

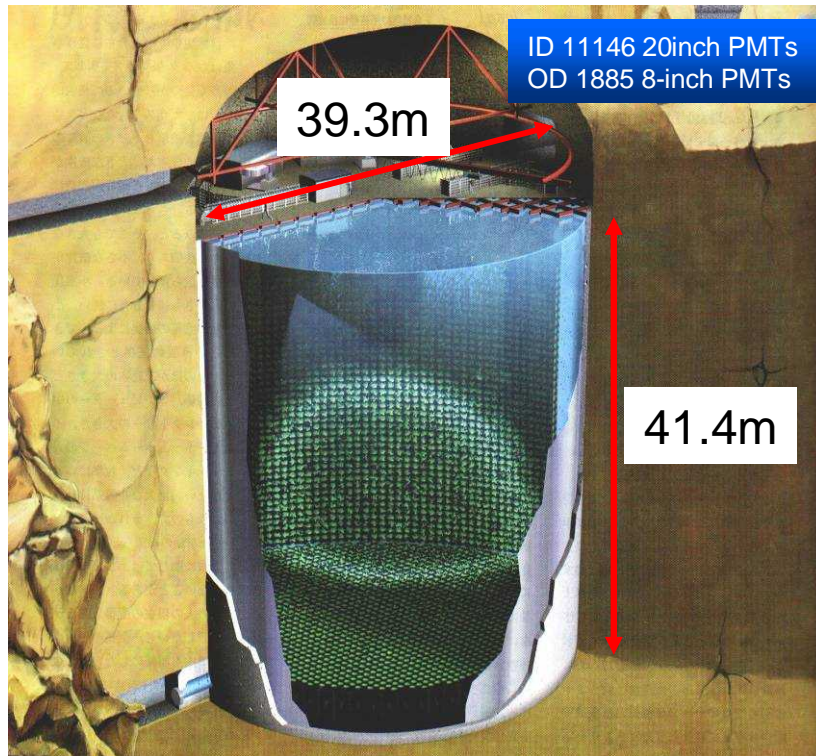
Oscillation analysis results using Super-Kamiokande atmospheric neutrino data

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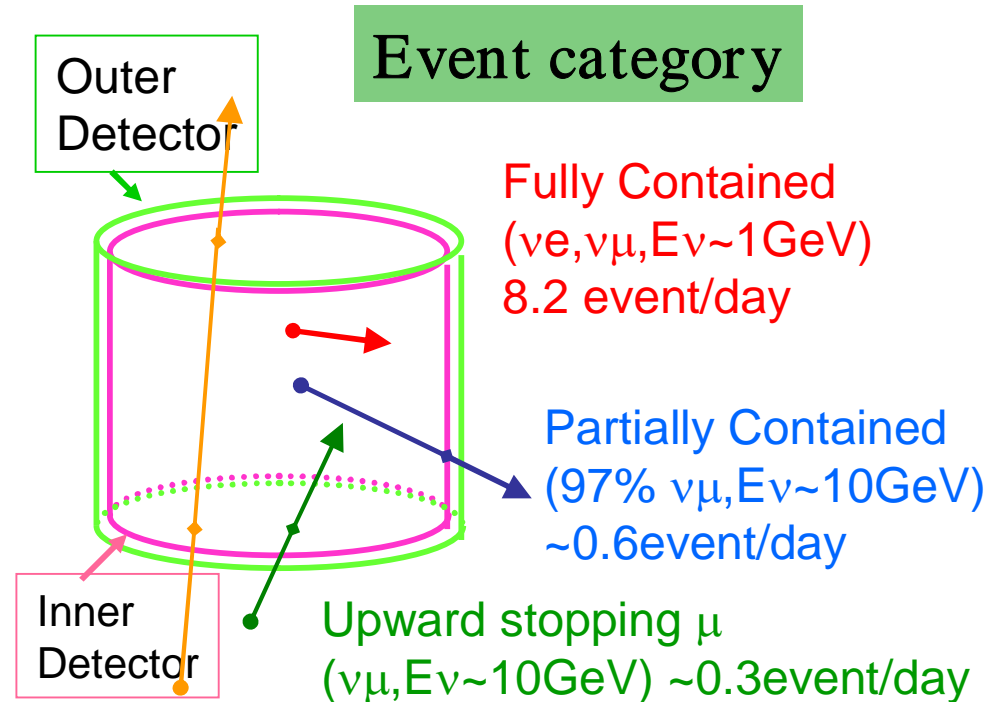
Joint Meeting of Pacific Region Particle Physics Communities

Oct. 31, 2006

Super-Kamiokande

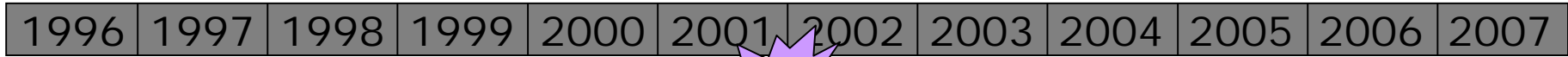


- Water Cherenkov detector
- 50kton (22.5 kton fid.), cylindrical shape
- Deep underground (2700 w.m.e)

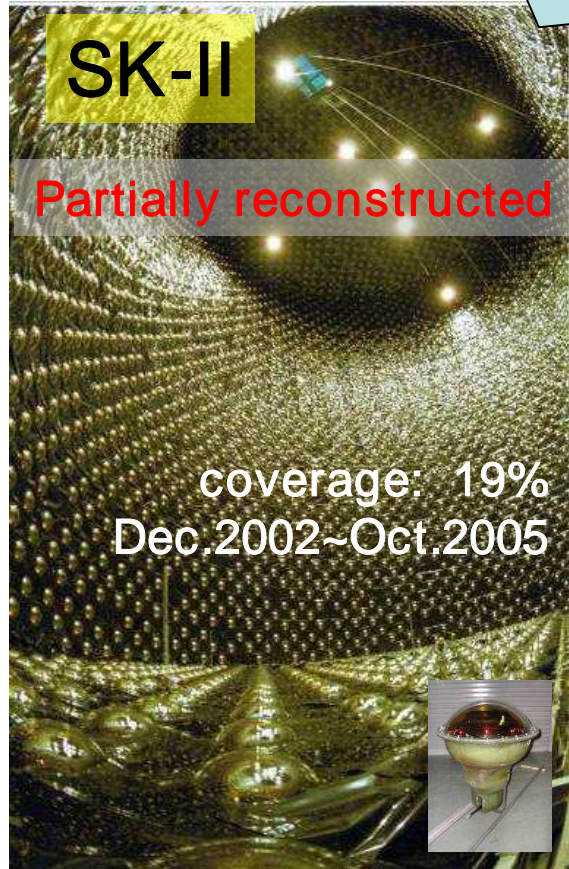
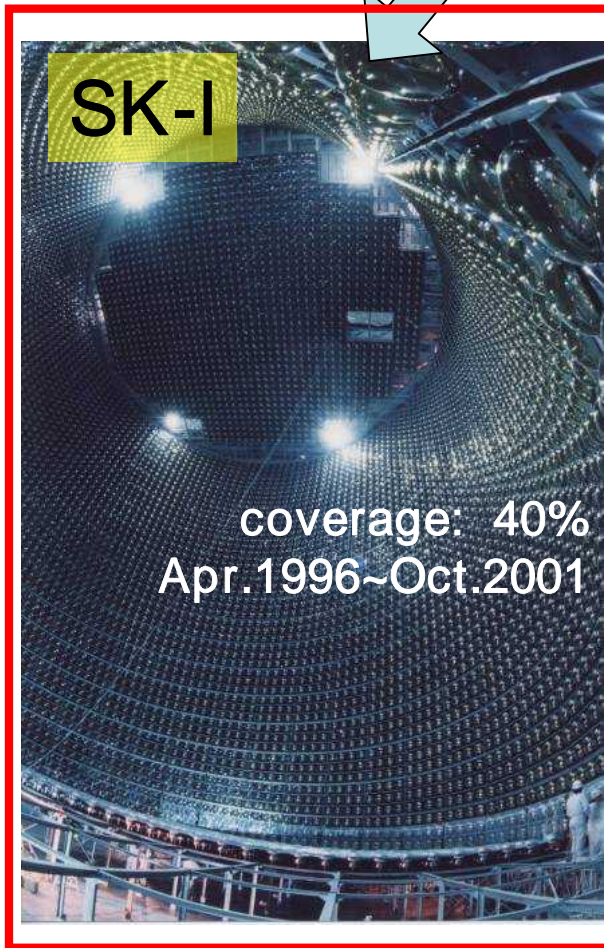


