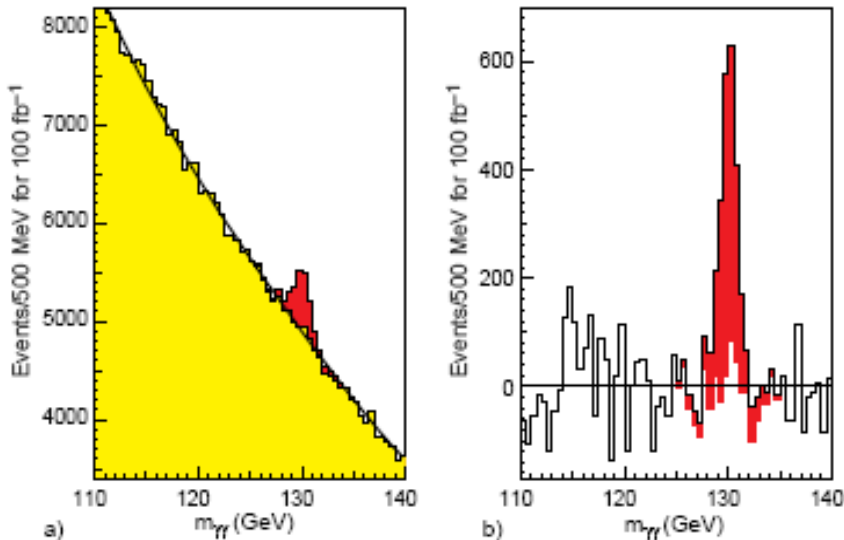
Extras

ECAL: Higgs $\rightarrow \gamma\gamma$

- The reconstructed mass the the Higgs depends on the Energy of both photons as well as the angle between the two.
- The error of the photon energy is very important

$$m_{\gamma\gamma}^2 = 2E_{\gamma 1}E_{\gamma 2}(1 - \cos\theta_{\gamma 1, \gamma 2})$$

$$\frac{\sigma_{m_{\gamma\gamma}}}{m_{\gamma\gamma}} = \frac{1}{2} \left[\frac{\sigma_{E_{\gamma 1}}}{E_{\gamma 1}} \oplus \frac{\sigma_{E_{\gamma 2}}}{E_{\gamma 2}} \oplus \frac{\sigma_{\theta_{\gamma\gamma}}}{\tan(\theta_{\gamma\gamma}/2)} \right]$$



Term	% expected	σ_E for 100GeV
a	$2.7\%/\sqrt{E}$	270 MeV
c	155 MeV in E_T	265 + 355 MeV
b	0.5%	500 MeV

Selection from MC

