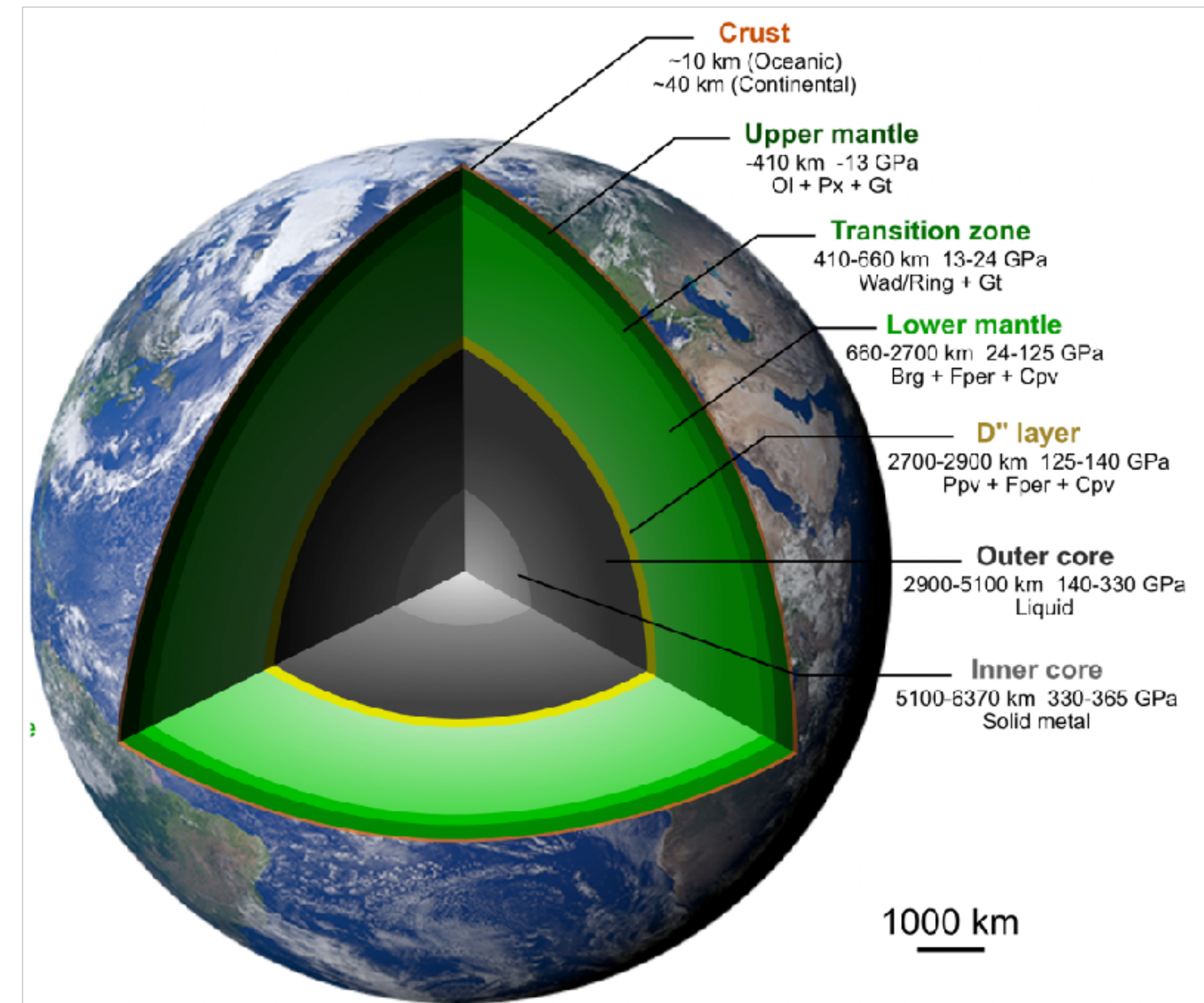


Geoneutrinos: KamLAND, Hanohano, OBD

~ Exploring the Earth's deep interior with Geoneutrinos ~

Hiroko Watanabe,
Research Center for Neutrino Science,
Tohoku University

April 28-May 2, 2025, University of Hawaii



Member of KamLAND experiment since 2007

After KamLAND started in 2000

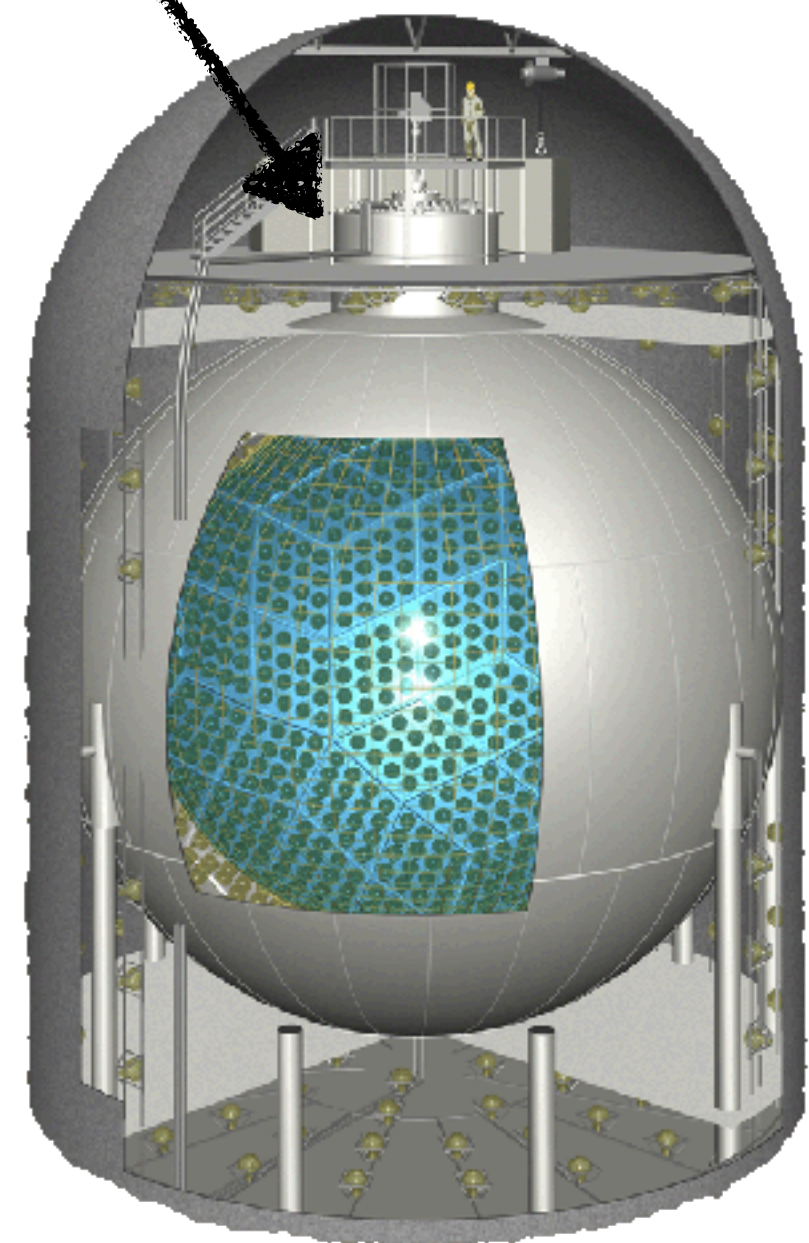
After 1st geo-neutrino paper publication in 2005