Upward through-going μ ($\nu_\mu, E_\nu \sim 100\text{GeV}$) $\sim 1.1\text{event/day}$

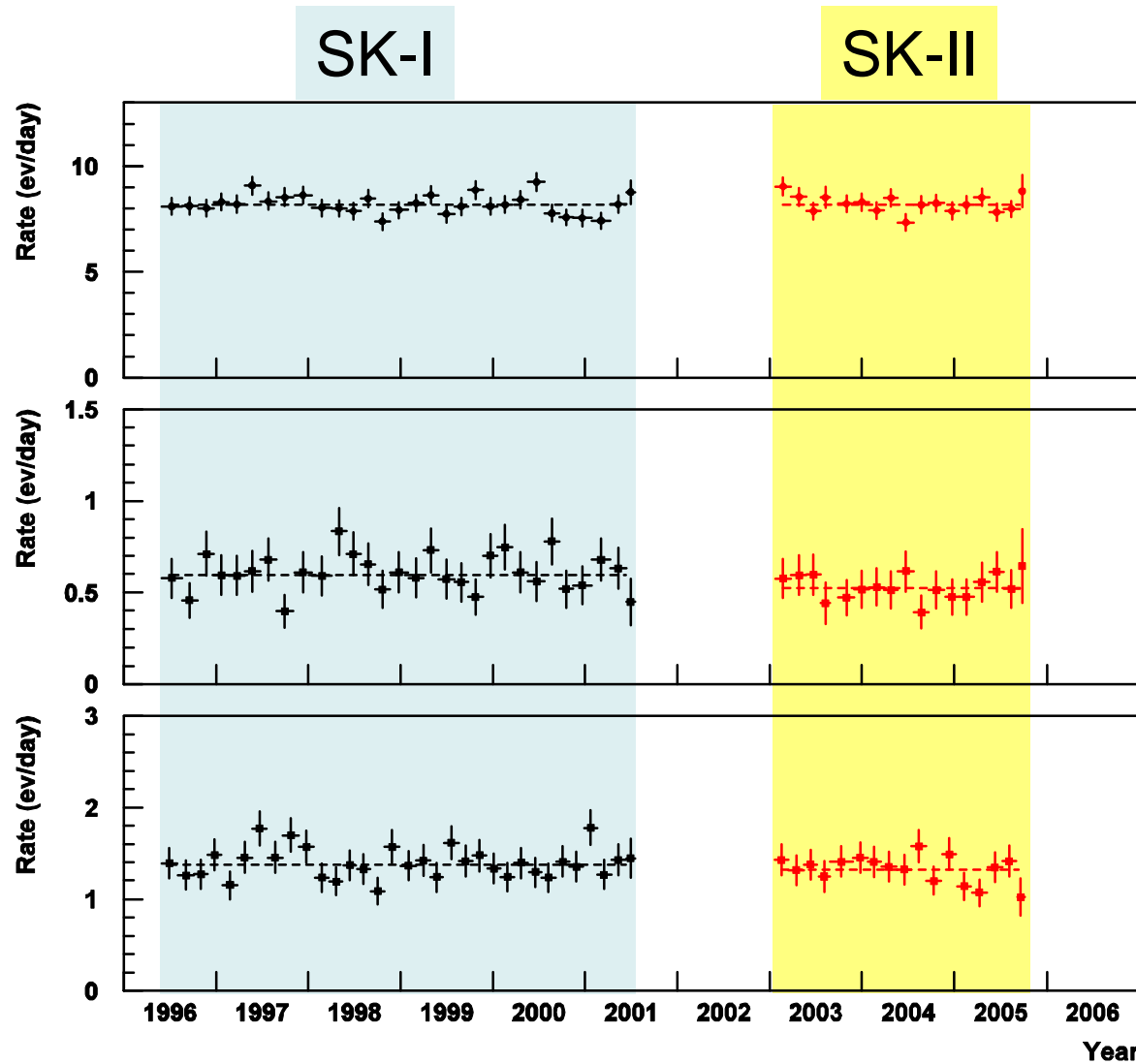
History of Super-Kamiokande



accident



Event rate



FC

SK-II: 8.22 ± 0.10 ev/d
SK-I : 8.19 ± 0.07 ev/d

PC

SK-II: 0.53 ± 0.03 ev/d
SK-I : 0.60 ± 0.02 ev/d

UP μ (stop+thru)

SK-II: 1.42 ± 0.04 ev/d
SK-I : 1.45 ± 0.03 ev/d

Preliminary

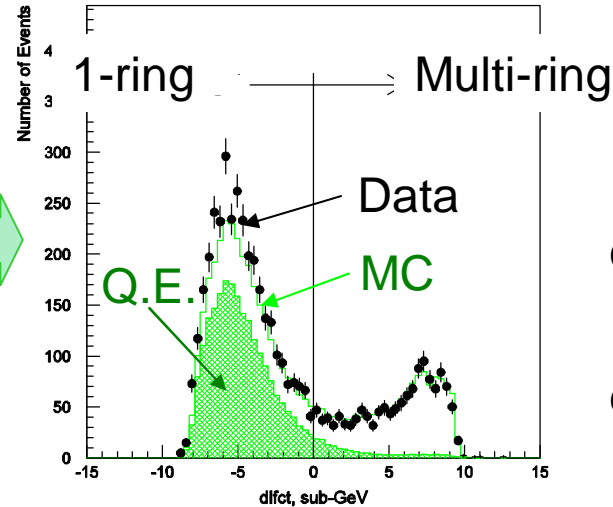
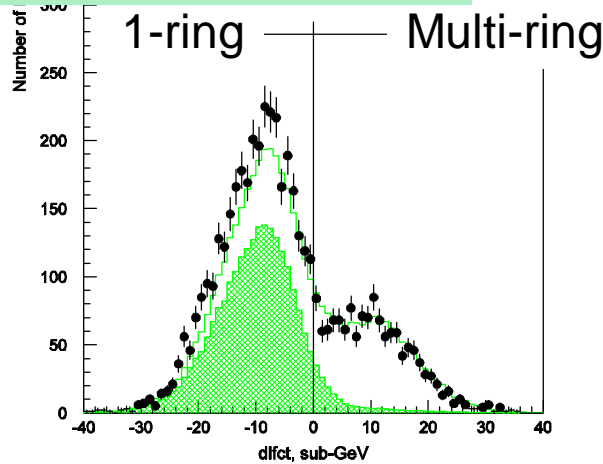
SK-II performance (Ring counting)

