

# **Current Status of Solar Neutrinos at Super-Kamiokande**

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# Super-Kamiokande Collaboration

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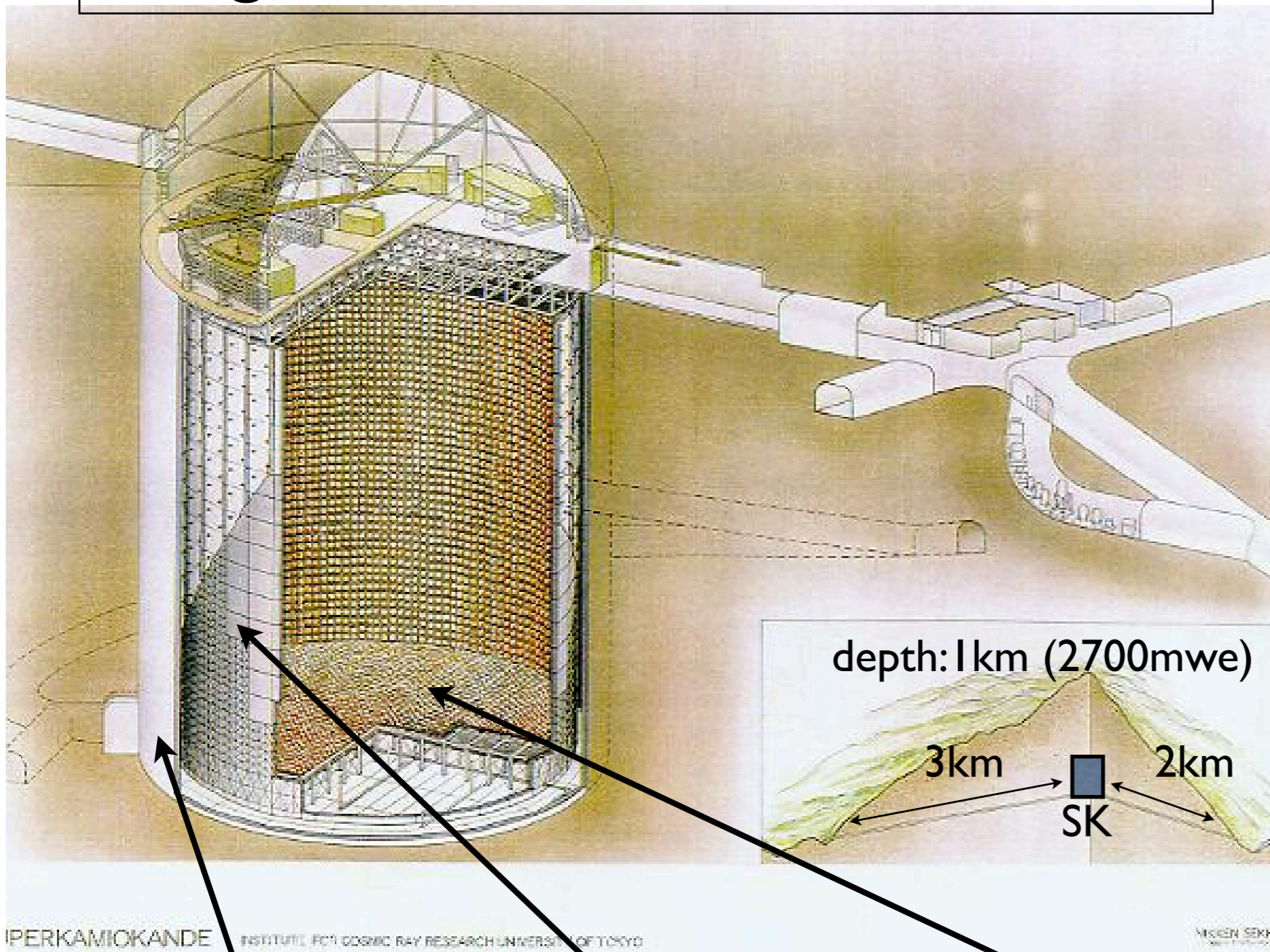
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4 Brookhaven National Laboratory, USA  
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8 Duke University, USA  
9 George Mason University, USA  
10 Gifu University, Japan  
11 University of Hawaii, USA  
12 Indiana University, USA  
13 KEK, Japan  
14 Kobe University, Japan  
15 Kyoto University, Japan  
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17 Louisiana State University, USA  
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19 Massachusetts Institute of Technology, USA  
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26 Osaka University, Japan  
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32 Tokai University, Japan  
33 Tokyo Institute for Technology, Japan  
34 University of Tokyo, Japan  
35 Warsaw University, Poland  
36 University of Washington, USA

**~ 130 collaborators**  
**36 institutions**  
**5 countries**

# The Super-Kamiokande Detector

Height: 41.4m    Diameter: 39.3m



## SK-I (1996~2001)

- 50000ton water
- ~11200 of 20inch PMTs
- Fid. vol. 22.5kt
- Photo coverage 40%
- Stopped by the accident in Nov. 2001

## SK-II (2002~2005)

- ~5200 of 20inch PMTs
- Photo coverage 19%

## SK-III (Jul. 2006~)

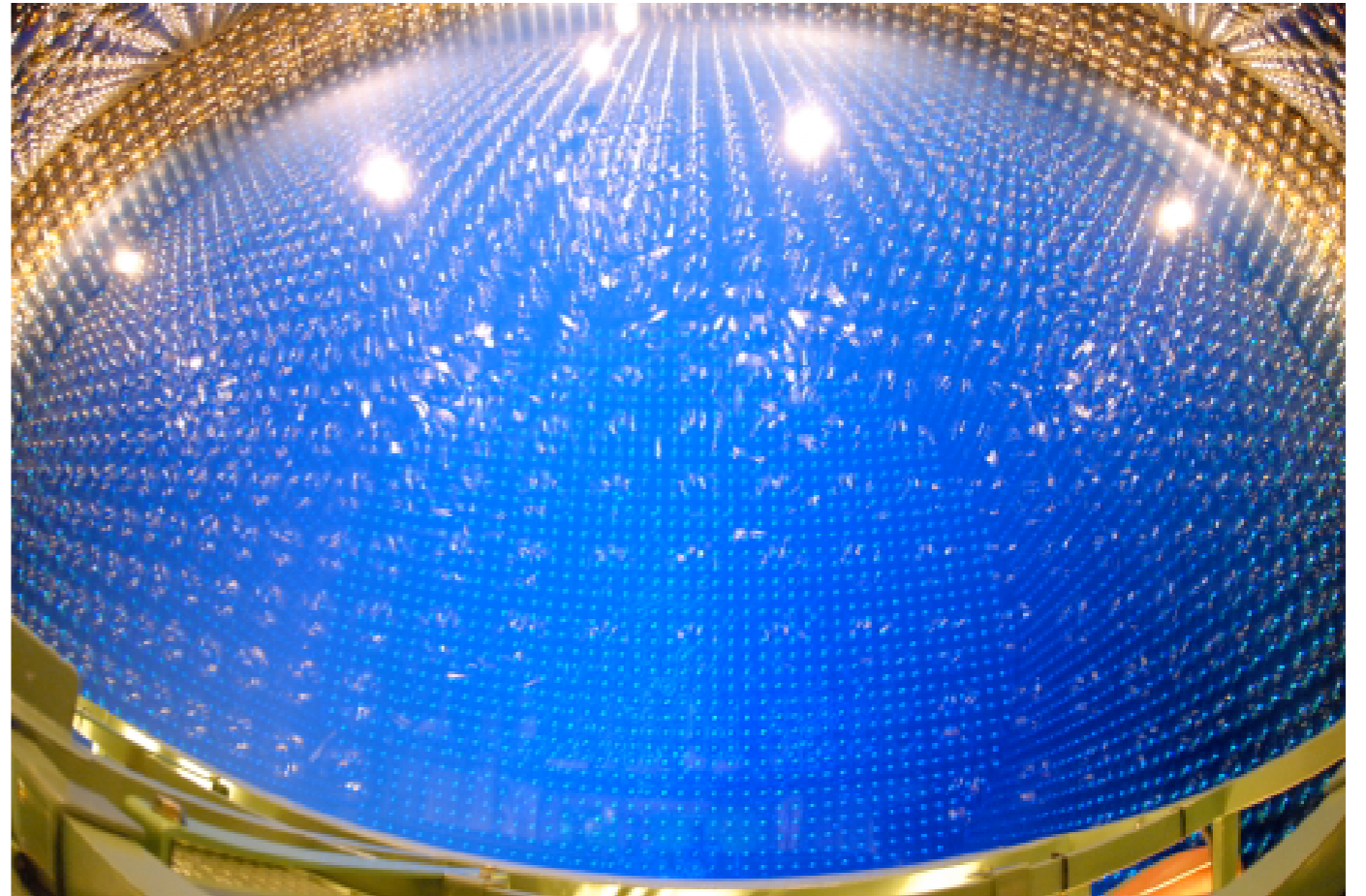
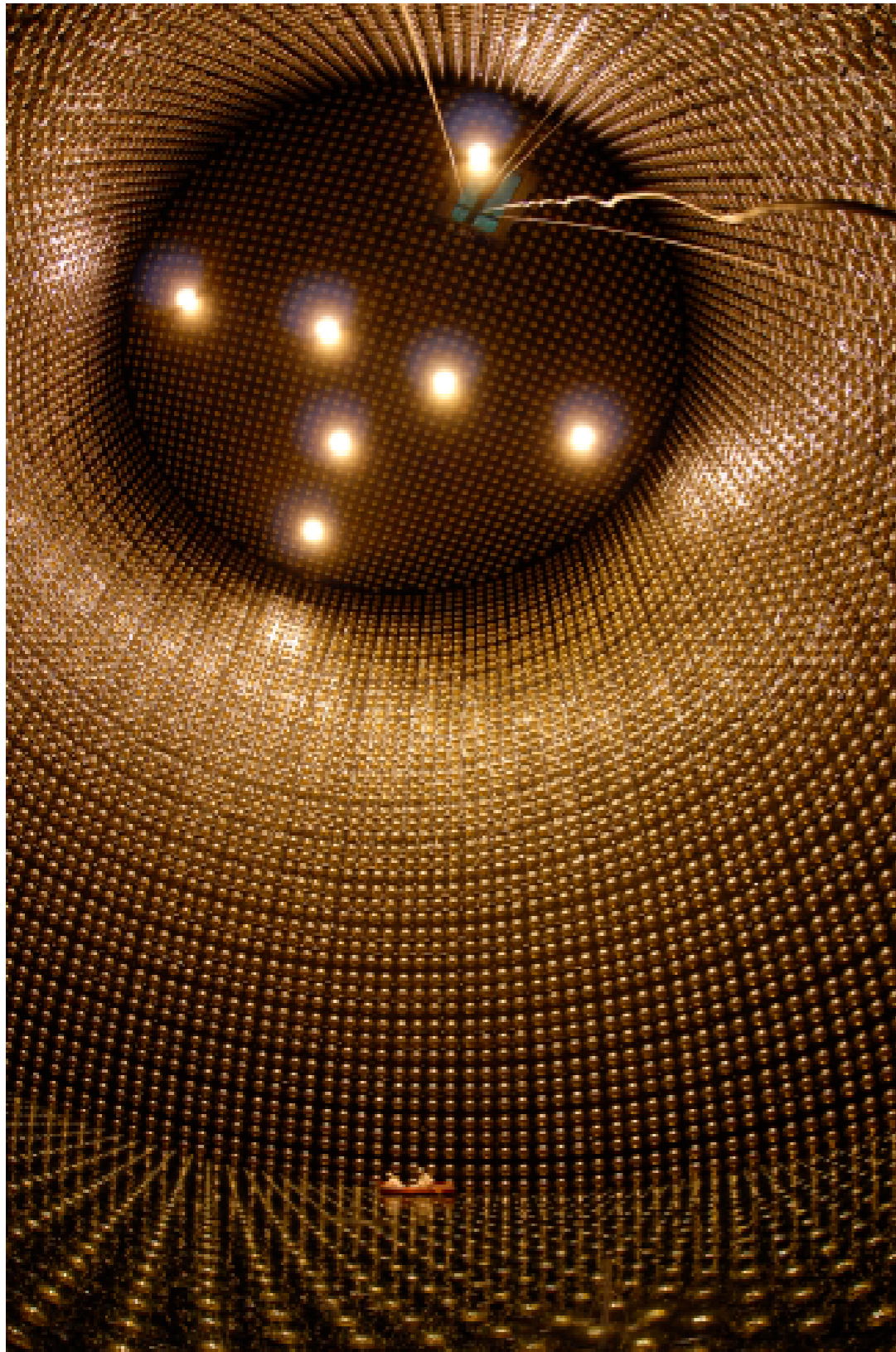
- 40% coverage
- OD Segmentation