After “Hanohano” idea started to be discussed around 2005



My research:

- * Neutrino observation with KamLAND
- * “Neutrino Geoscience”
: interdisciplinary science field
→ **Ocean Bottom Detector**



with John

2/23

in 2017

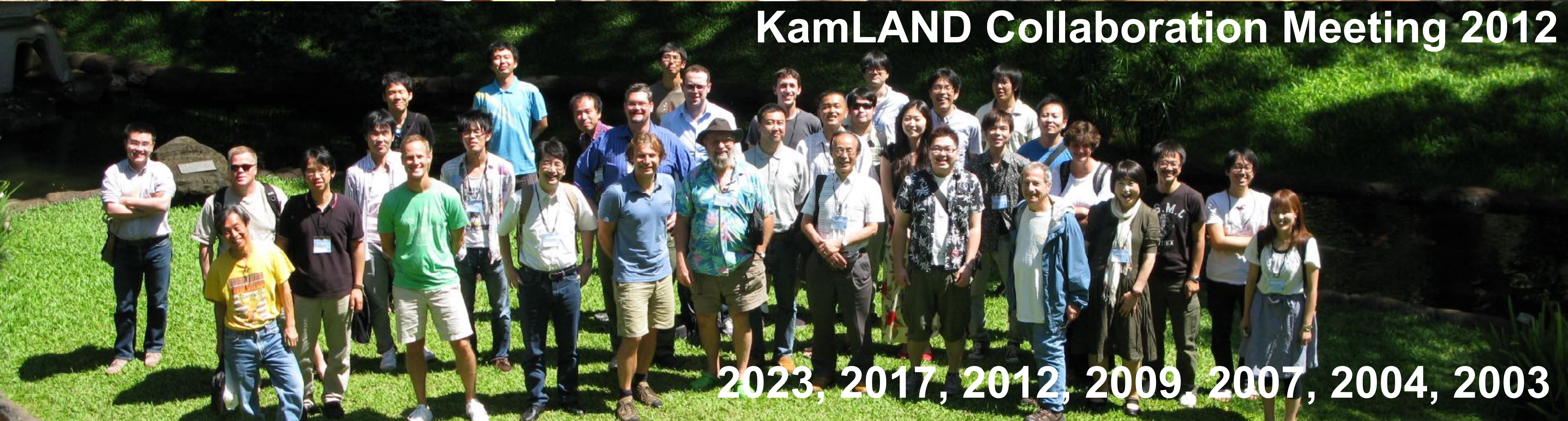
Chiba



Sendai



KamLAND Collaboration Meeting 2012



2023, 2017, 2012, 2009, 2007, 2004, 2003

① Crust : Solid

- Volume : 2 %, Mass : ~0.5 %

Heat-producing elements : 40 %

- Changes of seismic wave and chemical composition decide boundary

② Mantle : Solid

- Volume : 82 %, Mass : 68%

Heat-producing elements : 60 %

- Homogenous or Inhomogeneous? Chemical composition?

