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 Monitoring of nuclear reactors against illicit operation or fuel diversion: present proposals using conventional 1-ton detectors reach only > ~3 GWt reactor power

 Geological prospection, planetary tomography... the list gets much wilder.



collar JMPRPPC, Oct. 06

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2005: Geoneutrinos detected.

Dawn of the applied neutrino physics era?

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Collar JMPRPPC, Oct. 06

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As Reactor Fuel Burns the Composition Changes



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Three legged stool needed: mass, threshold, background

for Cosmological Physics

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No "light-bulb" moment: 5 years of R&D at UC



name-of-the-game: detection of << 1 keV recoils with large (> 1 kg) detectors

(25 y and counting... must use new technologies or at least alterations)



13 cm² APD on a 2 inch diameter silicon wafer



Single-photon pulses using LN₂ cooled LAAPDs (high QE)



Mass-produced 3M-UoC GEM and single-electron signals from quadruple GEM



Start with the foundations: ultra low-energy recoil calibrations at KSU reactor



Start with the foundations: ultra low-energy recoil calibrations at KSU reactor







Ti post-filter "switches off" the recoils, leaving all backgrounds unaffected

MCNP-POLIMI simulation





Modified-electrode $\underline{p-type}$ HPGe: A new tool in astroparticle & neutrino physics

Kavli Institute









Modified-electrode $\underline{p-type}$ HPGe: A new tool in astroparticle & neutrino physics

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The energy resolution and large mass of a HPGe plus the noise and threshold of a tiny x-ray detector???





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Modified-electrode $\underline{p-type}$ HPGe: A new tool in astroparticle & neutrino physics

Kavli Institute





From: otench@canberra.com To: "Juan I. Collar" <collar@uchicago.edu> Date: Sat, 3 Dec 2005 16:41:10 -0500 X-Uchicago-PMX-Id: 192.153.25.189: jB3LtheQ001154 [Sat Dec 3 15:55:44 2005] X-Uchicago-Spam: Gauge=XXI, Probability=21%

Hello Juan,

We just got the first results in and they seem to be outstanding. The pulse resolution is about 160eV(FWHM) and Co-60 is well under 2.0 KV(FWHM). The detector should be shipping from France soon- in time for Christmas. It is too late now to change hardware but this mignt be done in future.

Best regards, Orren

Developed during 2005 by CANBERRA/EURYSIS (the one of three contacted companies up to the challenge) Funded by NNSA.







<u>Mass and threshold in place for reactor experiment</u>, background... almost there (anti-Compton shield & Al part replacement underway)

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Extensive detector characterization early 2006 (submitted to Phys. Rev. C)





Extensive detector characterization early 2006 (submitted to Phys. Rev. C)



Shielding studies at 6 m.w.e. (comparable to reactor site)





Extensive detector characterization early 2006 (submitted to Phys. Rev. C)







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Relevant ROI for power reactor experiment has been explored





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Definitive check on DAMA soon X50 improvement from clean Al, x10 from anti-Compton



First Physics Results expected Fall 06:

These light WIMPs remain compatible with DAMA & all other searches (accelerator bounds are model-dependent)

Next: replacement with <0.2 ppb U cryostat, develop low-bckg version of anti-Compton shield... and deploy to power reactor.



Nothing radioclean yet (need cash! \otimes)





What else can you do with such a detector? J.I. Collar PRD 59 063514



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Damour & Krauss PRL 81 5726

MAJORANA: can we avoid segmentation altogether? (cost, speed, simplicity, much lower front-end backgrounds) Does this device have anything to offer in a ββ context?

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AT THE UNIVERSITY OF CHICAGO **Optimal E-resolution** allows one to think $\beta\beta$ T = 77 Kpulser spectrox - UNI CHICAGO - CO 60@ 12 ÁS.SPE Energie: 1173.2 Co-60 n-type 1989 MH: 1.71 ~15 keV FWHM Brut: 73199.1 01.00E+6 LMH: 69446. 1.82 Isotope: Brut: CO-60 6145 Net: 6104 Isotope: 59.7FWHM 0.48position peak position (keV) FWHM (keV) 59.60.44 Compare resolution with n-type 1989 (now same as coax HPGe of same mass, 1.8KeV FWHM) 0.4059.50.36 Also x20 improvement in charge collection 59.4-0.3210 15202530 35 5 distance from closed end, d (mm) in going to p-type

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How does a multiple-site interaction look in a modified-electrode HPGe?









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Kavli Institute for Cosmological Physics At THE UNIVERSITY OF CHICAGO Better signal acceptance / background rejection than an 8-segment clover HPGe (even before close packing!) all with a single-channel device















Kavli Institute Advantages of single channel p-type modified for Cosmological Physics at THE UNIVERSITY OF CHICAGO electrode vis-à-vis segmentation for MAJORANA:

- Very efficient PSA rejection of multiples. All with one channel.
- Excellent energy resolution (1.8 keV Co-60, may drop some as noise is further improved)
- Increase speed of deployment/manufacture <u>as long as</u>... (is gradient of impurities reproducible? How important?)
- Increase simplicity of construction and analysis (one channel)
- Decrease cost (detectors and DAQ). Improve production time (cosmogenics)
- Decrease front end-associated radioactive backgrounds, thermal load, photon path.
- Increase stability (prototype performance stable for >5 continuous mo. and counting)
- Intrinsic to p-type: ruggedness (a must when arraying) and decreased sensitivity to surface contaminations.
- Several others (e.g., rejection of ALL alphas via PIXE -studies underway-)
- CANBERRA and PHDs Co. receptive to further fabrication (and further work on noise reduction).







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Disadvantages:

- Technology too new: too many unknowns in reproducibility, cost, speed of production, largest crystal size that can be produced, waste (important for $\beta\beta$), etc.
- Canberra's position: We need to build 6–10 more to know (they admit "lucking out". Hopefully this will not change).

Solution:

- This fits perfectly with planned coherent v program. Recent NSF/DOE proposal centered around this theme <- Help from rest of MAJORANA collaboration to maximize <u>synergy</u>: PNNL already funded to build more of these, ORNL seeking funding.
- Several kg of modified-electrode p-type HPGe's built by 2007!
- GOAL: Be by early next year counting at the Columbia Generating Station (Richland, WA, 12 mi. from PNNL) and simultaneously further developing the technique (i.e., building more of these). San Onofre? (offers more depth)











Reactor Monitoring: Right technological timing (HPGe technology flourishing:

Kavli Institute for Cosmolog Segmentation, encapsulation, arrays and (silent) mechanical cooling)

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CLUSTER

MAJORANA



RHESSI





GERDA





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New generation of recondensing Dewars add no microphonic noise and need topping (not refilling) every ~ 1yr (can be filled from N2 gas cylinder!) Ideal for reactor deployment.



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CLUSTER array for EUROBALL (7 encapsulated HPGe detectors)

Hexagonal tapering - diam .: 70 mm - height: 78 mm FWHM resolution : < 2.3 keV Efficiency: > 55% Alu wall thickness: 0.7 mm Cap-to-Ge distance: 0.7 mm.

11 kg, encapsulated, single cold finger (CANBERRA)

With 100 eV threshold, the equivalent of ~1 ton liquid scintillator (plus an additional ~x10 in rate beyond!)

A reality

fast approaching?



SANDS Sees Reactor Turn-on in Detail (Antineutrino Rate, Running Average)





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Coherent neutrino detection:



I want to believe!