50,000 ton  
stainless steel tank

Outer Detector (OD)  
1885 8in. PMTs

Inner Detector (ID)  
11129 20in. PMTs

# The Super-Kamiokande Detector



*April 12th, 2006*

**PMT mounting completed**

*July 11th*

**Tank filled**

*October 3rd*

**Water purification progressing**

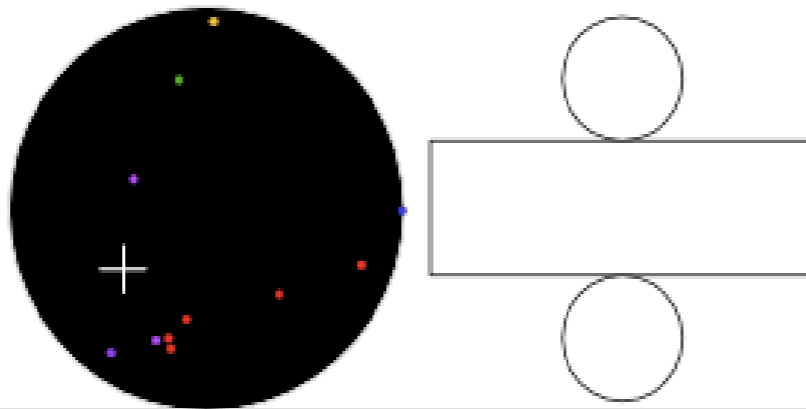


# Solar Neutrinos at Super-K

## Typical low energy event

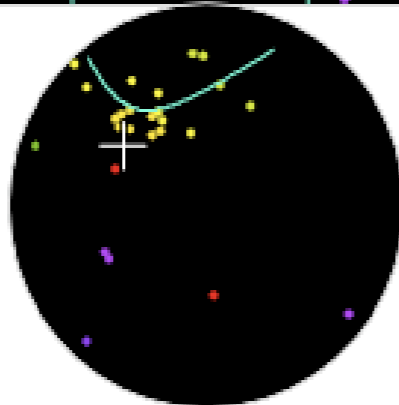
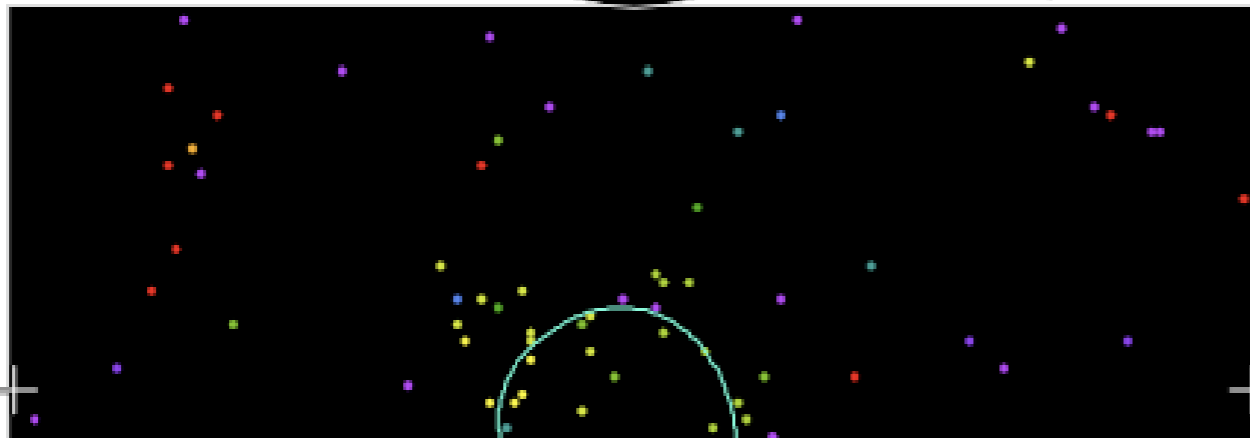
### Super-Kamiokande

Run 1742 Event 102496  
 96-09-31:07:13:23  
 Inner: 103 hits, 123 pE  
 Outer: -1 hits, 0 pE (in-time)  
 Trigger ID: 0x03  
 E= 9.086 GeV=0.77 GeV= 0.949  
 Solar Neutrino

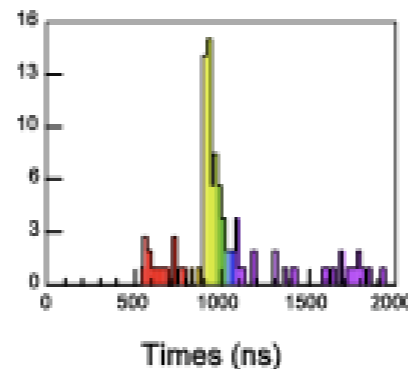


### Time (ns)

\* < 815  
 \* 815- 835  
 \* 835- 855  
 \* 855- 875  
 \* 875- 895  
 \* 895- 915  
 \* 915- 935  
 \* 935- 955  
 \* 955- 975  
 \* 975- 995  
 \* 995-1015  
 \* 1015-1035  
 \* 1035-1055  
 \* 1055-1075  
 \* 1075-1095  
 \* >1095



$E_e = 9.1 \text{ MeV}$   
 $\cos\theta_{\text{sun}} = 0.95$



$$\nu_x + e^- \longrightarrow \nu_x + e^-$$

$$\sigma(\nu_{\mu(\tau)} e^-) \approx 0.15 \times \sigma(\nu_e e^-)$$

## Timing information

Vertex position

Ring Pattern

Direction

Number of hit PMTs

Energy

Resolutions (for 10 MeV electron)