③ Outer Core : Fluid

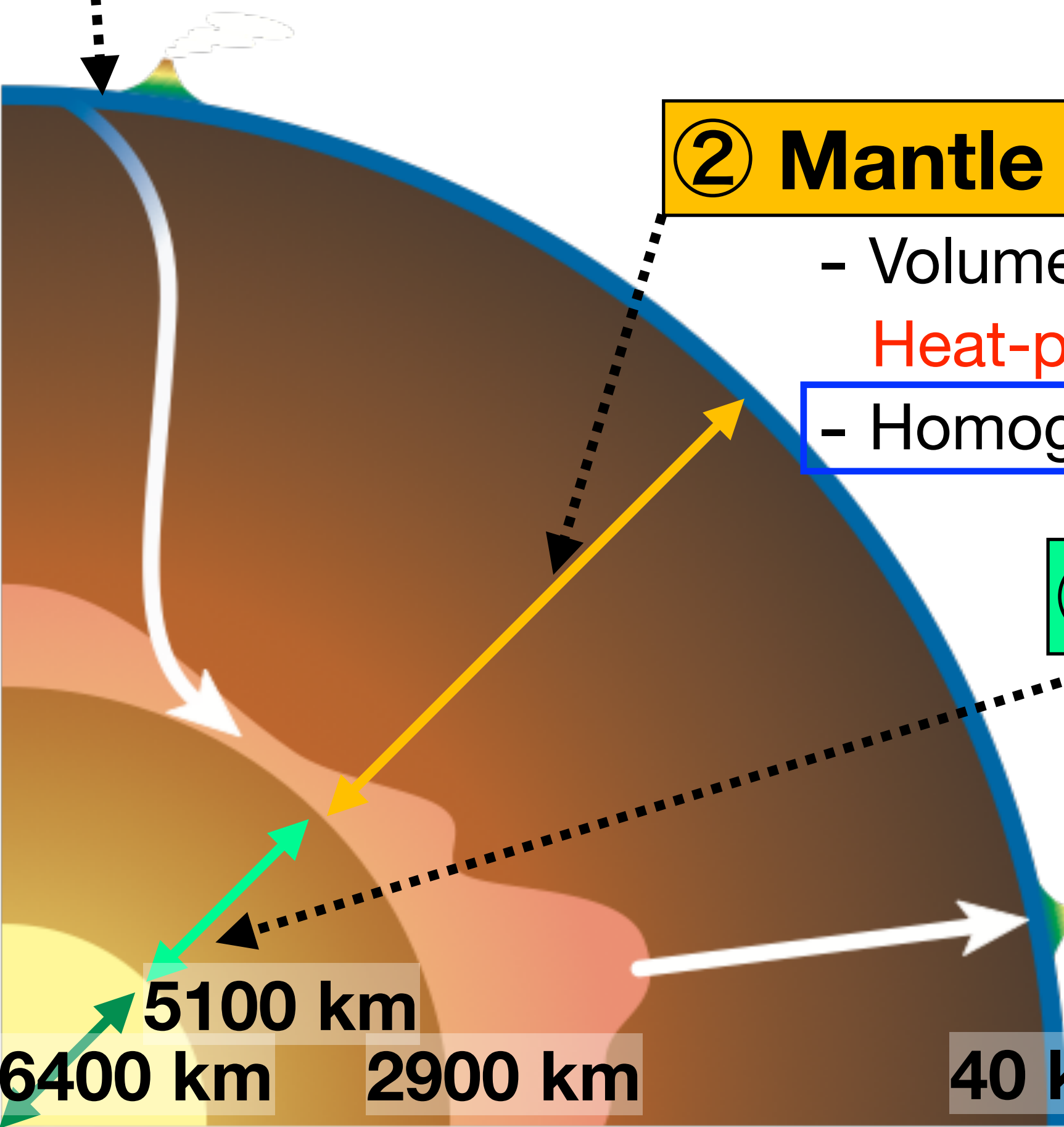
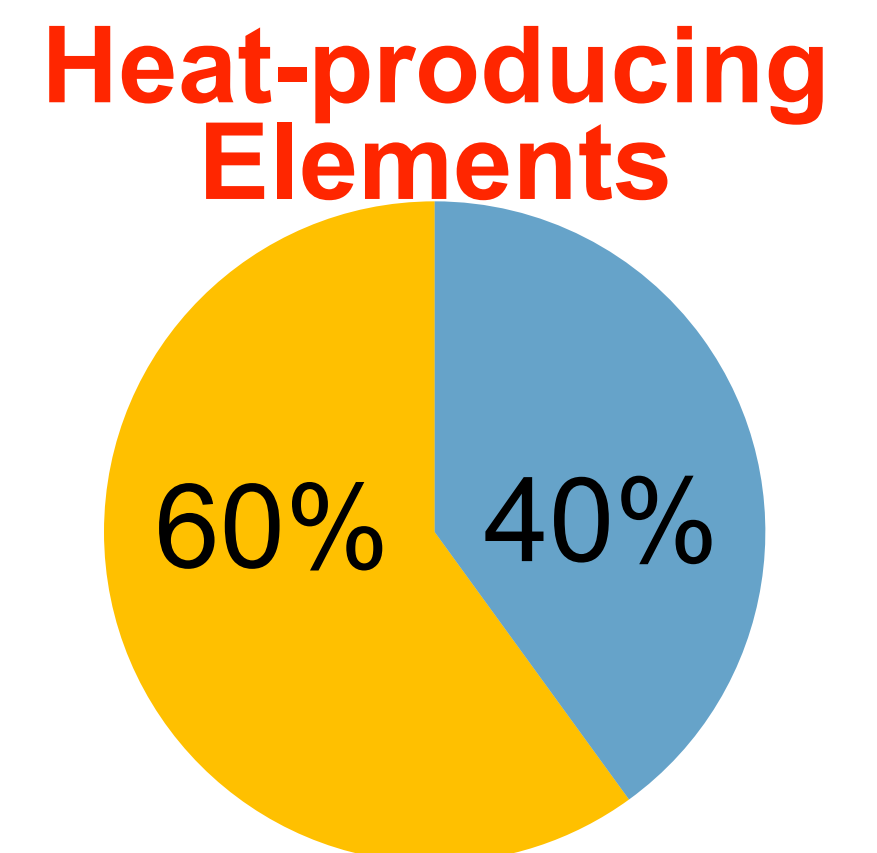
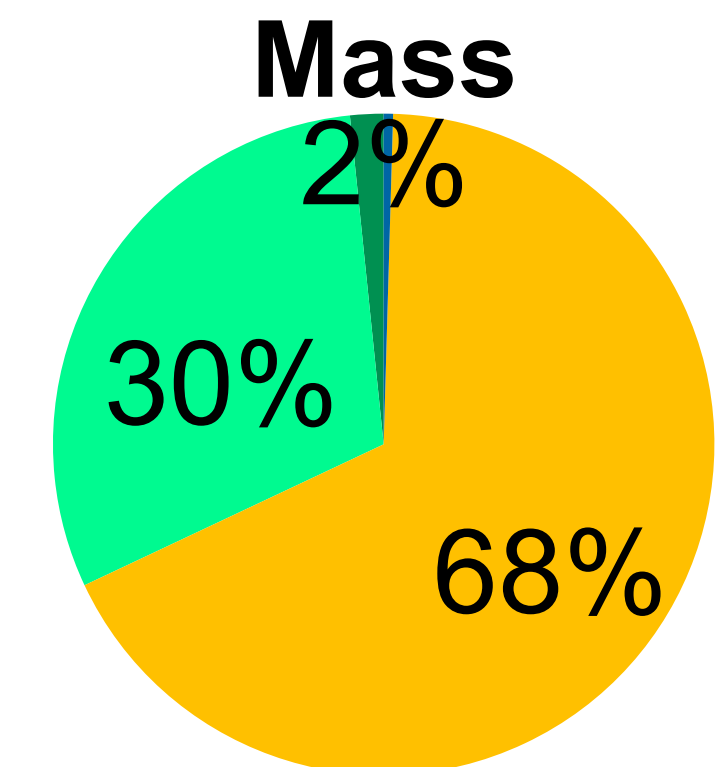
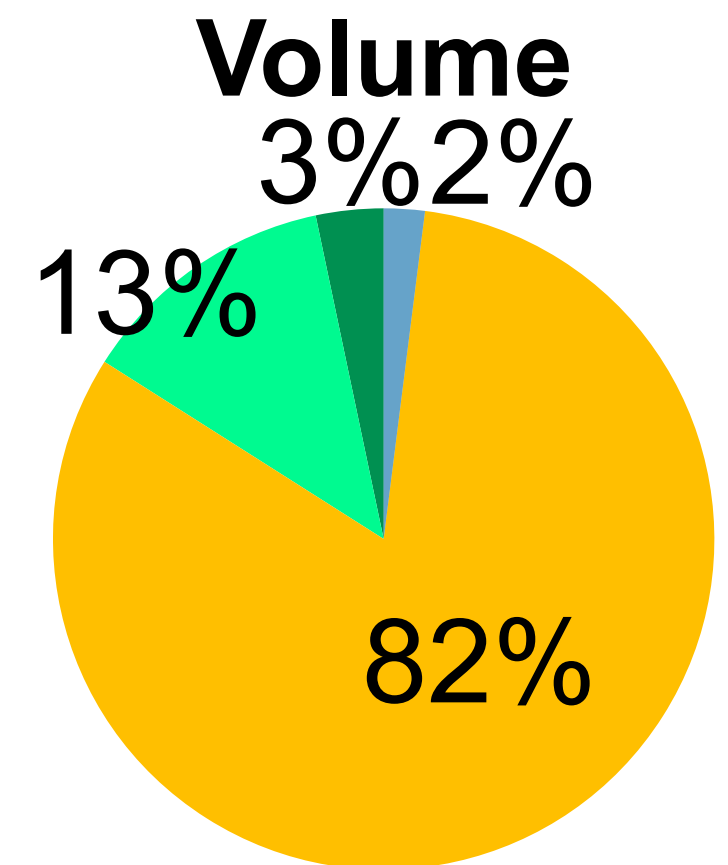
④ Inner Core : Solid