preliminary

Old ring counting

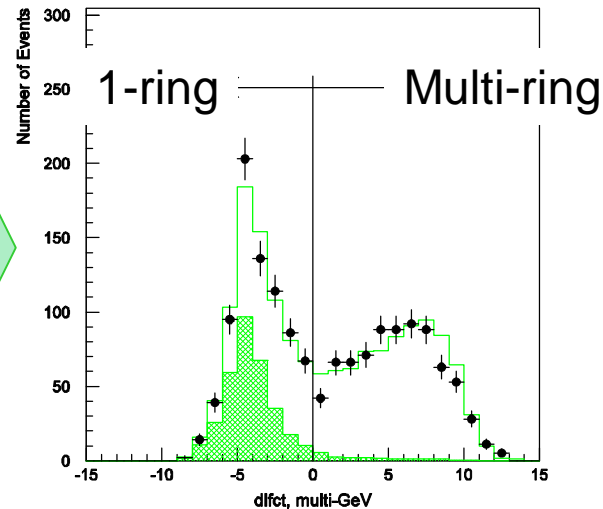
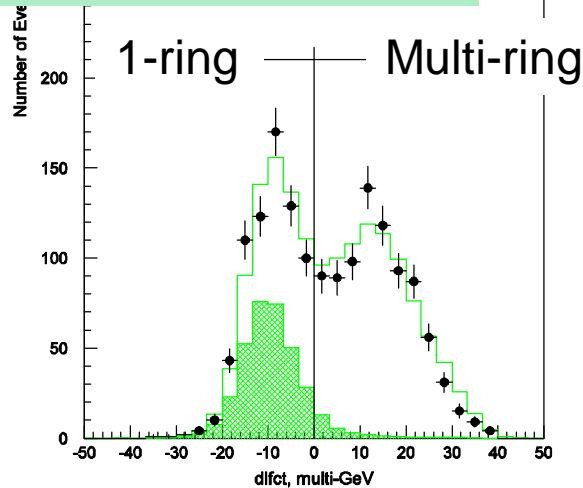
New ring counting

Sub-GeV(<1.33GeV)



- improved ring-counting algorithm
- better separation, especially in multi-GeV
- reduced systematic error

Multi-GeV(>1.33GeV)



SK-II data summary

Live time : 804 days (FC,PC)
828 days (UP μ)

Honda 2003, w/o oscillation

Preliminary

	FC sub-GeV		FC multi-GeV	
	Data	MC	Data	MC
1R	3565	3769.9	796	937.7
e-like	1842	1554.5	417	426.2
μ -like	1723	2215.4	379	511.5
2R	938	1033.8	318	377.1
$\geq 3R$	364	477.0	430	526.8
Total	4867	5280.8	1544	1841.5

	PC	
	Data	MC
Total	427	592.3

	UP μ	
	Data	MC
stop	228	350.8
thru	889	832.0

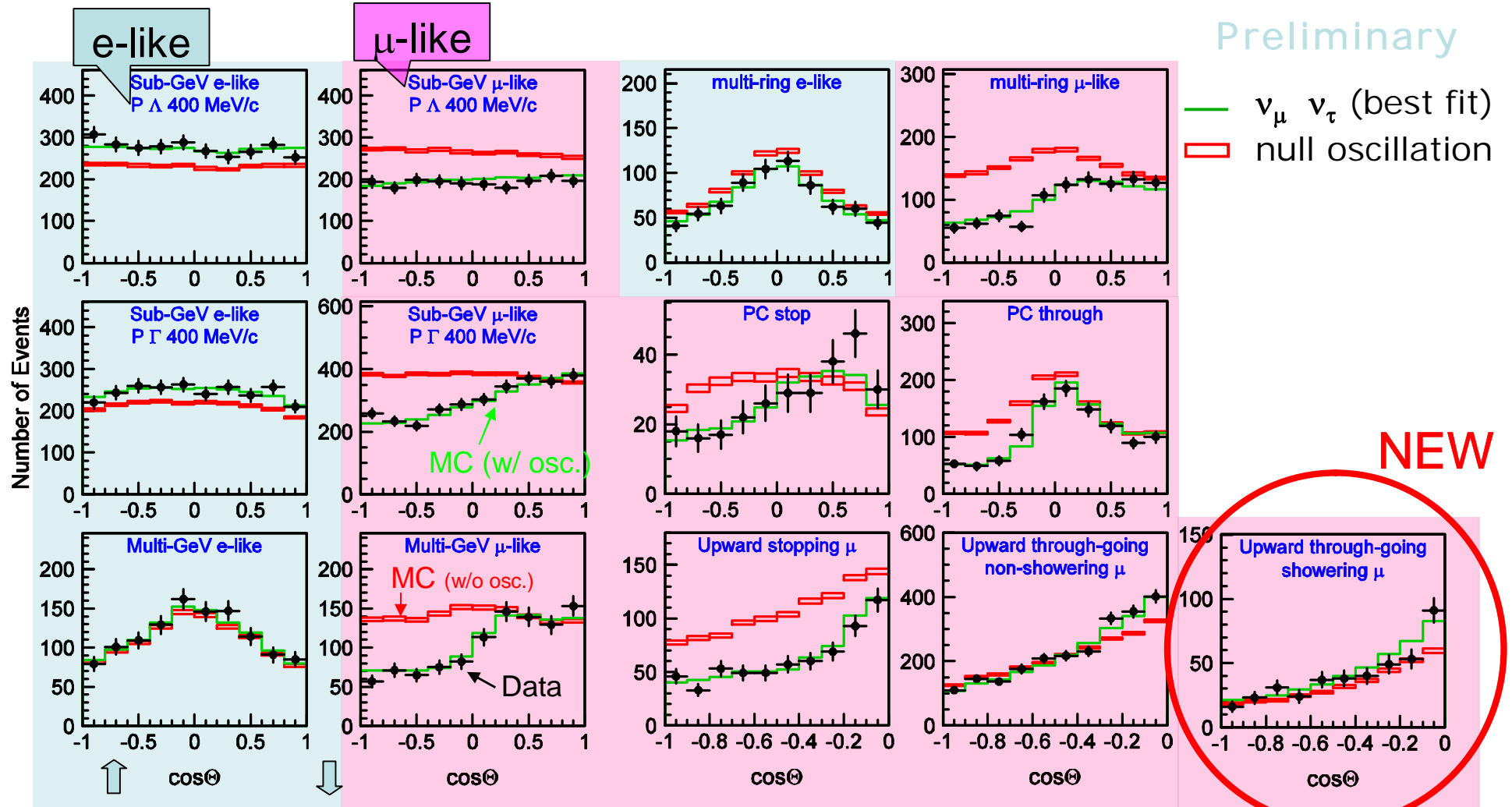
SK-II double ratio

$$\frac{(\mu/e)_{\text{DATA}}}{(\mu/e)_{\text{MC}}} = \begin{cases} 0.656 \pm 0.022 \pm 0.033 & \text{(FC Sub-GeV)} \\ 0.746_{-0.044}^{+0.047} \pm 0.055 & \text{(FC Multi-GeV + PC)} \end{cases}$$

SK-I

$$\begin{cases} 0.658 \pm 0.016 \pm 0.035 \\ 0.702_{-0.030}^{+0.032} \pm 0.099 \end{cases}$$

Zenith Angle Distributions (SK-I + SK-II)