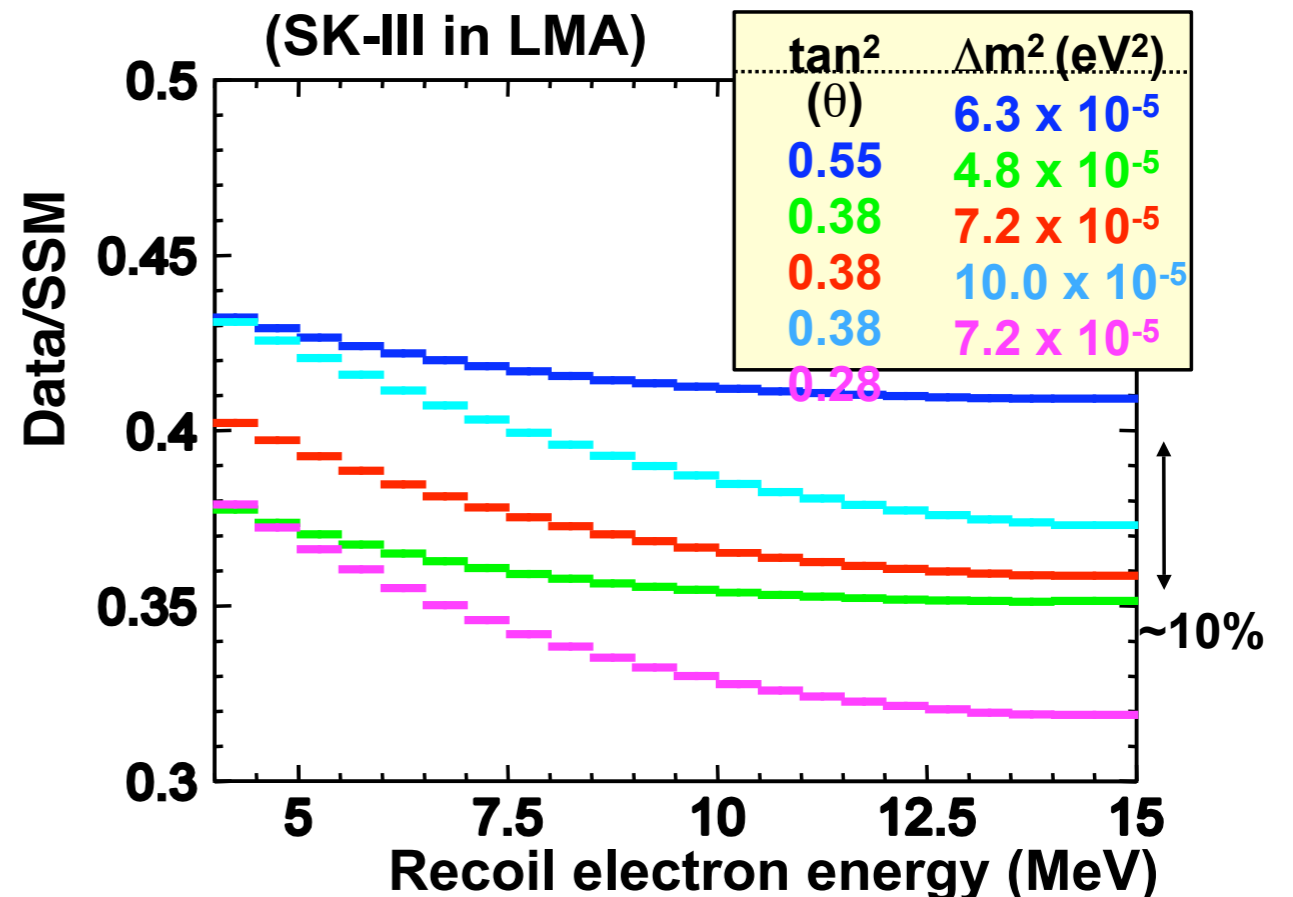
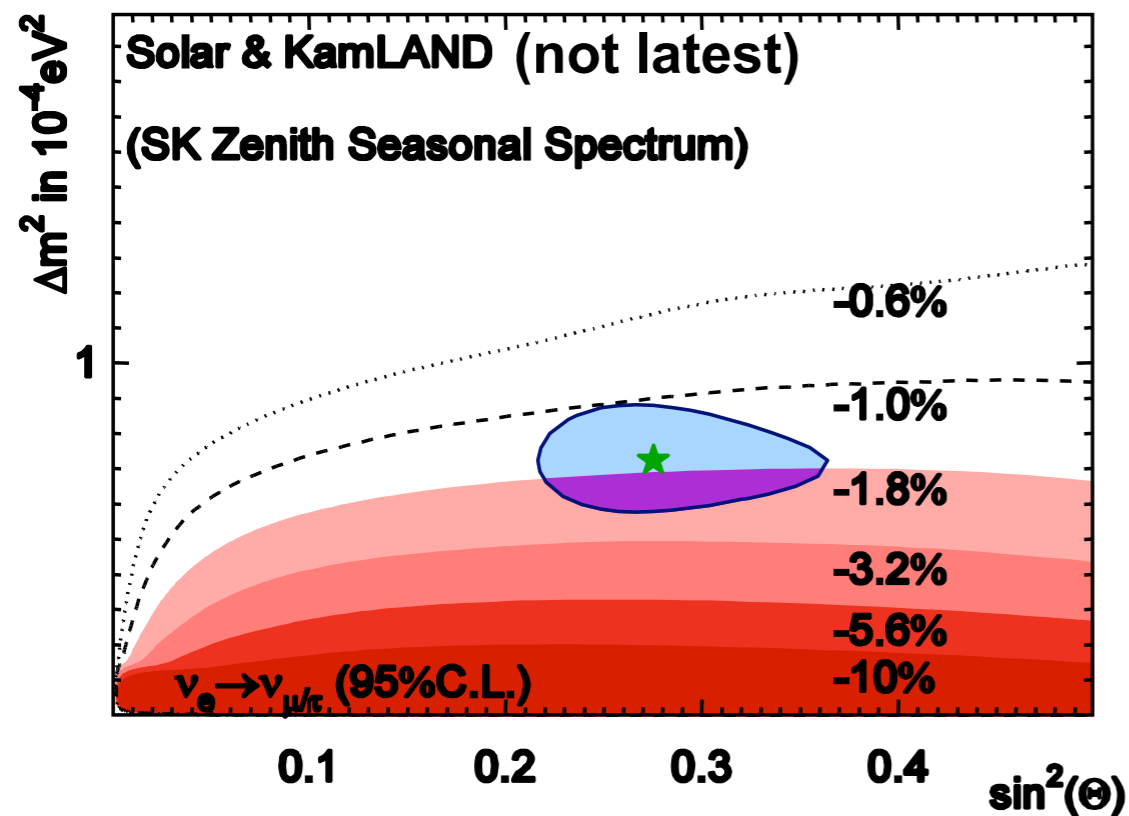
Energy: 14%

Vertex: 87 cm

Direction: 26 deg.

# Solar Neutrinos at Super-K

- **High Statistics** ~15 events/day with  $E_e > 5$  MeV ( $^8\text{B}$  & hep)
- **Time Variations** Day/Night, Seasonal
- **Energy Spectrum** Sensitive to  $\nu$  oscillation parameters
- **Precise Energy Calibration** by electron LINAC
- **Flux Independent Analysis**



# SK-II Latest Results

## Full SK-II run period

Analysis periods & energy thresholds

Dec. 24th, 2002 - July 15th, 2003 (159 days)

8.0 - 20. MeV

July 15th, 2003 - Oct. 5th 2005 (631 days)