- Volume : 16%, Mass : 32%

Heat-producing elements : negligible

- Fe (+little Ni) Density : 10% lighter than Fe

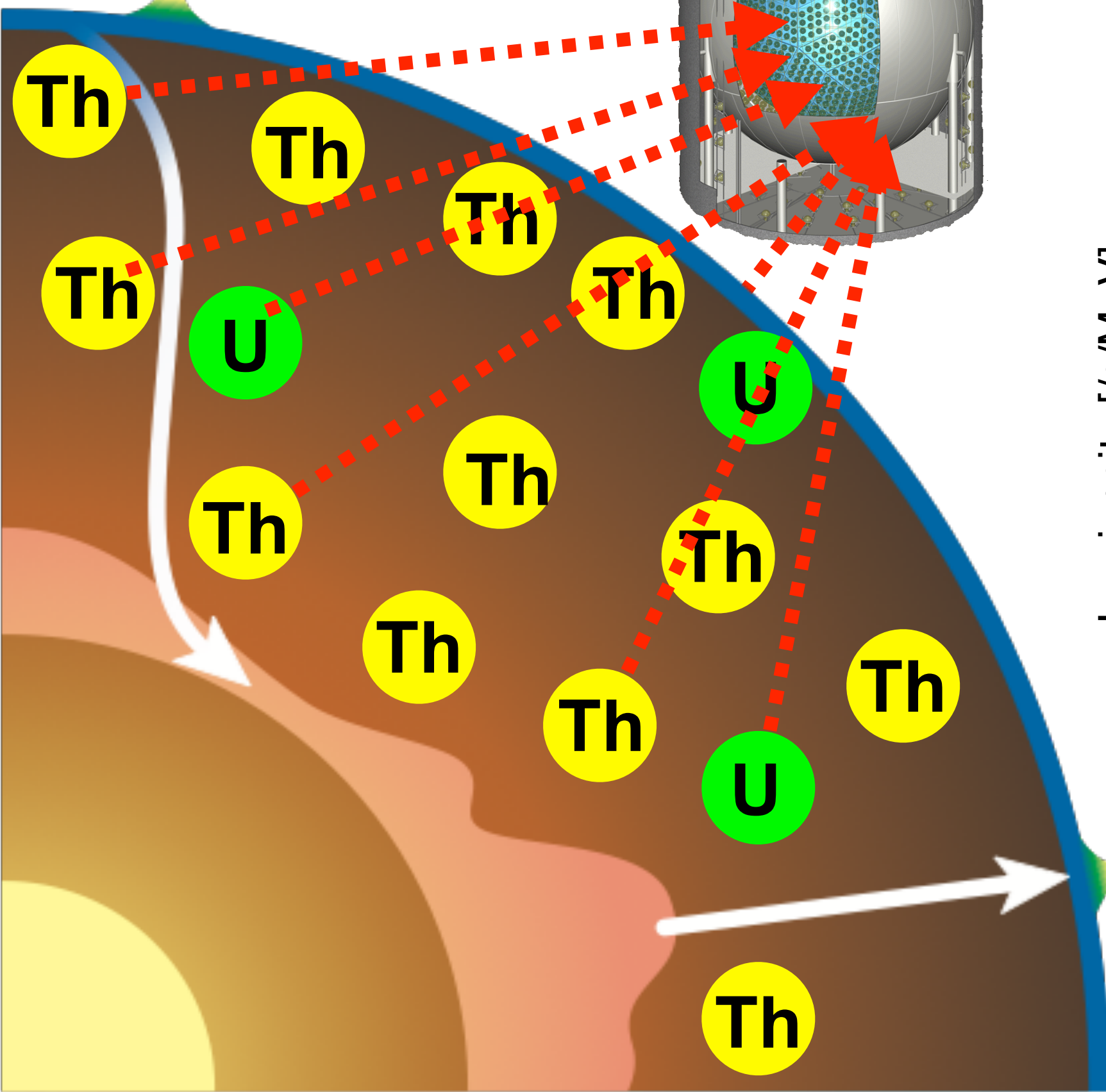
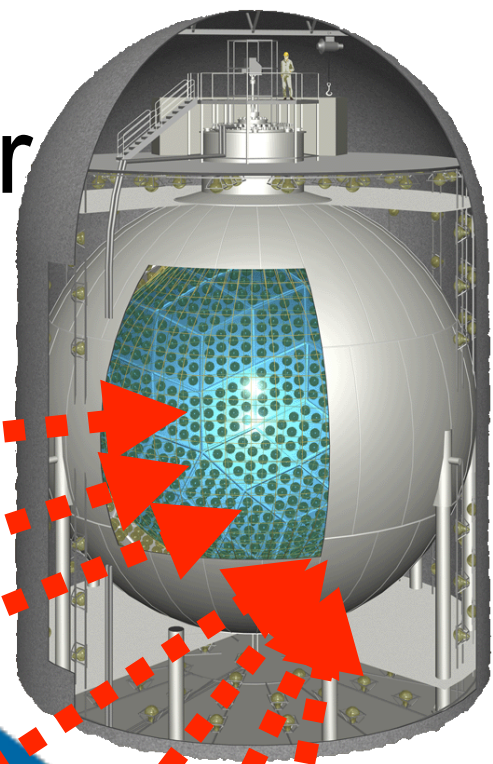
- Light elements? Amount?



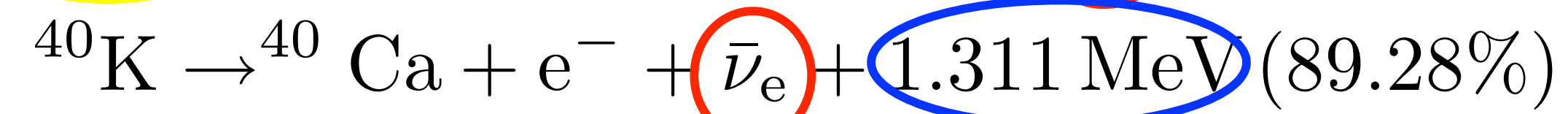
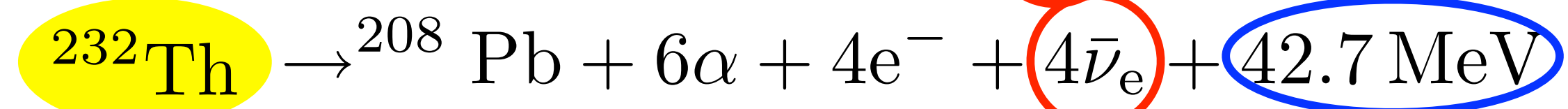
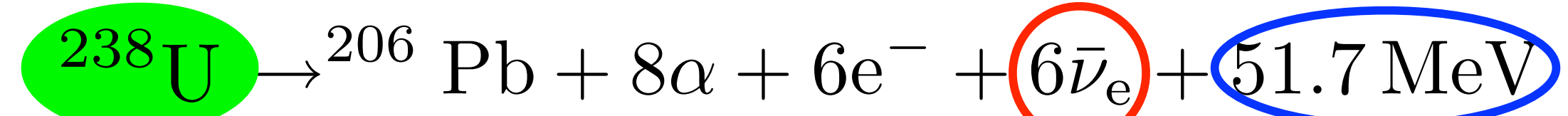
Electron-antineutrinos from natural radioactive decays

$$\bar{\nu}_e \ 4.1 \times 10^6 / \text{cm}^2 / \text{sec}$$

Anti-neutrino Detector
(e.g. KamLAND)