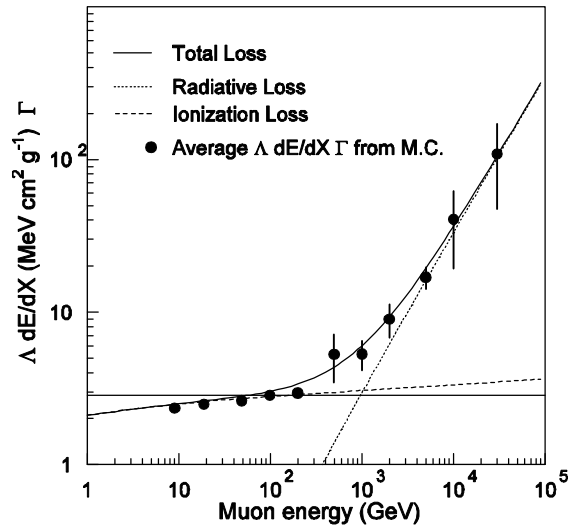


Excellent agreement with $\nu_\mu \rightarrow \nu_\tau$ oscillation hypothesis

upward-going showering muon

Upward showering muon

muon dE/dx in water

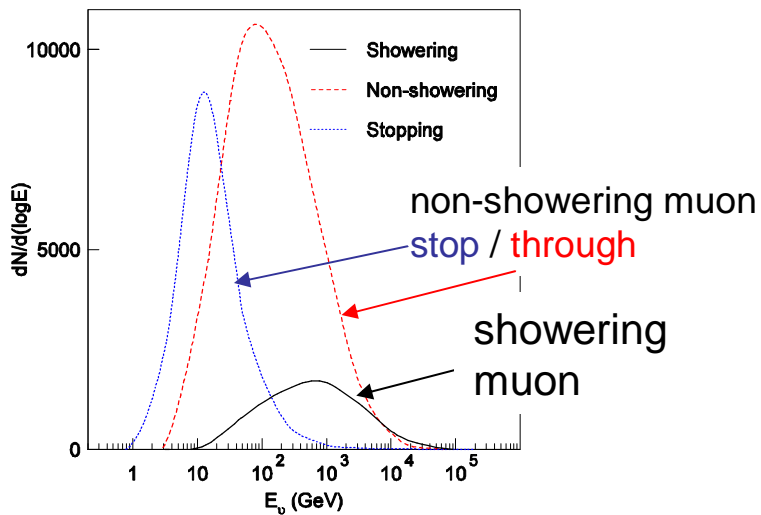


(Radiation loss) > (Ionization loss) @ $P_\mu > \sim 1 \text{ TeV}$

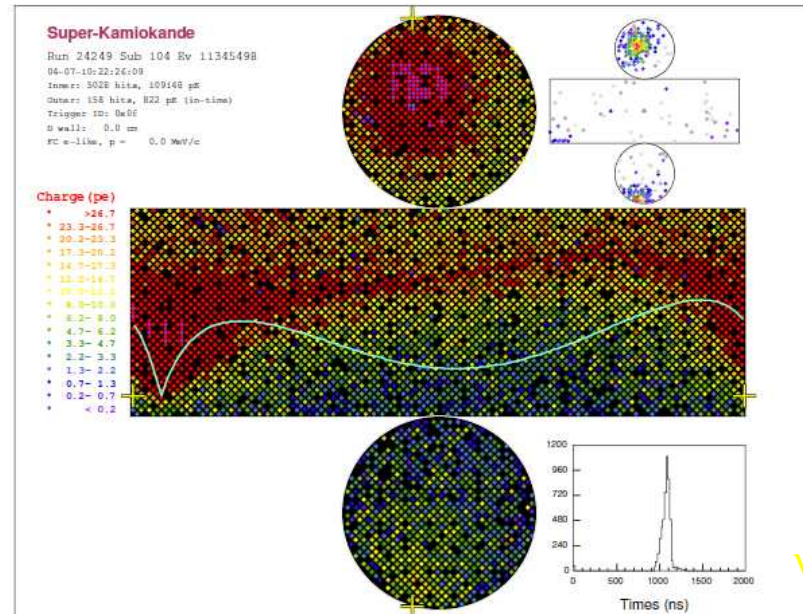
$\langle E_\nu \rangle \sim 1 \text{ TeV}$

Separation using charge profile

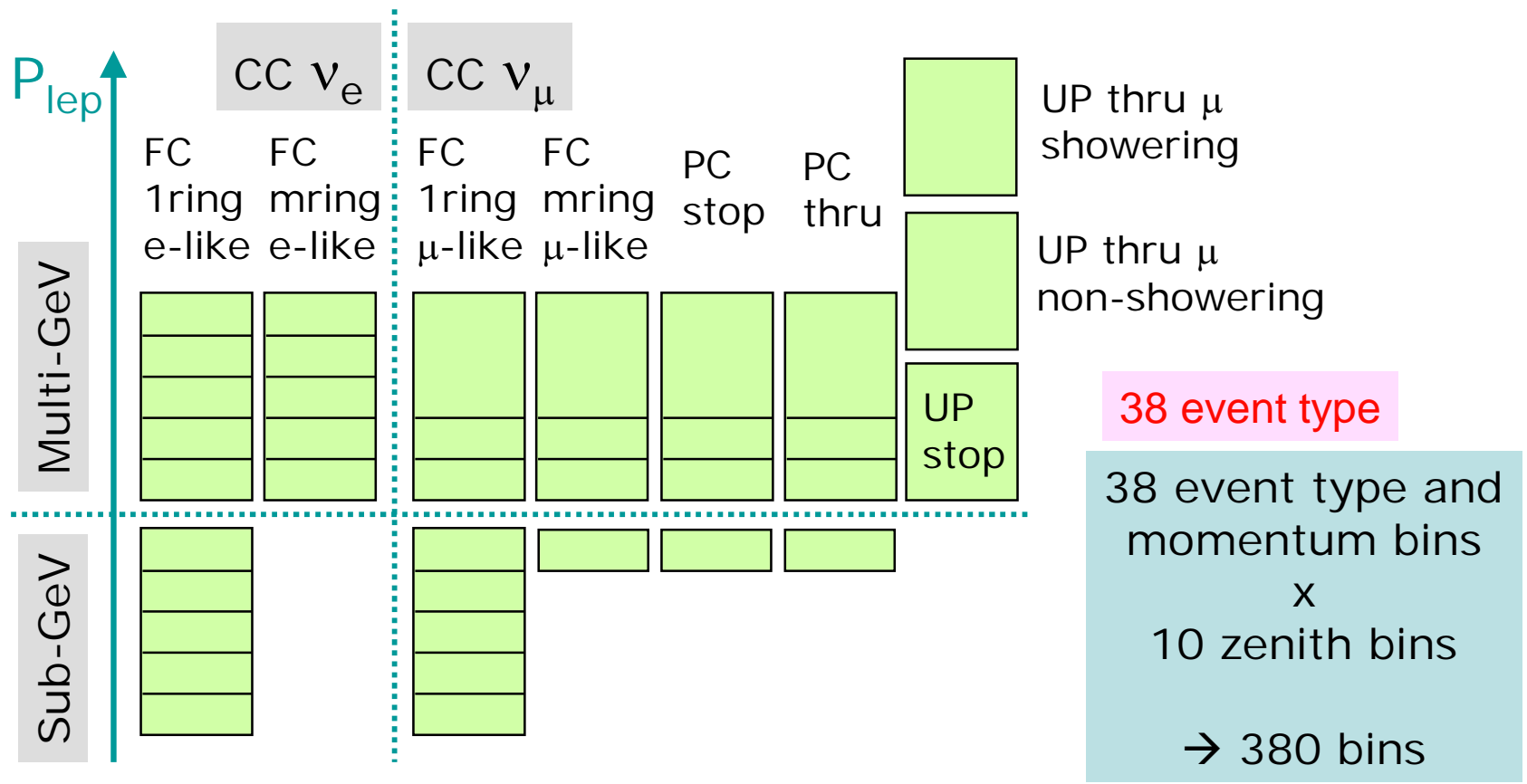
309 (91) events observed in SK-I (SK-II)



typical showering muon (SK-II)