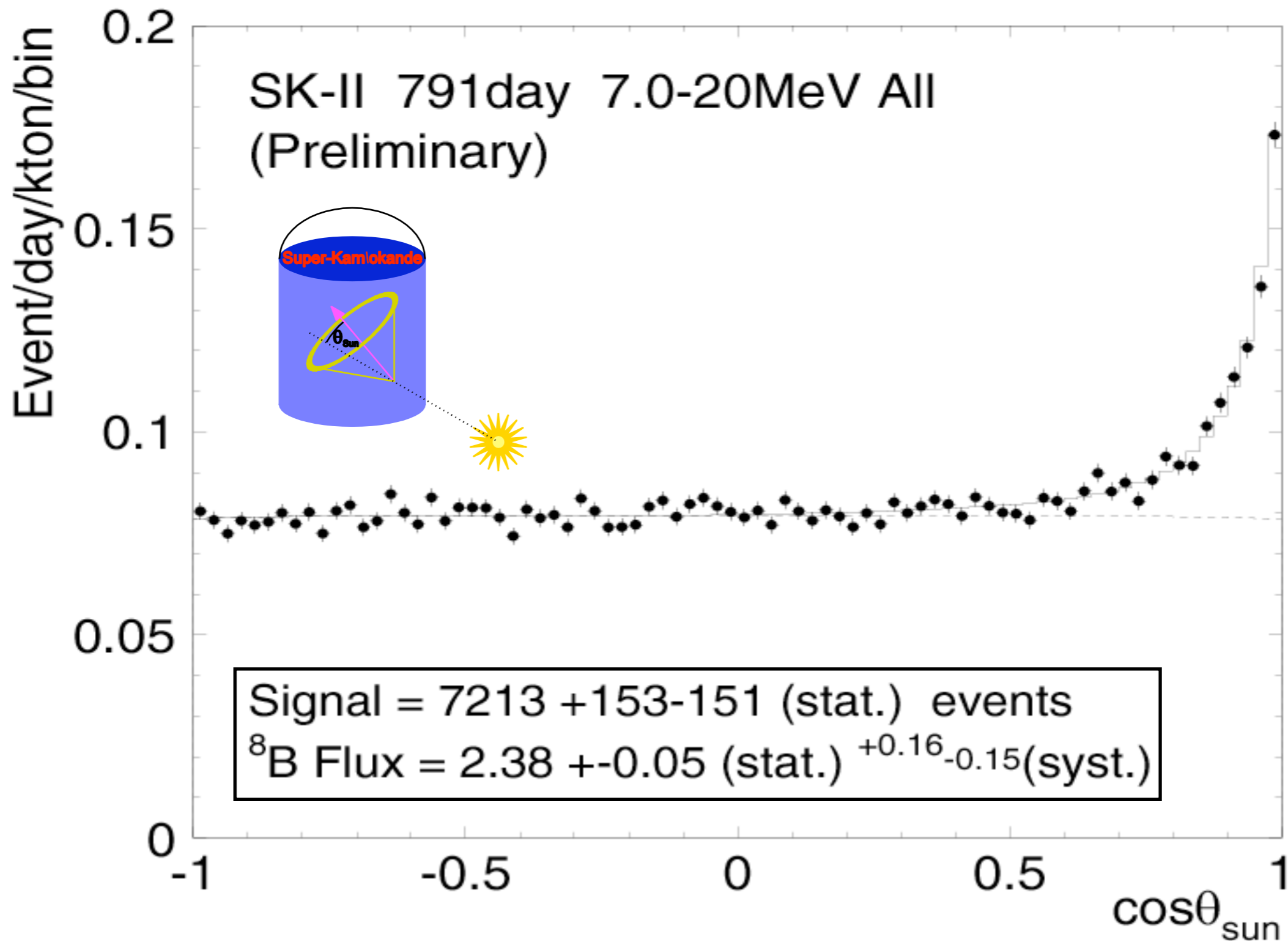
7.0 - 20.0 MeV for total flux

7.5 - 20.0 MeV for day/night

Total live time: 791 days

Systematic errors are under study

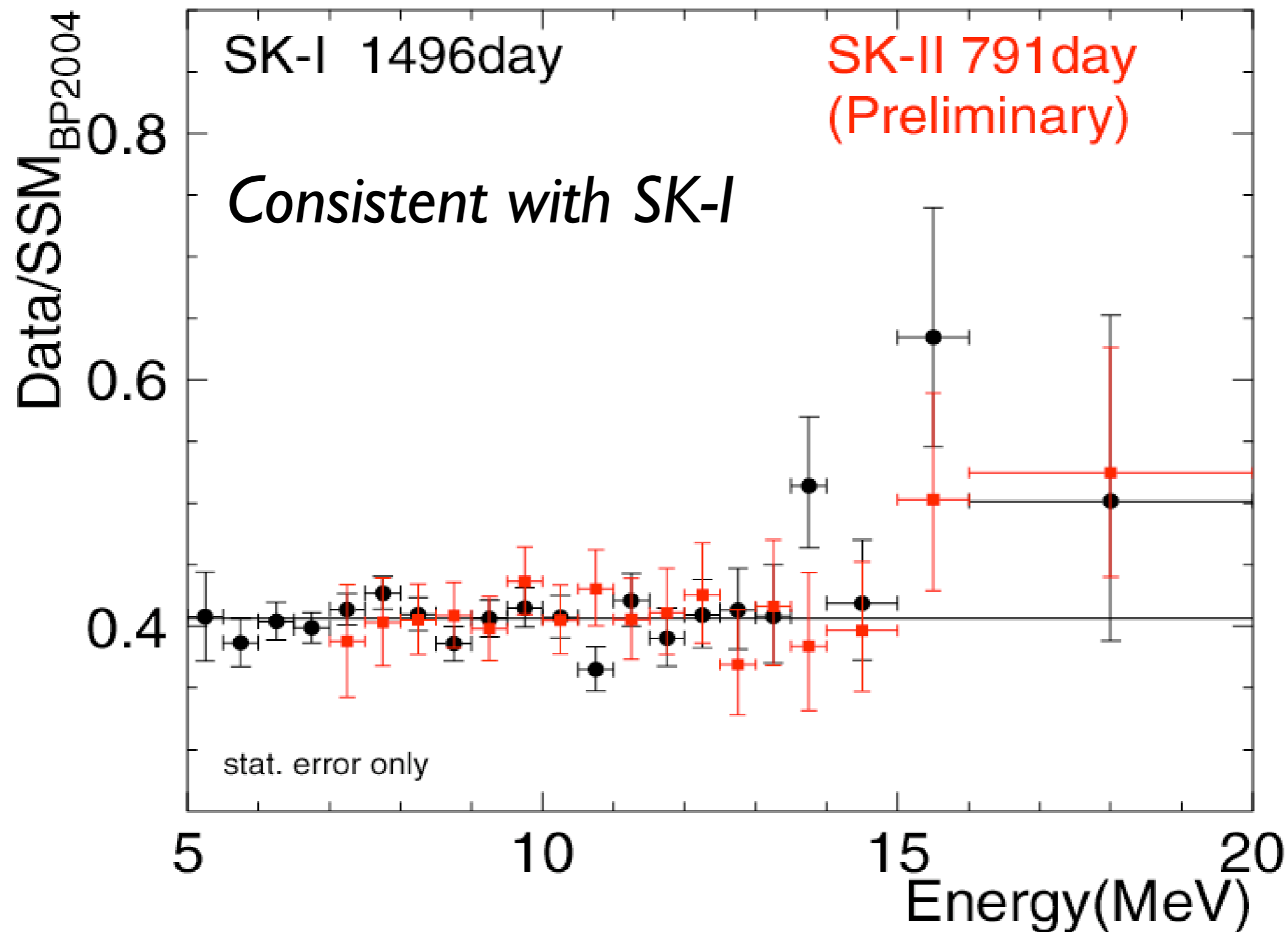
# SK-II Flux



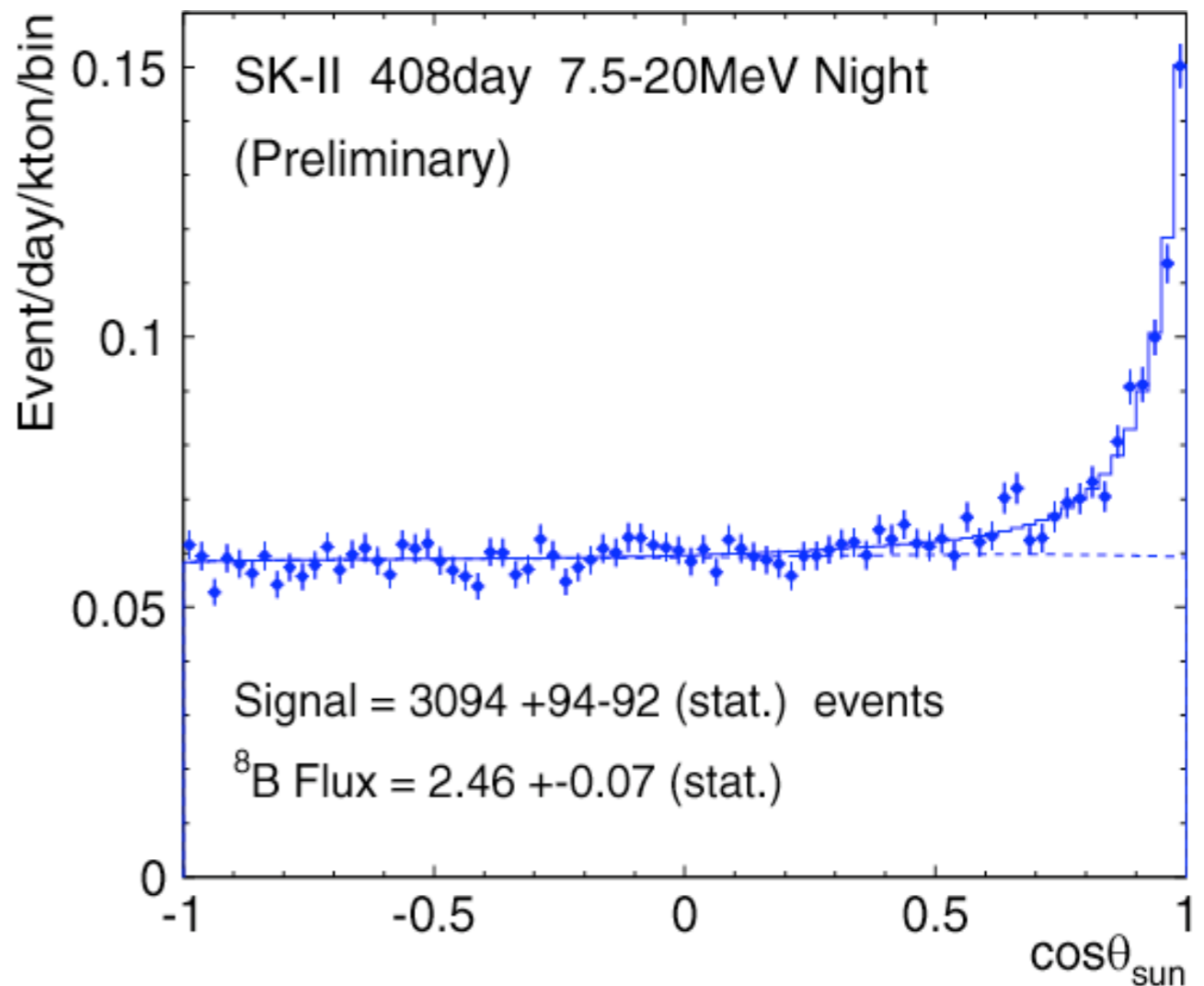
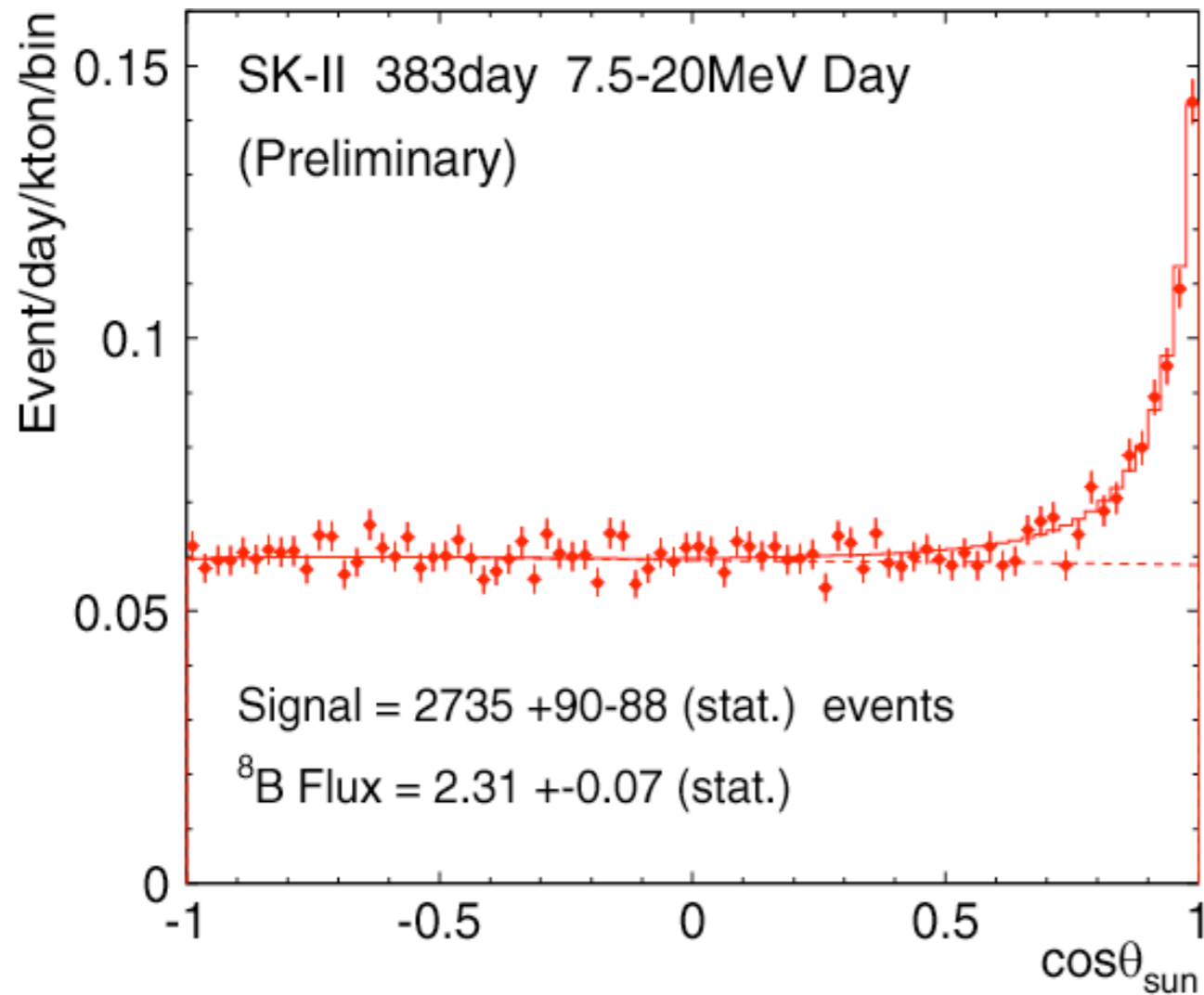
**SK-I result: 2.35 +/- 0.02(stat.) +/- 0.08(syst.)**