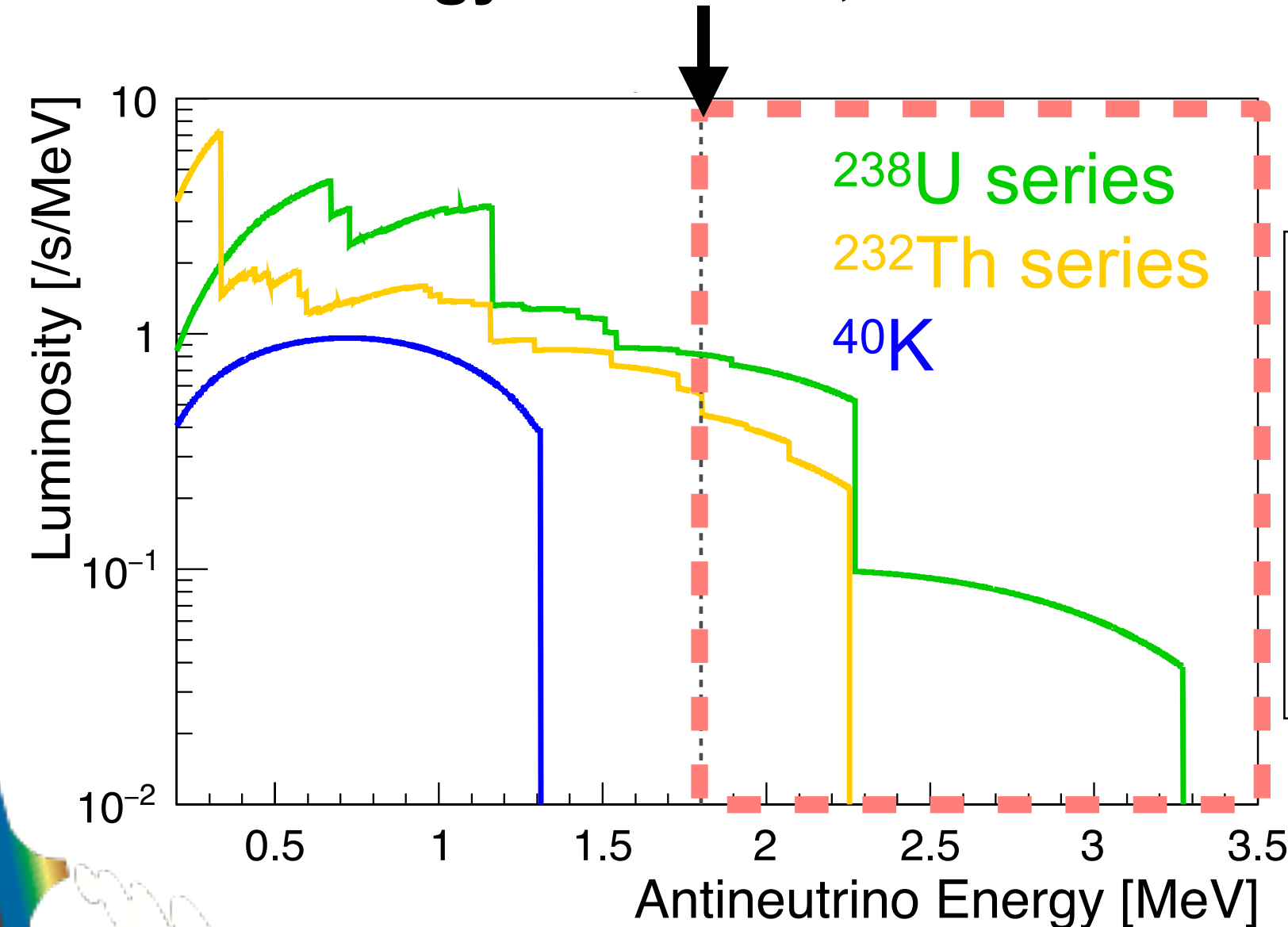


β -decay

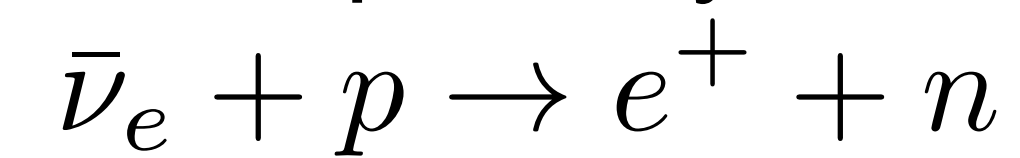


geo-neutrinos

Energy threshold, 1.8 MeV



inverse β -decay

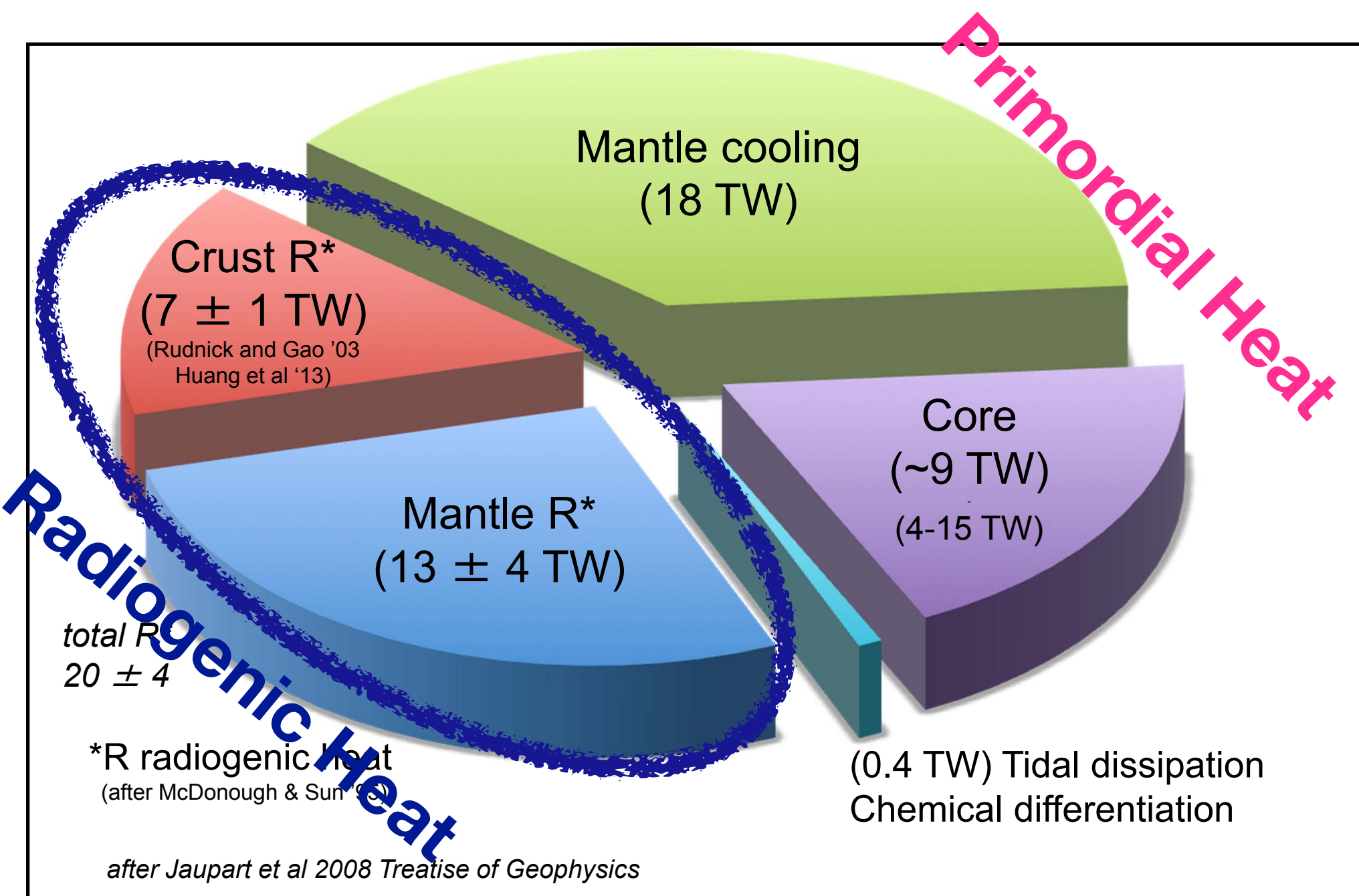


* Only geo-neutrinos from **U** and **Th** are detectable right now
* ^{40}K geo-neutrino detection needs another technology.