SK-I + SK-II combined analysis (data binning)



380 bins for SK-I + 380 bins for SK-II → 760 bins in total

Statistical Treatment of Binned Data

- “ Pull Approach ” G.L.Fogli et al. PRD66 (2002) 053010
- χ^2 incorporates Poissonian uncertainties.

number of p, Θ ,
event-type bins

number of sys. effects
(normalization is free)

$$\chi^2 \equiv \sum_{i=1}^{760} \left(2(N_i^{\text{exp}} - N_i^{\text{obs}}) + 2N_i^{\text{obs}} \ln \frac{N_i^{\text{obs}}}{N_i^{\text{exp}}} \right) + \sum_{j=1}^{68} \left(\frac{\varepsilon_j}{\sigma_j} \right)^2$$

$$N_i^{\text{exp}} = N_i^0 \cdot P(\nu_\alpha \rightarrow \nu_\beta) \cdot \left(1 + \sum_{j=1}^{70} \underline{f_j^i} \cdot \varepsilon_j \right)$$

fractional change in
predicted event rate
due to variation in
systematic param. ε

At each point on a $(\sin^2 2\theta, \Delta m^2)$ grid, χ^2 is minimized by varying systematic parameters ε_j .

Systematic errors

used in previous analyses

hep-ex/0501064

neutrino flux (14)

neutrino interaction (12)

solar activity (1)

event selection
and reconstruction (19)

BG sub. from showering (1)

shower / non-shower sep.(1)

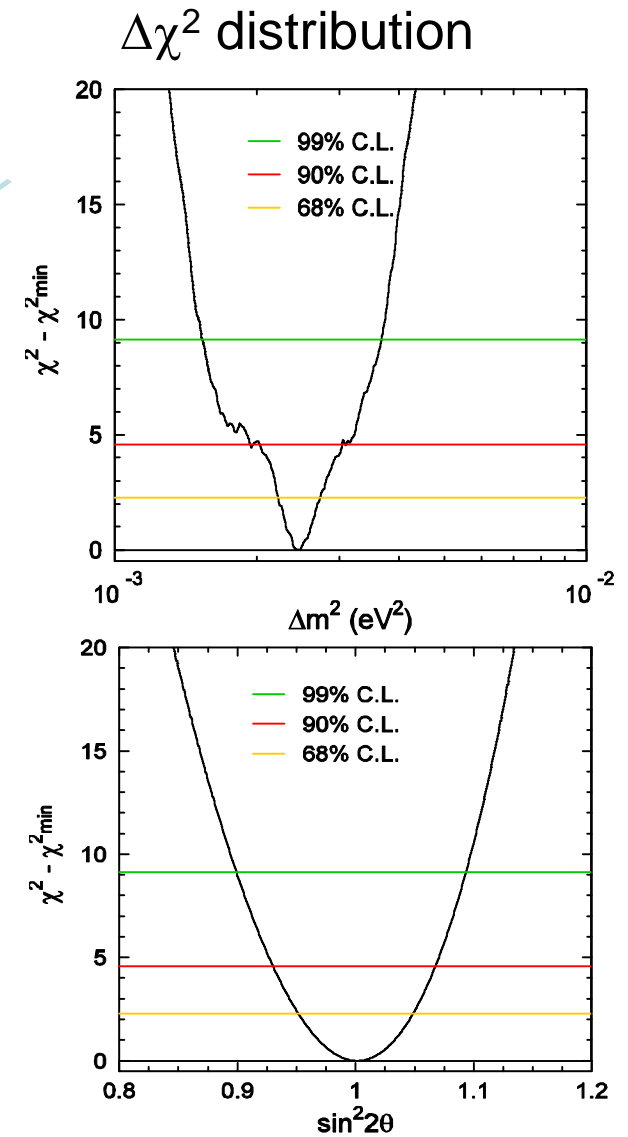
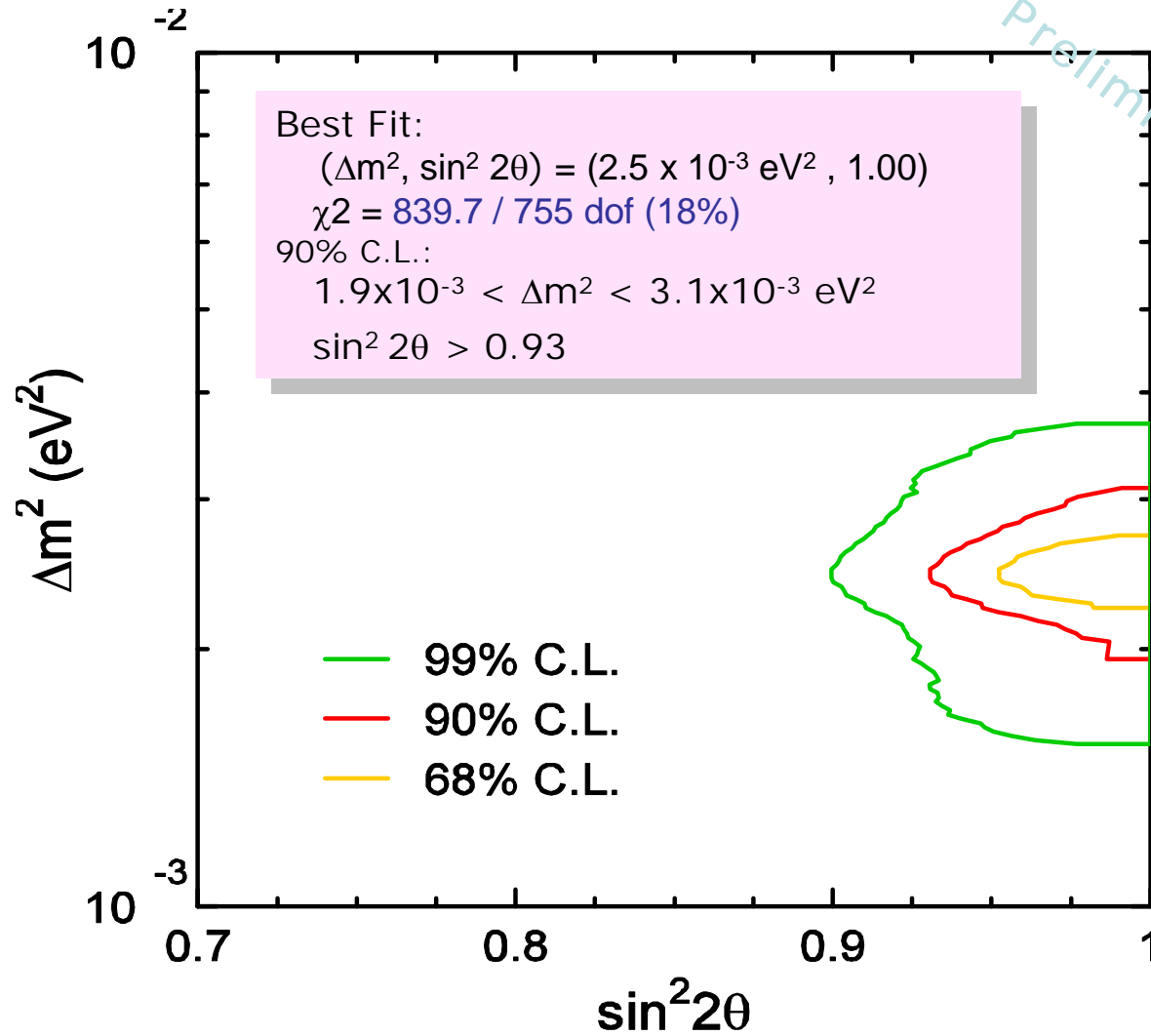
Identical for SK-I and SK-II