# SK-II Energy Spectrum



# Day/Night

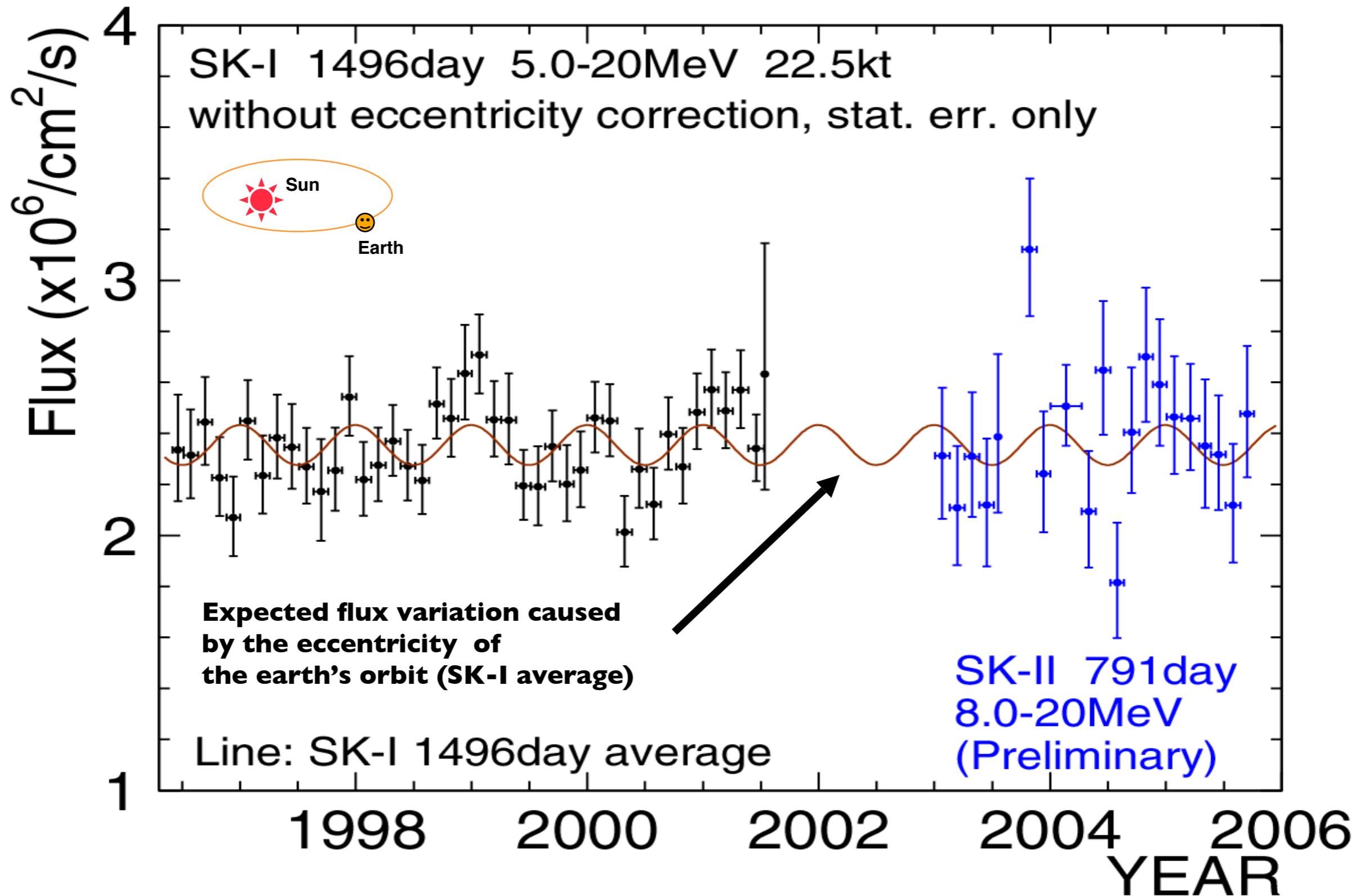


$$\frac{(\text{Day-Night})}{(\text{Day+Night})/2} = -0.064 \pm 0.043 (\text{stat.}) \quad \text{Preliminary}$$

(systematic errors are under study)

SK-I D/N Asymmetry:  $-0.021 \pm 0.020$   $^{+0.013}_{-0.012}$

# Time Variation



# SK Oscillation Analysis

$$\chi_{SK}^2(\beta, \eta, \delta_B, \delta_{SK1,S}, \delta_{SK1,R}, \delta_{SK2,S}, \delta_{SK2,R})$$

$$\begin{aligned}
 &= \sum_{i=1}^{21} \frac{(d_{SK1,i} - (\beta_m b_{SK1,i} + \eta_m h_{SK1,i}) f(E_i, \delta_B, \delta_{SK1,S}, \delta_{SK1,R}))^2}{\sigma_{SK1,i}^2} \\
 &+ \sum_{i=1}^{17} \frac{(d_{SK2,i} - (\beta_m b_{SK2,i} + \eta_m h_{SK2,i}) f(E_i, \delta_B, \delta_{SK1,S}, \delta_{SK1,R}))^2}{\sigma_{SK2,i}^2} \\
 &+ \left(\frac{\delta_B}{\sigma_B}\right)^2 + \left(\frac{\delta_{SK1,S}}{\sigma_{SK1,S}}\right)^2 + \left(\frac{\delta_{SK1,R}}{\sigma_{SK1,R}}\right)^2 + \left(\frac{\delta_{SK2,S}}{\sigma_{SK2,S}}\right)^2 + \left(\frac{\delta_{SK2,R}}{\sigma_{SK2,R}}\right)^2 \\
 &+ 2\Delta \log(\mathcal{L})
 \end{aligned}$$

$$d_i = \frac{\text{Data}_i}{{}^8 B_i^{\text{SSM}} + \text{hep}_i^{\text{SSM}}}, \quad b_i = \frac{{}^8 B_i^{\text{osc}}(\Delta m^2, \tan^2 \theta)}{{}^8 B_i^{\text{SSM}} + \text{hep}_i^{\text{SSM}}}, \quad h_i = \frac{\text{hep}_i^{\text{osc}}(\Delta m^2, \tan^2 \theta)}{{}^8 B_i^{\text{SSM}} + \text{hep}_i^{\text{SSM}}}$$

# SK Oscillation Analysis

$$\chi_{SK}^2(\beta, \eta, \delta_B, \delta_{SK1,S}, \delta_{SK1,R}, \delta_{SK2,S}, \delta_{SK2,R})$$

flux factors

energy correlated systematic error

$$\begin{aligned} &= \sum_{i=1}^{21} \frac{(d_{SK1,i} - (\beta_m b_{SK1,i} + \eta_m h_{SK1,i}) f(E_i, \delta_B, \delta_{SK1,S}, \delta_{SK1,R}))^2}{\sigma_{SK1,i}^2} \\ &+ \sum_{i=1}^{17} \frac{(d_{SK2,i} - (\beta_m b_{SK2,i} + \eta_m h_{SK2,i}) f(E_i, \delta_B, \delta_{SK1,S}, \delta_{SK1,R}))^2}{\sigma_{SK2,i}^2} \\ &+ \left(\frac{\delta_B}{\sigma_B}\right)^2 + \left(\frac{\delta_{SK1,S}}{\sigma_{SK1,S}}\right)^2 + \left(\frac{\delta_{SK1,R}}{\sigma_{SK1,R}}\right)^2 + \left(\frac{\delta_{SK2,S}}{\sigma_{SK2,S}}\right)^2 + \left(\frac{\delta_{SK2,R}}{\sigma_{SK2,R}}\right)^2 \end{aligned}$$

$$+ 2\Delta \log(\mathcal{L})$$

time variation

predicted oscillated spectrum

$$d_i = \frac{\text{Data}_i}{{}^8 B_i^{\text{SSM}} + \text{hep}_i^{\text{SSM}}}, \quad b_i = \frac{{}^8 B_i^{\text{osc}}(\Delta m^2, \tan^2 \theta)}{{}^8 B_i^{\text{SSM}} + \text{hep}_i^{\text{SSM}}}, \quad h_i = \frac{\text{hep}_i^{\text{osc}}(\Delta m^2, \tan^2 \theta)}{{}^8 B_i^{\text{SSM}} + \text{hep}_i^{\text{SSM}}}$$