Number of geo $\bar{\nu}_e \propto$ amount of **U** **Th**, **radiogenic heat**

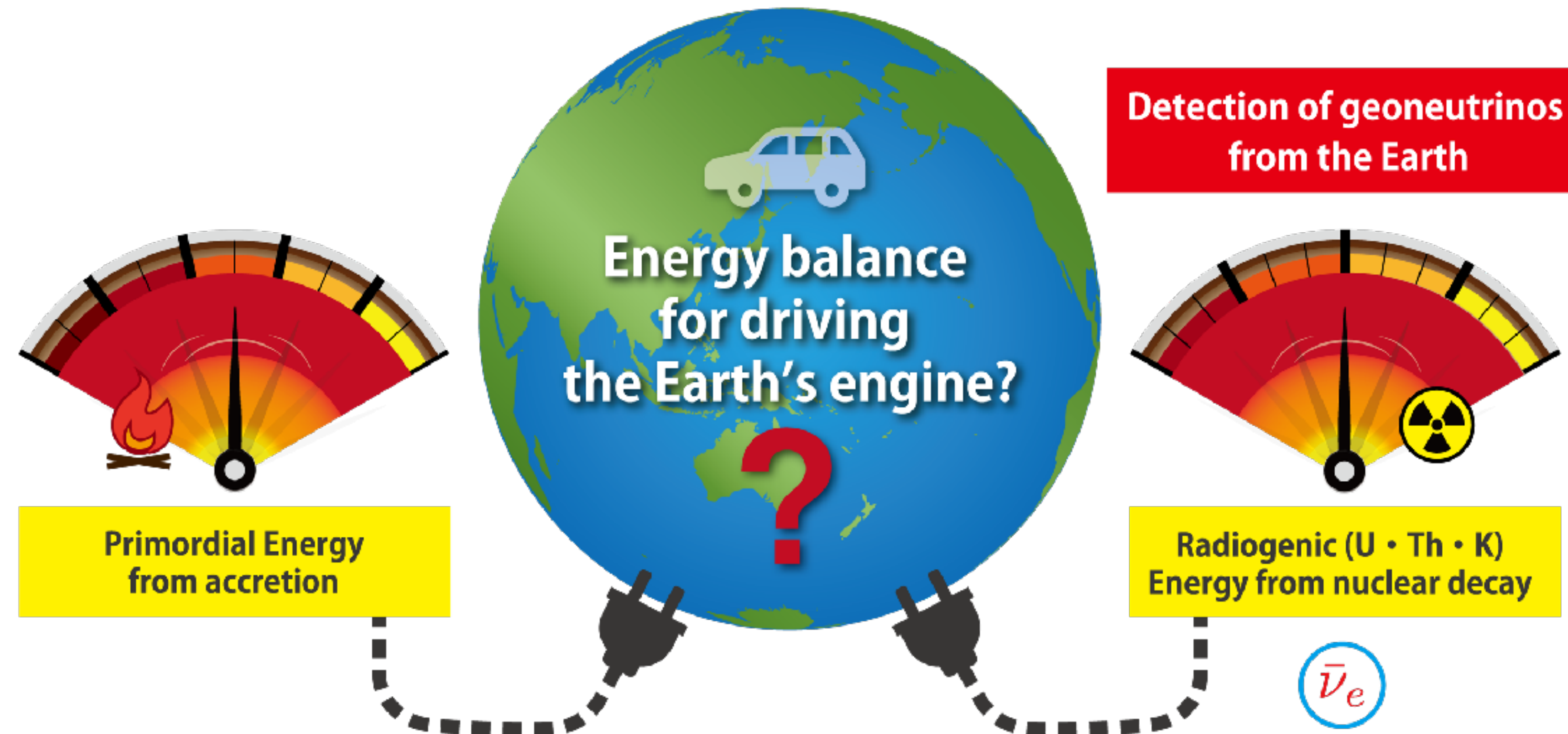
✓ Surface heat flow
 46 ± 3 TW

example of Earth model



Primordial Heat

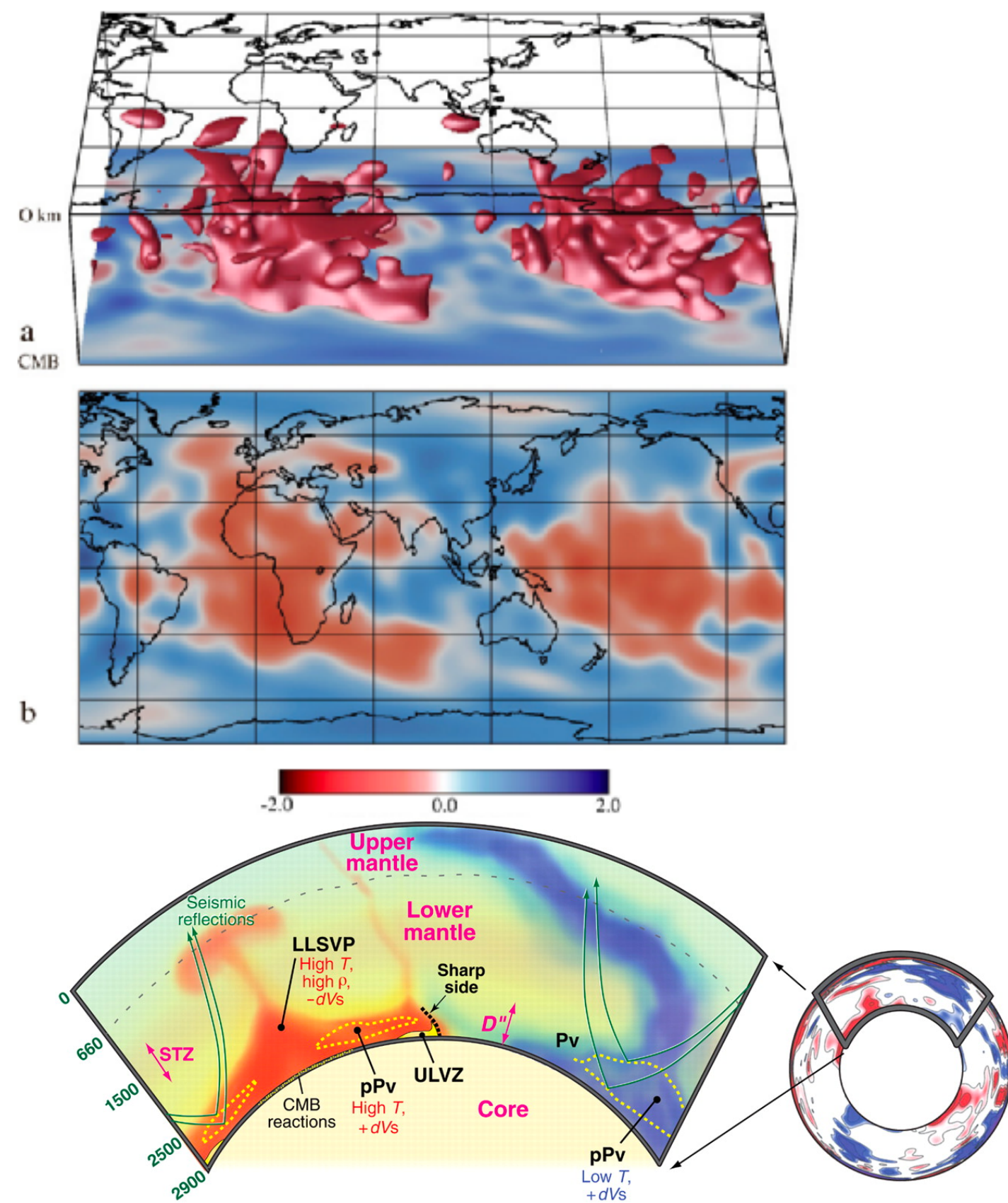
- * Releases of gravitational energy through accretion or metallic core separation
- * Latent heat from the growth of inner core



Q : How much radiogenic heat contributes to Earth's heat?