Regarded as independent
between SK-I and SK-II

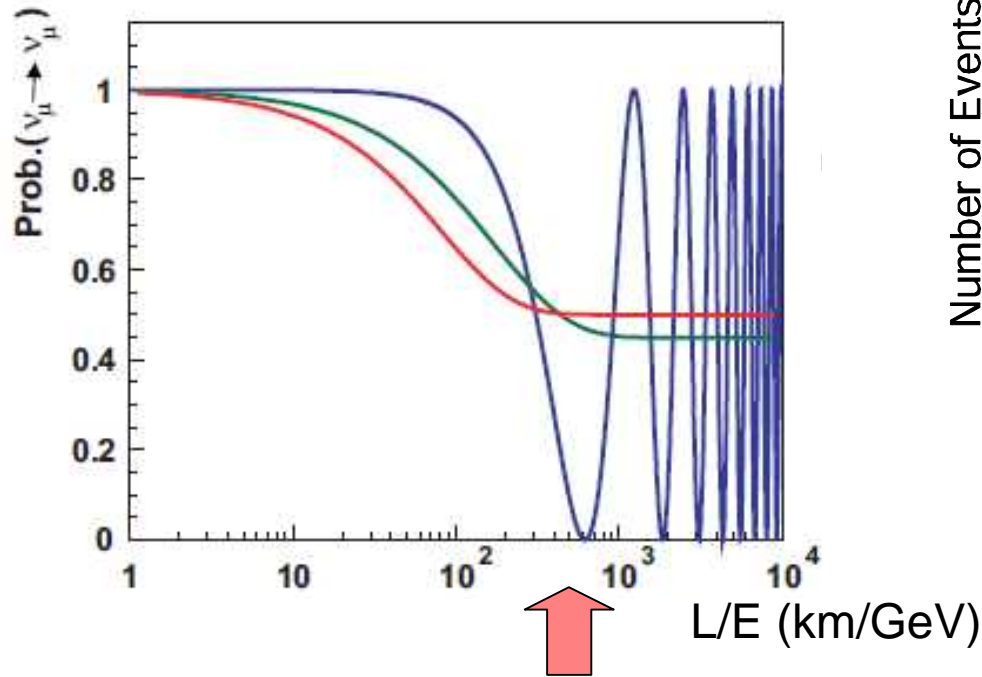
The total number of systematic errors is :

$$\text{Flux (14) + Interaction (12) + SK-I (22) + SK-II (22) = 70}$$

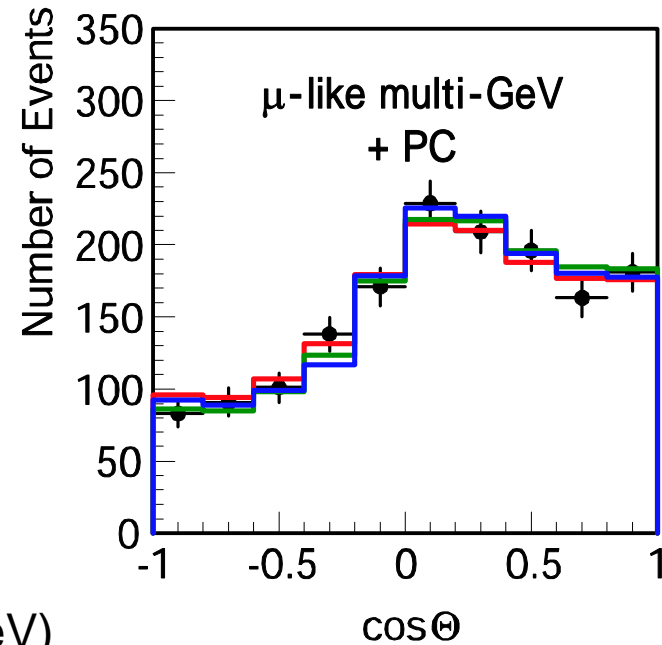
Allowed region



L/E analysis

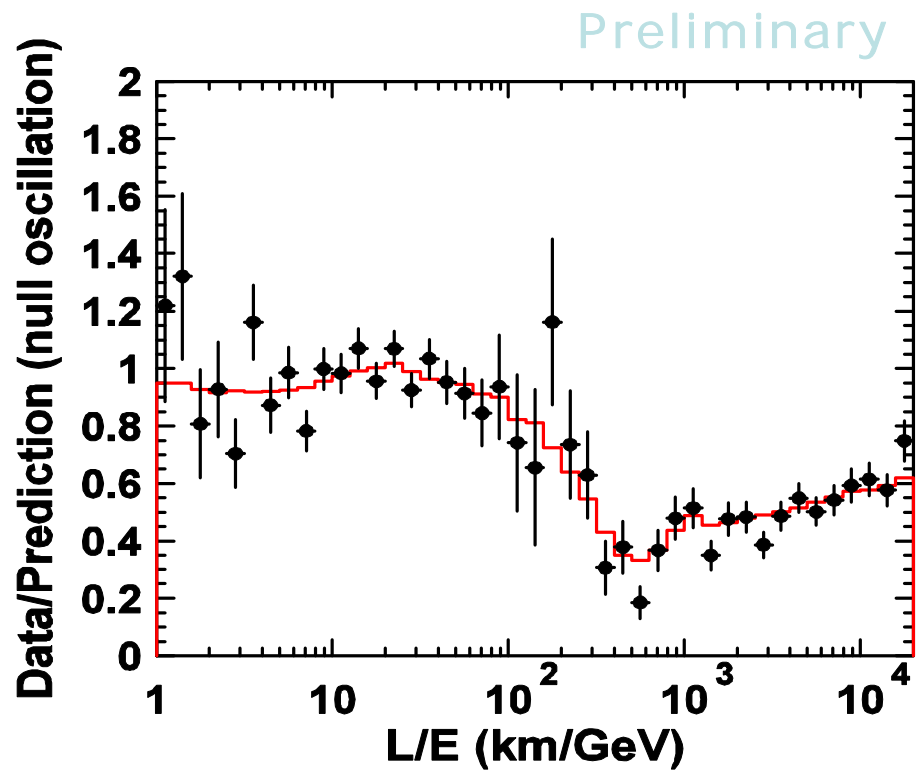
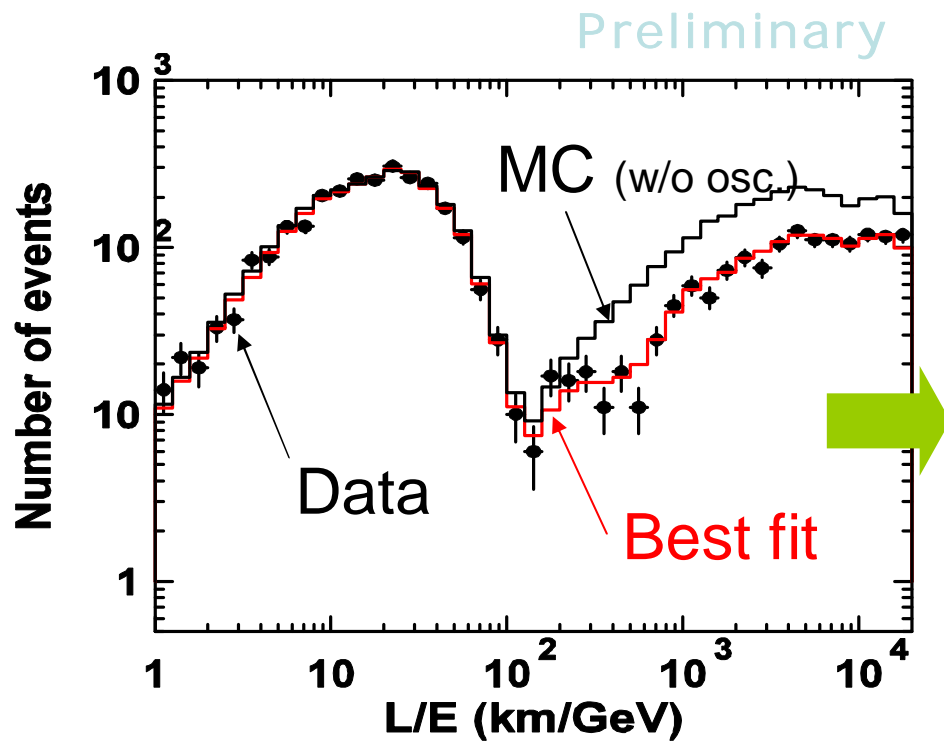