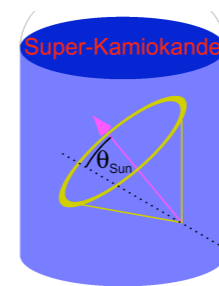
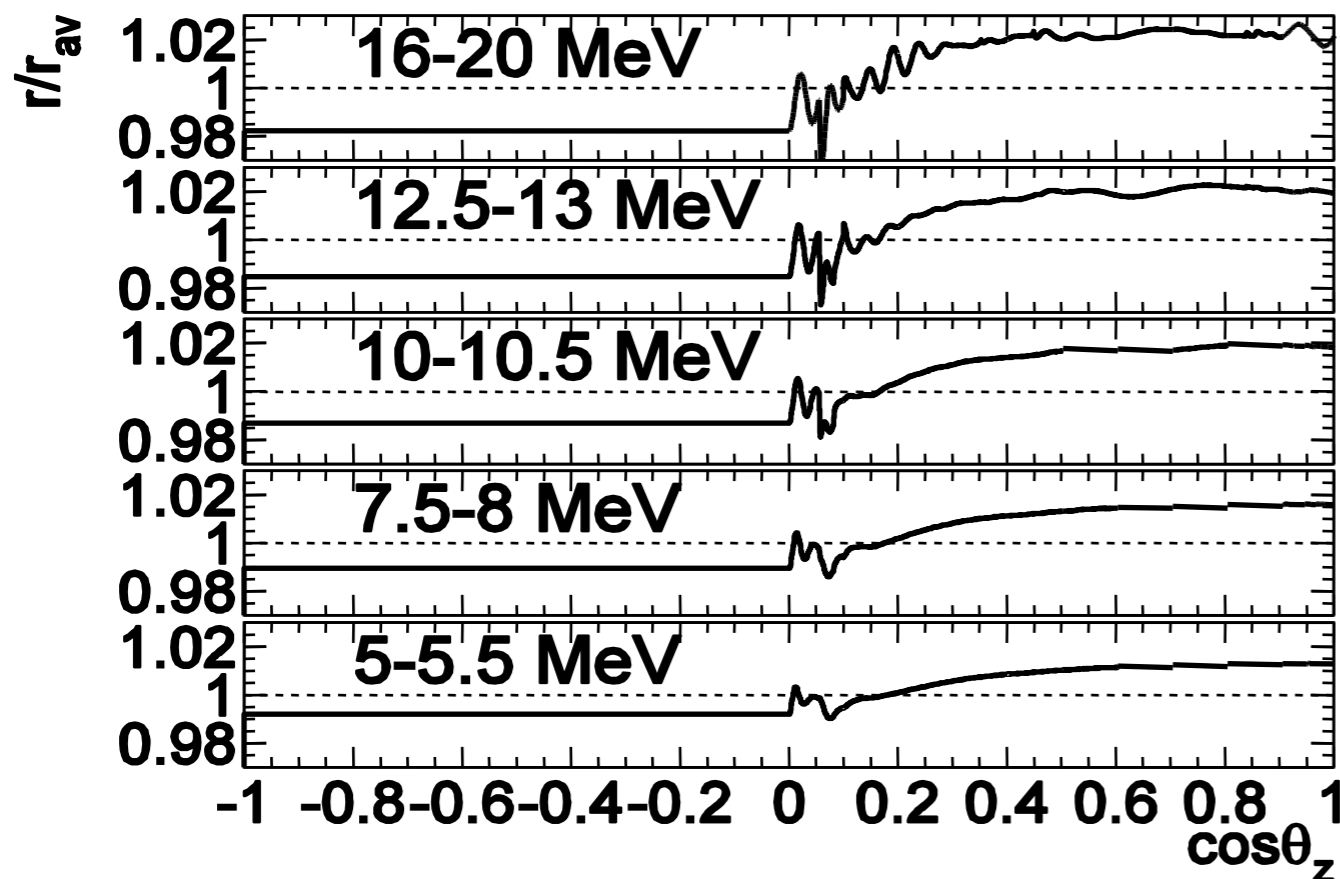
unoscillated spectrum

# Unbinned Time Variation

$$2\Delta \log(\mathcal{L})$$

$$\mathcal{L} = e^{-(\sum_i B_i + S)} \prod_{i=1}^{N_{bin}} \prod_{j=1}^{n_i} (B_i \cdot b_{ij} + S \cdot Y_i \cdot \underbrace{p(\cos \theta_{ij}, E_j) \cdot z_i(t_j)}_{\text{---}})$$

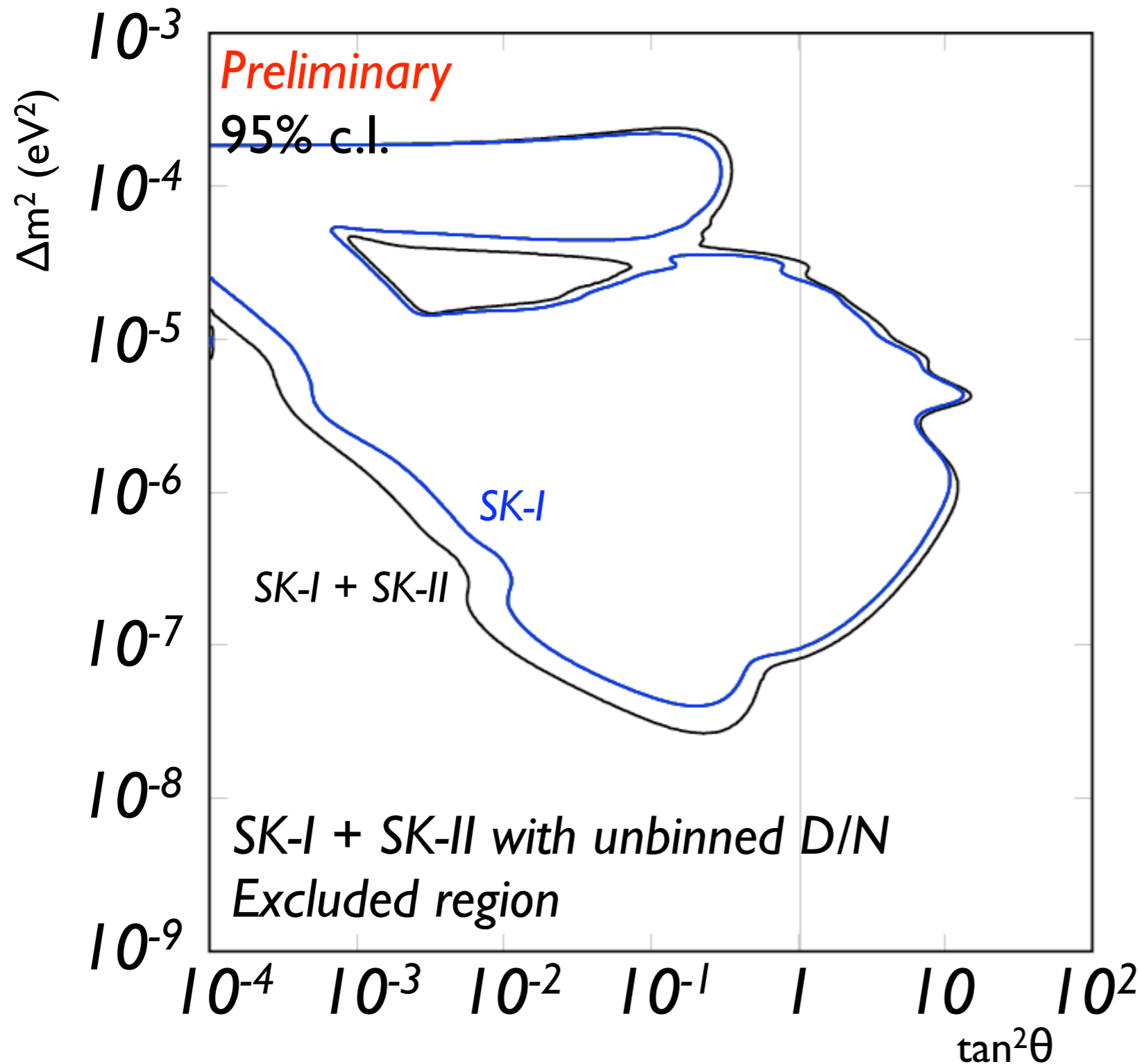
Predicted solar zenith angle variations ( $\Delta m^2 = 6.3 \times 10^{-5} \text{ eV}^2, \tan^2 \theta = 0.55$ )



SK zenith angle  $\cos \theta_{ij}$  of the  $j^{\text{th}}$  event in the  $i^{\text{th}}$   $E$  bin



# SK-I, SK-II Oscillation Analysis



## <sup>8</sup>B, hep free fit

### Common:

- <sup>8</sup>B flux scale
- Hep flux scale
- <sup>8</sup>B shape error

### SK-I only:

- Energy resolution
- Energy scale

### SK-II only:

- Energy resolution
- Energy scale

# SK-I, SK-II Global Analysis

$$\chi_{global}^2(\beta, \eta) = \chi_{SK}^2(\beta, \eta) + \chi_{SNO}^2(\beta, \eta) + \frac{(ADN_{CC} - ADN_{pred})^2}{\sigma_{ADN_{CC}}^2} + \chi_{radiochem}^2(\beta, \eta)$$