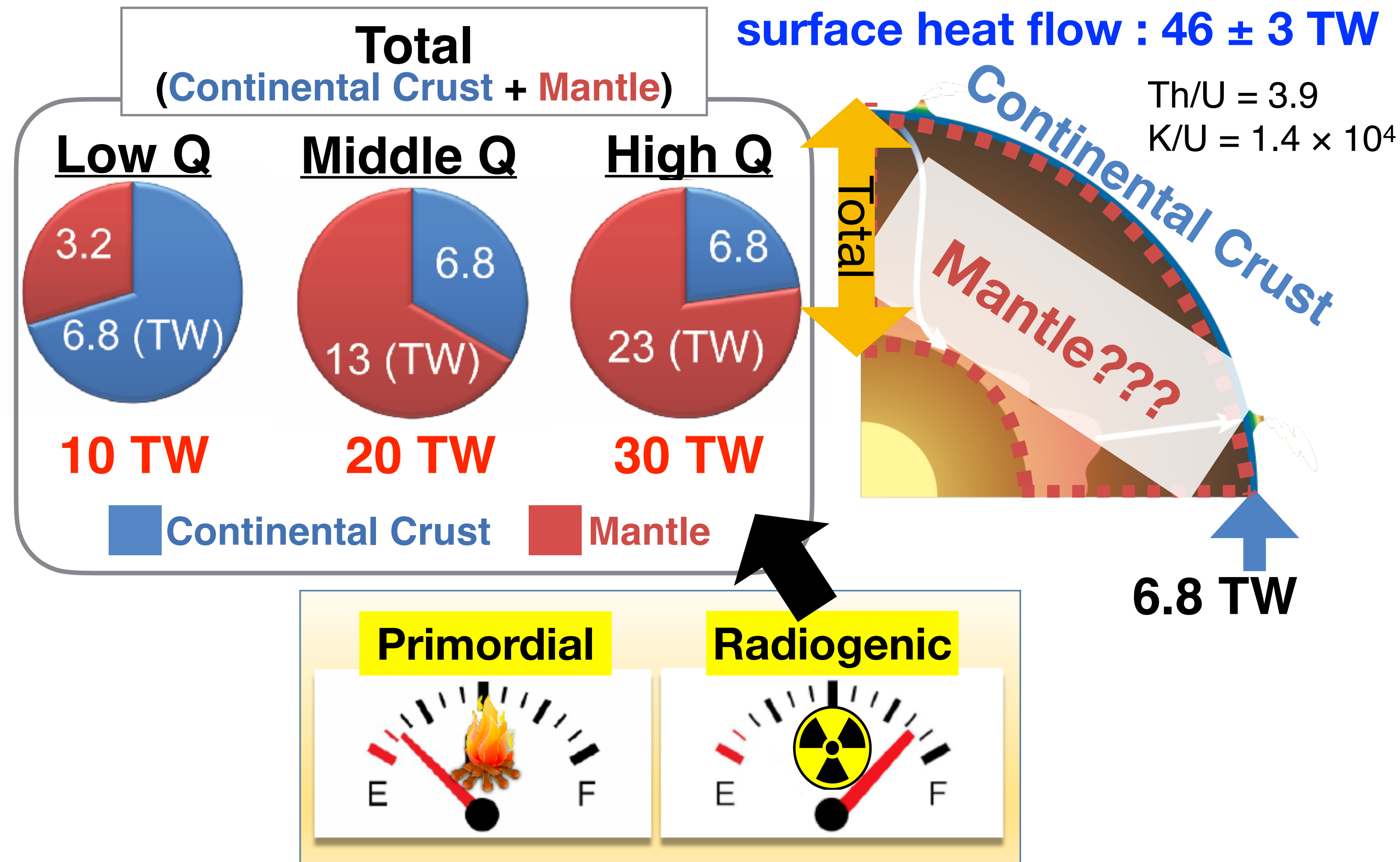
What is in the mantle?

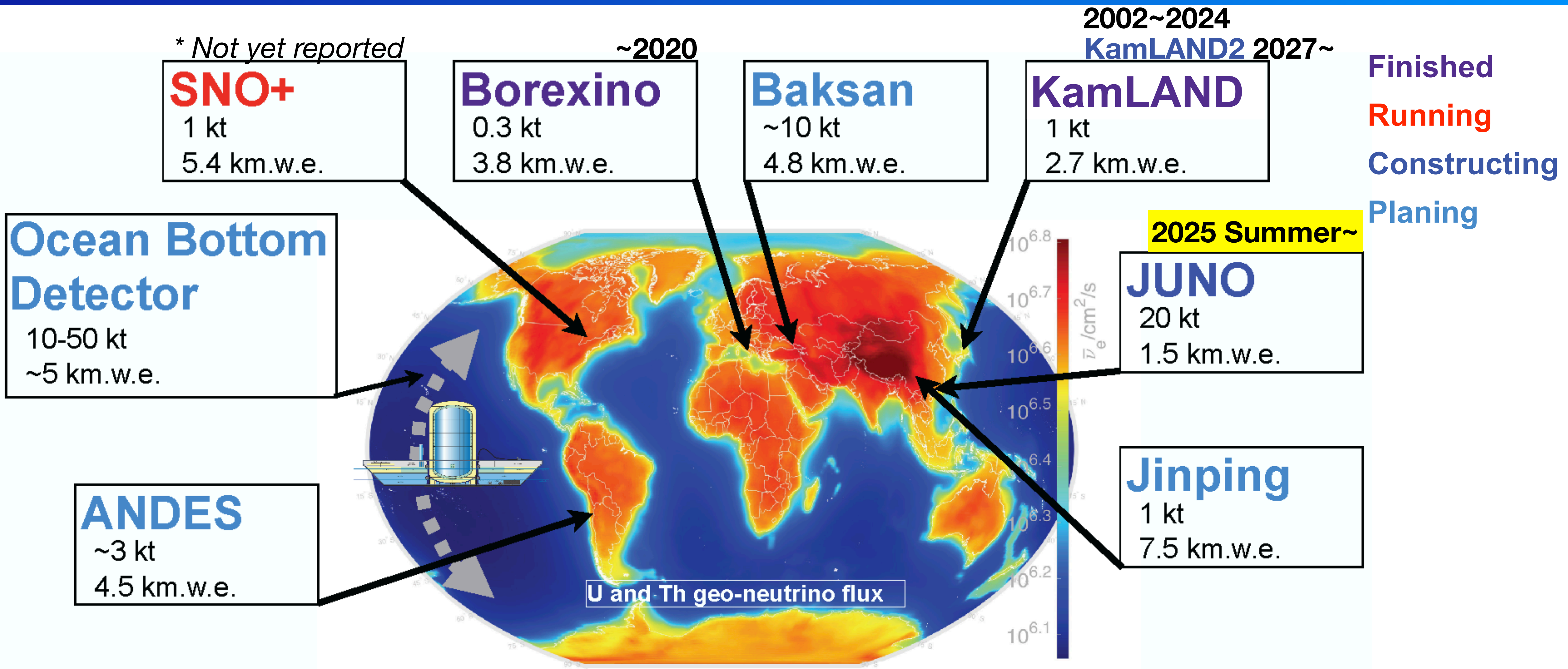
Many seismically imaged structures and chemical heterogeneities in the mantle