this dip should be observed



Neutrino oscillation :	$P_{\mu\mu} = 1 - \sin^2 2\theta \sin^2 \left(1.27 \frac{\Delta m^2 L}{E} \right)$
Neutrino decay :	$P_{\mu\mu} = \left(\cos^2 \theta + \sin^2 \theta \times \exp \left(-\frac{m}{2\tau} \frac{L}{E} \right) \right)^2$
Neutrino decoherence :	$P_{\mu\mu} = 1 - \frac{1}{2} \sin^2 2\theta \times \left(1 - \exp \left(-\gamma_0 \frac{L}{E} \right) \right)$

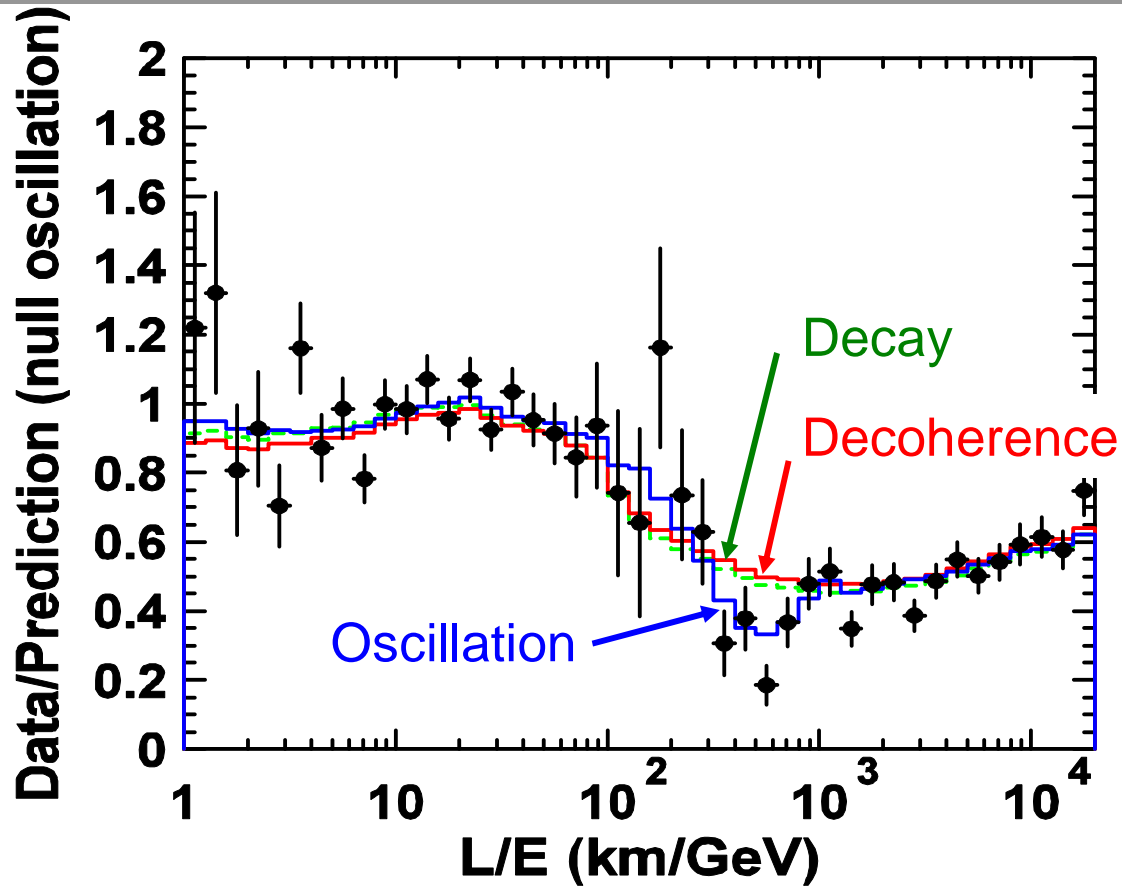
L/E distribution



Event sample:
SK1+SK2 FC μ -like, PC
w/ resolution cut

Best fit: $\Delta m^2=2.3 \times 10^{-3}, \sin^2 2\theta=1.00$
 $\chi^2_{\min}=83.9/83$ d.o.f
($\sin^2 2\theta=1.03, \chi^2_{\min}=83.4/83$ d.o.f)

L/E analysis result



$$\chi^2_{\text{osc}} = 83.9/83 \text{ d.o.f}$$

SK-I

$$\chi^2_{\text{dcy}} = 107.1/83 \text{ d.o.f}, \Delta\chi^2 = 23.2 (4.8 \text{ })$$

3.4 σ

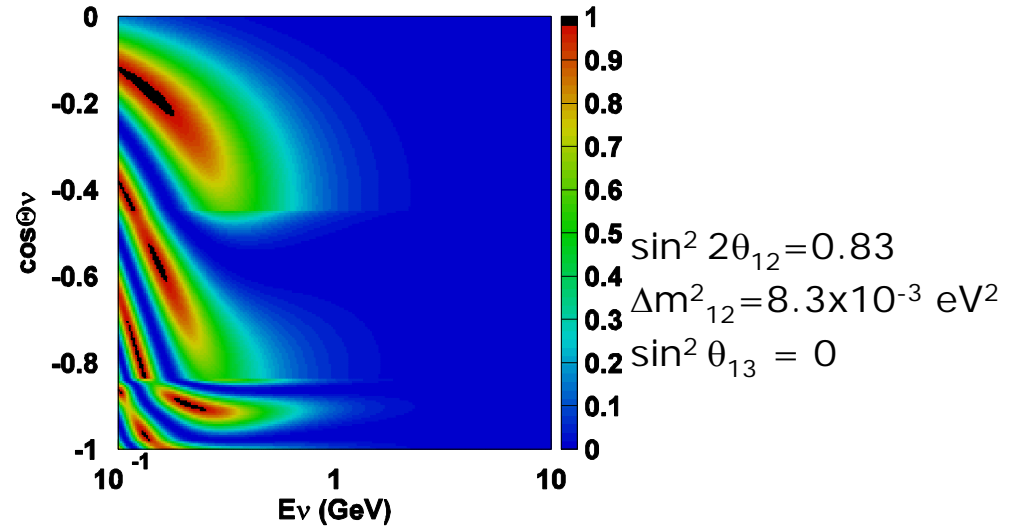
$$\chi^2_{\text{dec}} = 112.5/83 \text{ d.o.f}, \Delta\chi^2 = 27.6 (5.3 \text{ })$$