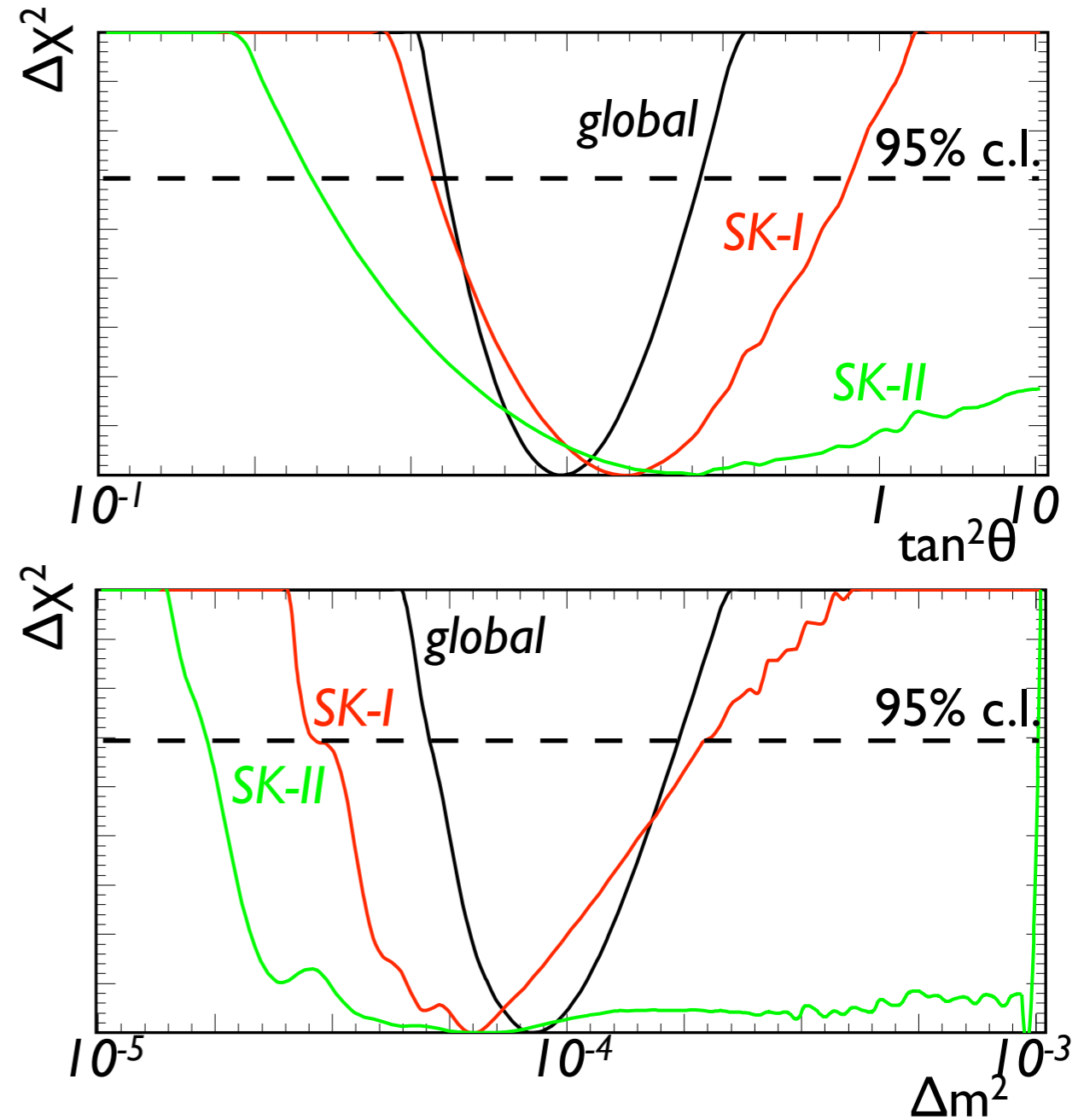
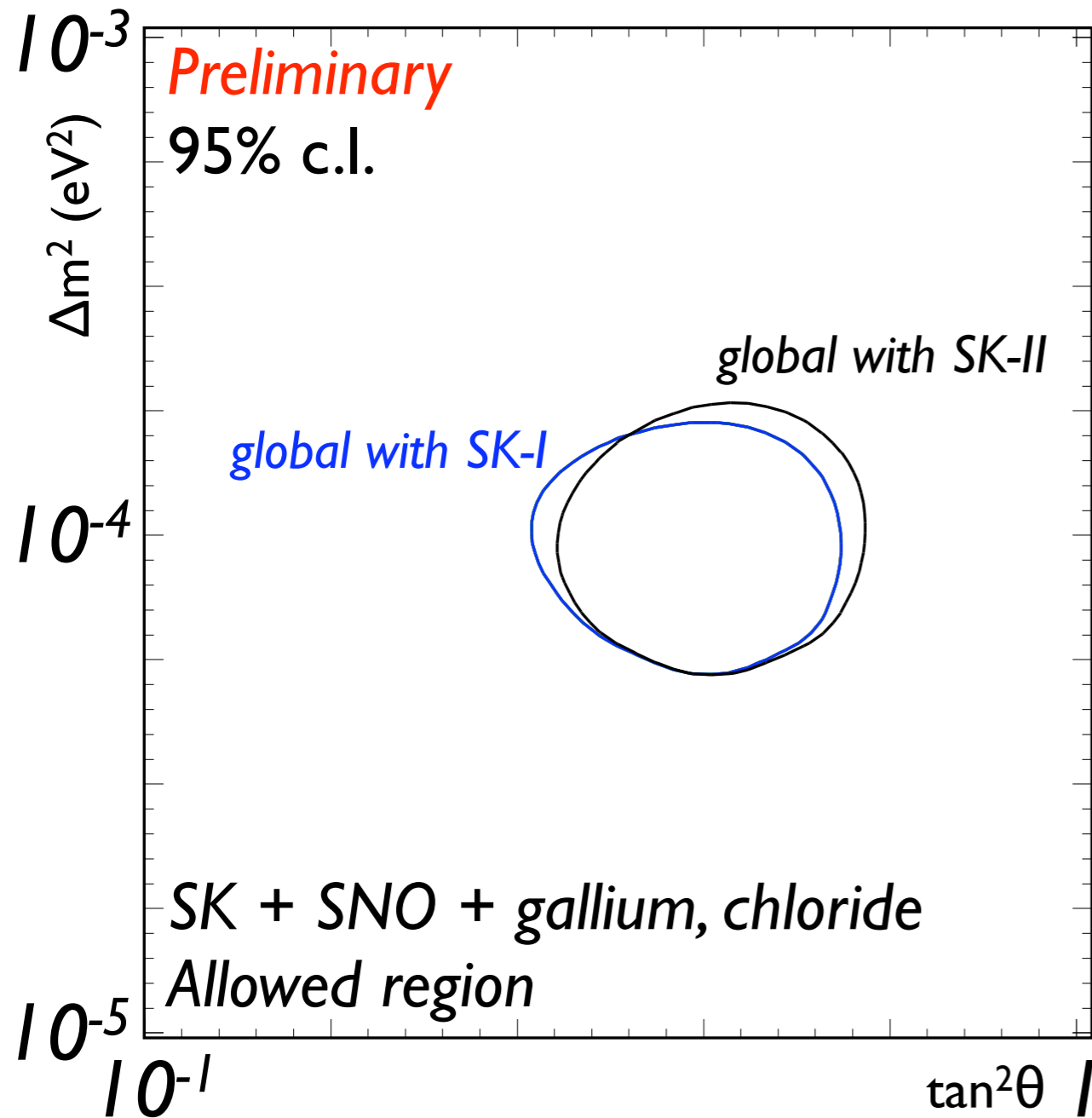
## SNO data:

Absolute flux fit (spectrum information summed)

Fluxes from 371-day Salt Phase (CC & NC)

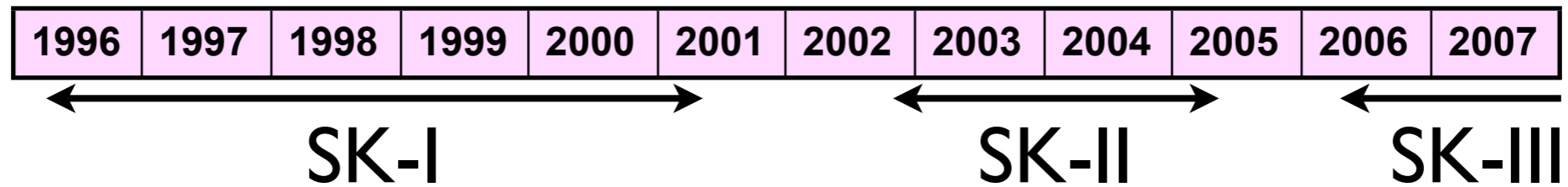
ADN from 306-day pure D<sub>2</sub>O Phase with NC=0

# SK-I, SK-II Global Analysis



SK-I, SK-II global fit:  $\Delta m^2 = 8.71 \times 10^{-5}$  (eV<sup>2</sup>)  
 $\tan^2\theta = 0.38$

# Future Plan In SK-III



Water filling was finished on July 11th, 2006  
Now, detector calibration is taking place

ID PMT:

SK-II = ~5,200

SK-III = 11,129

**Original energy & vertex resolutions  
for low-energy events is restored**

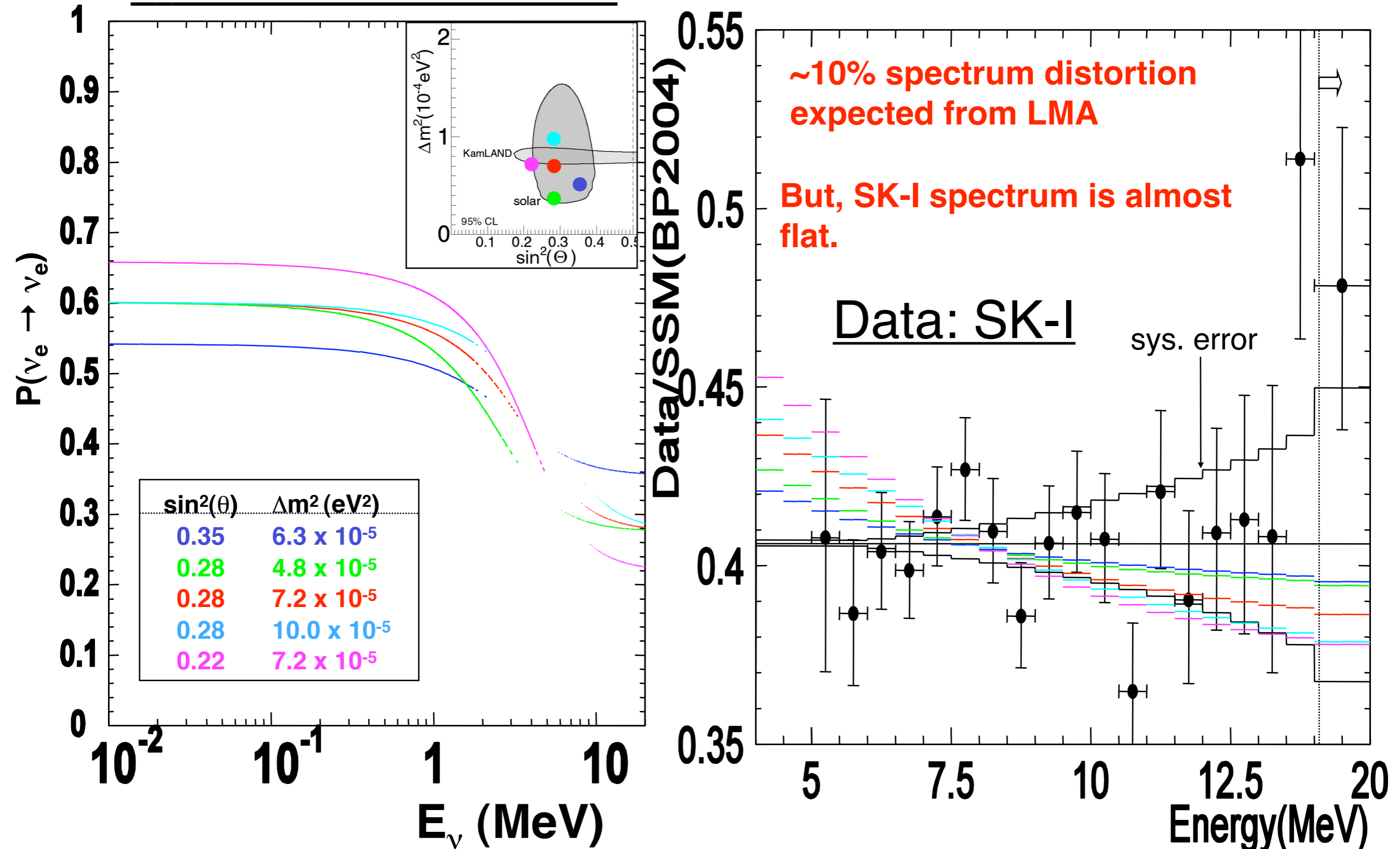
Goals

Solar neutrinos below 5.0 MeV with improved analysis tools  
and lower Rn background  
Precise study on spectrum distortion

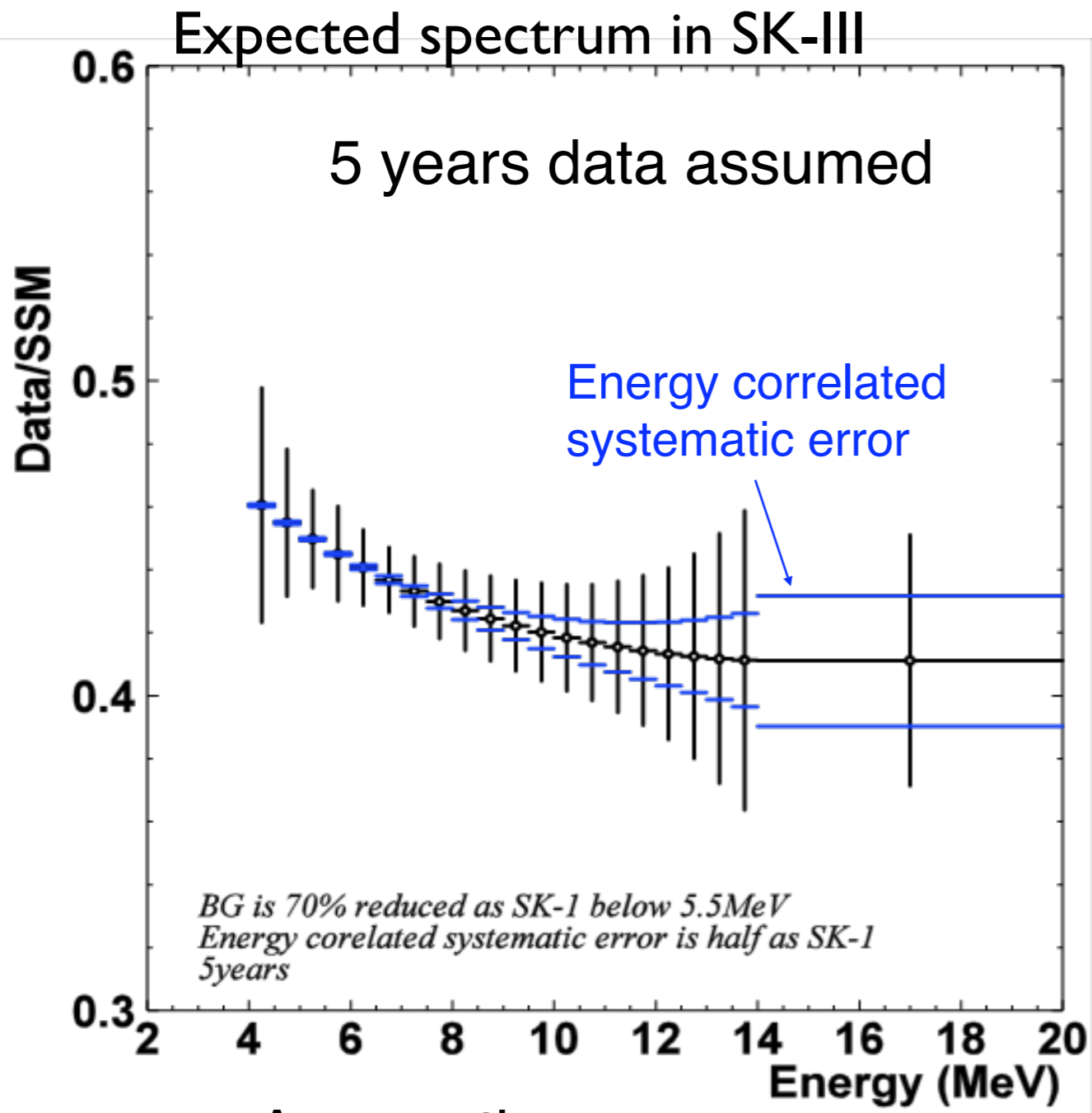
# Energy Spectrum Distortion

$\nu_e$  survival probability

Recoil electron spectrum



# Energy Spectrum Distortion

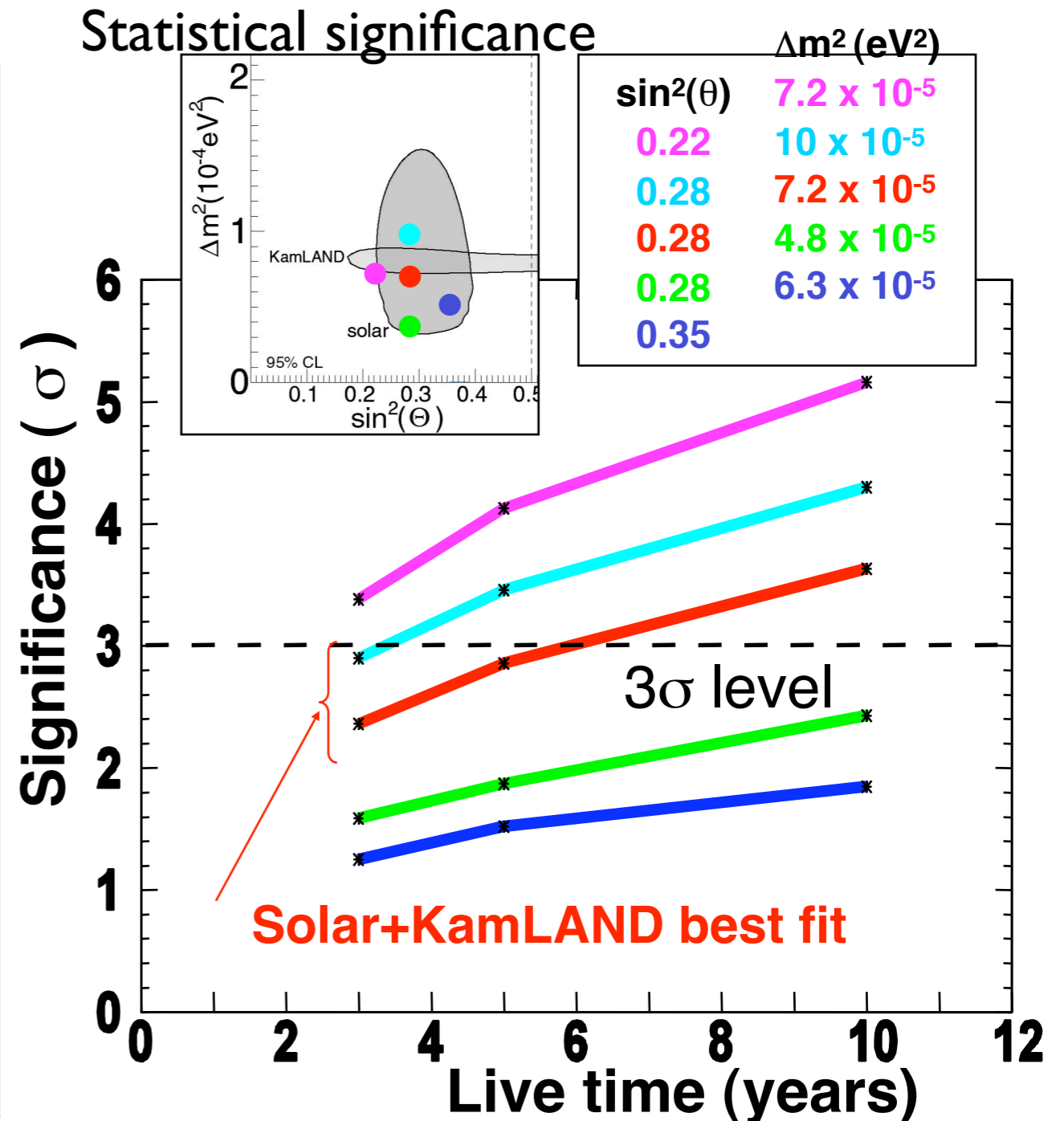


Assumption:

Correlated systematic error:  $\times 0.5$

4.0-5.5MeV background :  $\times 0.3$  of SK-I

(> 5.5MeV is same as SK-I)





# Summary

- SK-III has been started and is taking data
- SK-II data has been updated to its final 791 d
- Oscillation analysis with SK-I, SK-II data has been performed and shows consistency with final SK-I data set
- SK-II shows consistency within the global analysis
- Hope to see energy spectrum distortion in SK-III

*thank you*