3.8 σ

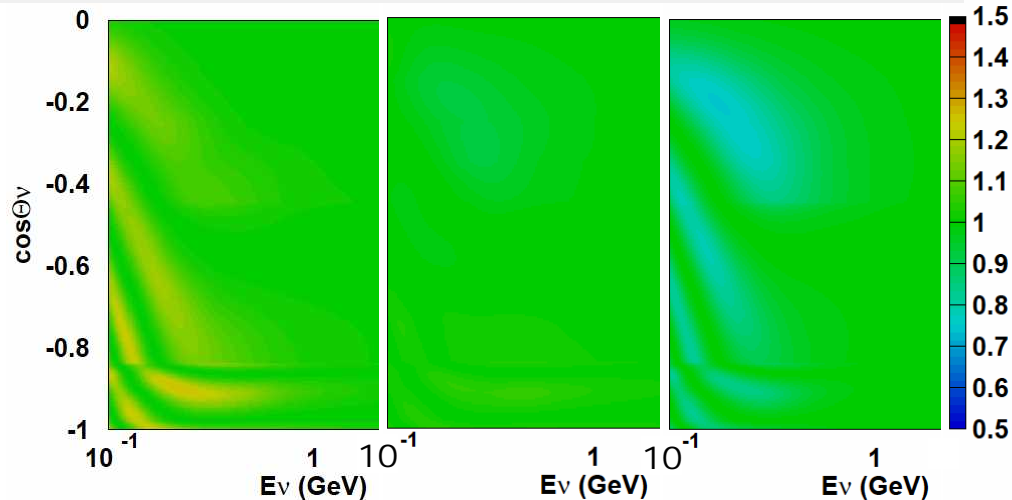
Solar term effect to atmospheric neutrinos

Due to the LMA solution, atmospheric neutrinos should also oscillate by $(\theta_{12}, \Delta m^2_{12})$.

P_2 : 2ν transition prob. $\nu_e \rightarrow \nu_{\mu, \tau}$ in matter driven by Δm^2_{12}



$\sin^2 2\theta_{23} = 0.96$	$= 1.0$	$= 0.96$
$\sin^2 \theta_{23} = 0.4$	$= 0.5$	$= 0.6$



$\frac{\nu_e \text{ flux (osc.)}}{\nu_e \text{ flux (no osc.)}}$

Discrimination between $\theta_{23} < \pi/4$ and $> \pi/4$ might be possible by studying low energy atmospheric ν_e and ν_μ events.

Oscillation analysis with solar terms

- Data set : SK-I (1489 days) + SK-II (804 days)

- Oscillation maps :

$$\Delta m^2_{12} = 10^{-4.16} \sim 10^{-4.04} \text{ eV}^2 \quad (5 \text{ points})$$

$$\sin^2 \theta_{12} = 0.22 \sim 0.40 \quad (7 \text{ points})$$

$$\sin^2 \theta_{13} = 0 \quad (\text{fixed})$$

$$\Delta m^2_{23} = 10^{-3} \sim 10^{-2.2} \text{ eV}^2 \quad (41 \text{ points})$$

$$\sin^2 \theta_{23} = 0.3 \sim 0.7 \quad (41 \text{ points})$$

→ 4 dimensional analysis

- Calculate the minimum χ^2 : projection to the $\sin^2 \theta_{23}$ axis

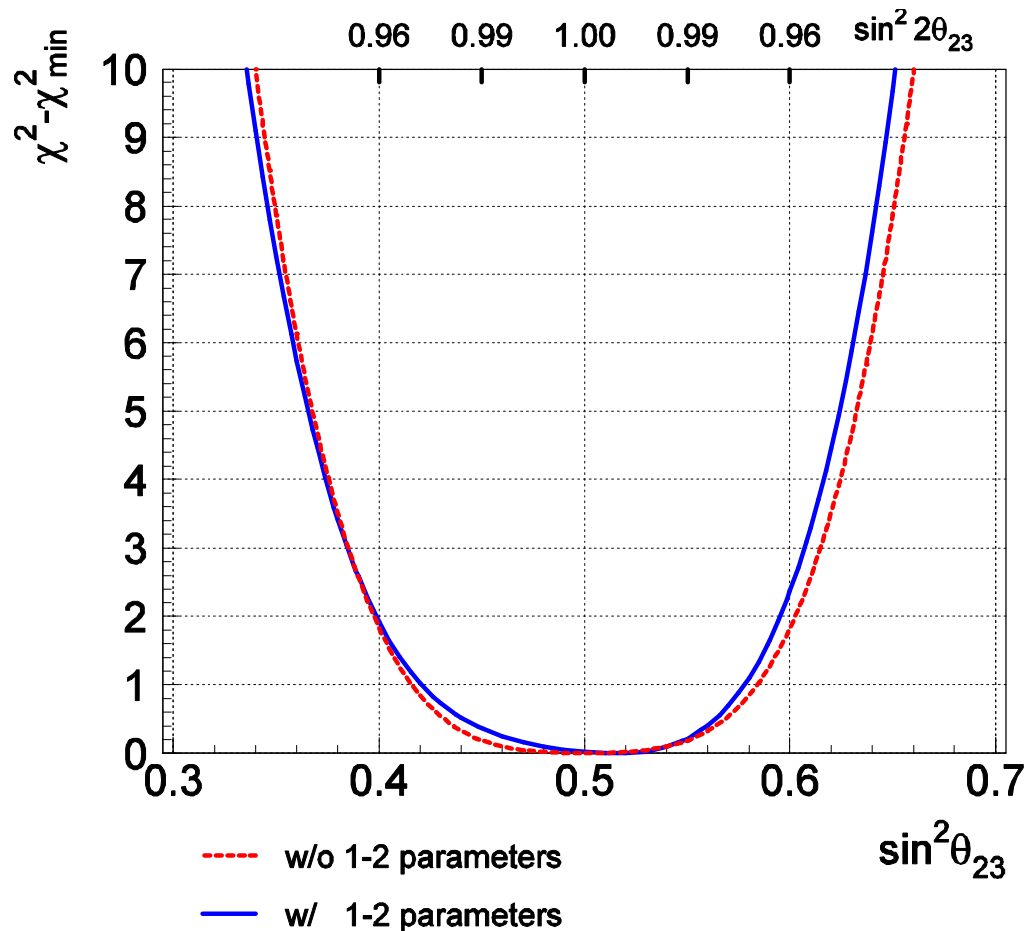
χ^2_{solar} from Solar- ν +KamLAND results are added in each $(\Delta m^2_{12}, \sin^2 \theta_{12})$ point.

deviation from the 2-3 full mixing ?

Result of $\sin^2 \theta_{23}$ determination

Preliminary

$\chi^2 - \chi^2_{\min}$ distribution as a function of $\sin^2 \theta_{23}$ where the other oscillation parameters are chosen to minimize χ^2



..... Solar terms off :

The 1-2 parameters (Δm^2_{12} , $\sin^2 \theta_{12}$) are fixed at zero.

best-fit : $\sin^2 \theta_{23} = 0.50$

— Solar terms on :

The 1-2 parameters are scanned.

best-fit : $\sin^2 \theta_{23} = 0.52$
($\sin^2 2\theta_{23} = 0.9984$)

Summary

- We have analyzed SK-I + SK-II atmospheric neutrino data in Super-Kamiokande
- Two-flavor oscillation analysis and L/E analysis were performed by these method:
 - Two-flavor oscillation without solar terms
 - $\sin^2 2\theta > 0.93$, $1.9 \times 10^{-3} \text{ eV}^2 < \Delta m^2 < 3.1 \times 10^{-3} \text{ eV}^2$ at 90 % C.L.
 - Oscillation with solar terms
 - No clear evidence for deviation of θ_{23} from $\pi/4$
 - L/E analysis
 - neutrino decay and decoherence hypotheses are excluded more significantly.