LLSVP (Large Low Shear Velocity Provinces)

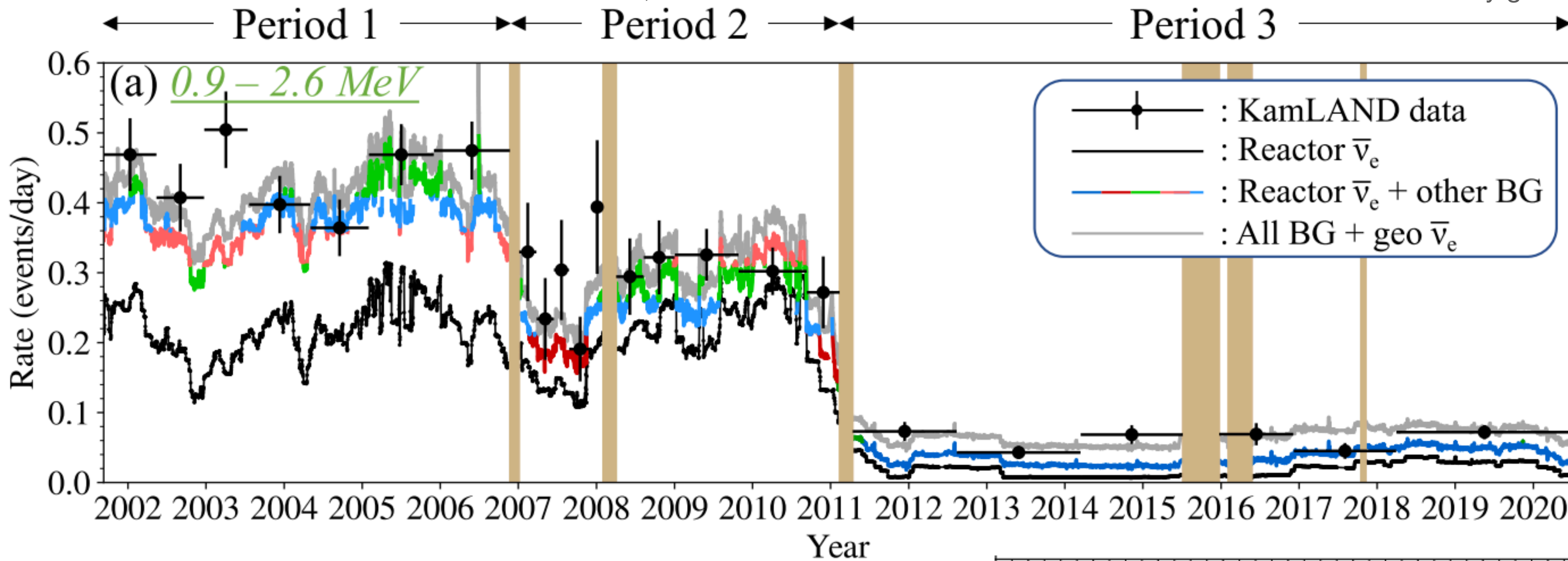
How much fuel is left to drive Plate Tectonics?





- Two experiments have published geoneutrino measurement results so far.
- New experimental data by **SNO+** & **JUNO** will be reported soon!

S. Abe et al, "Abundances of uranium and thorium elements in Earth estimated by geoneutrino spectroscopy", GRL, 49, e2022GL099566

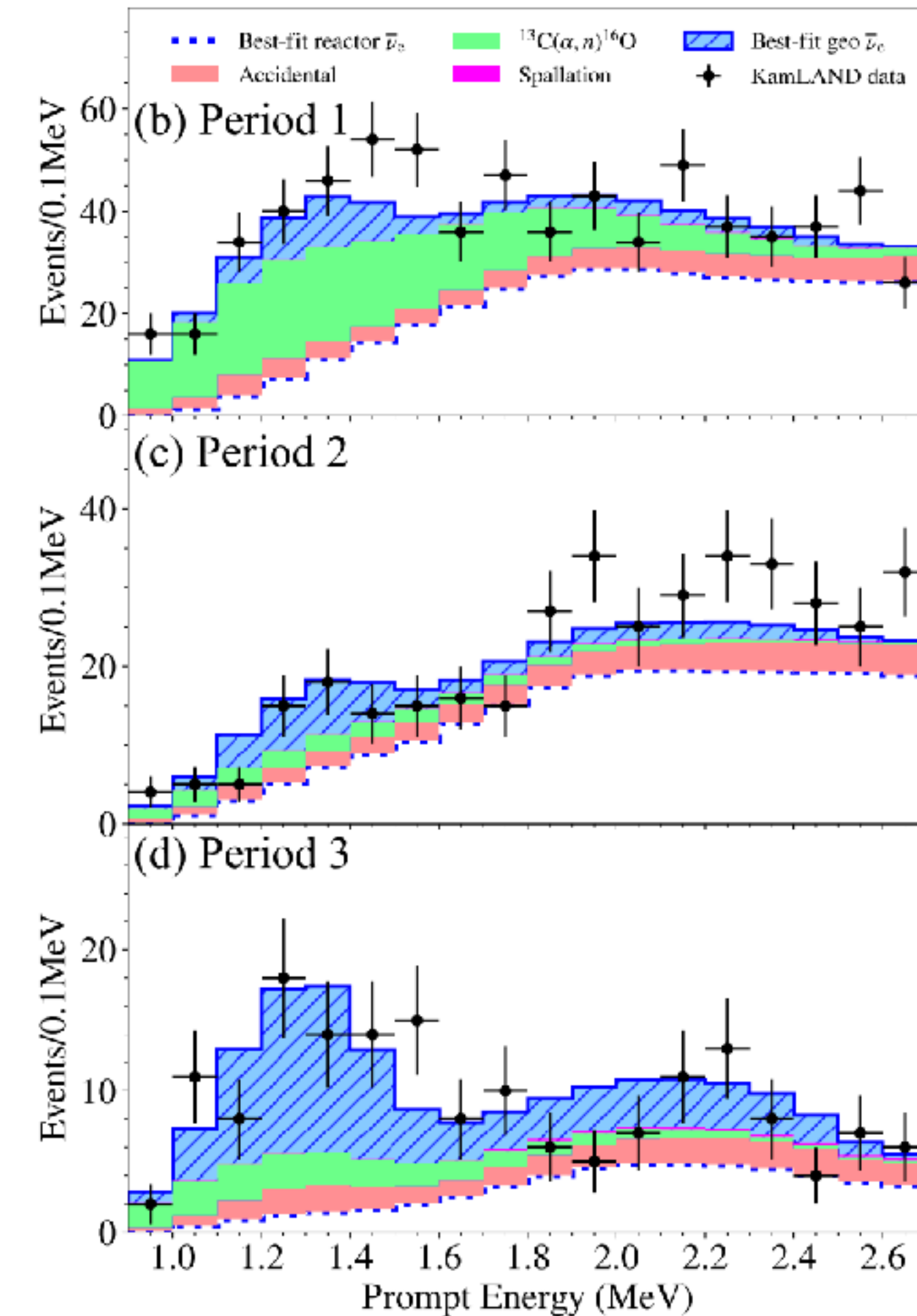
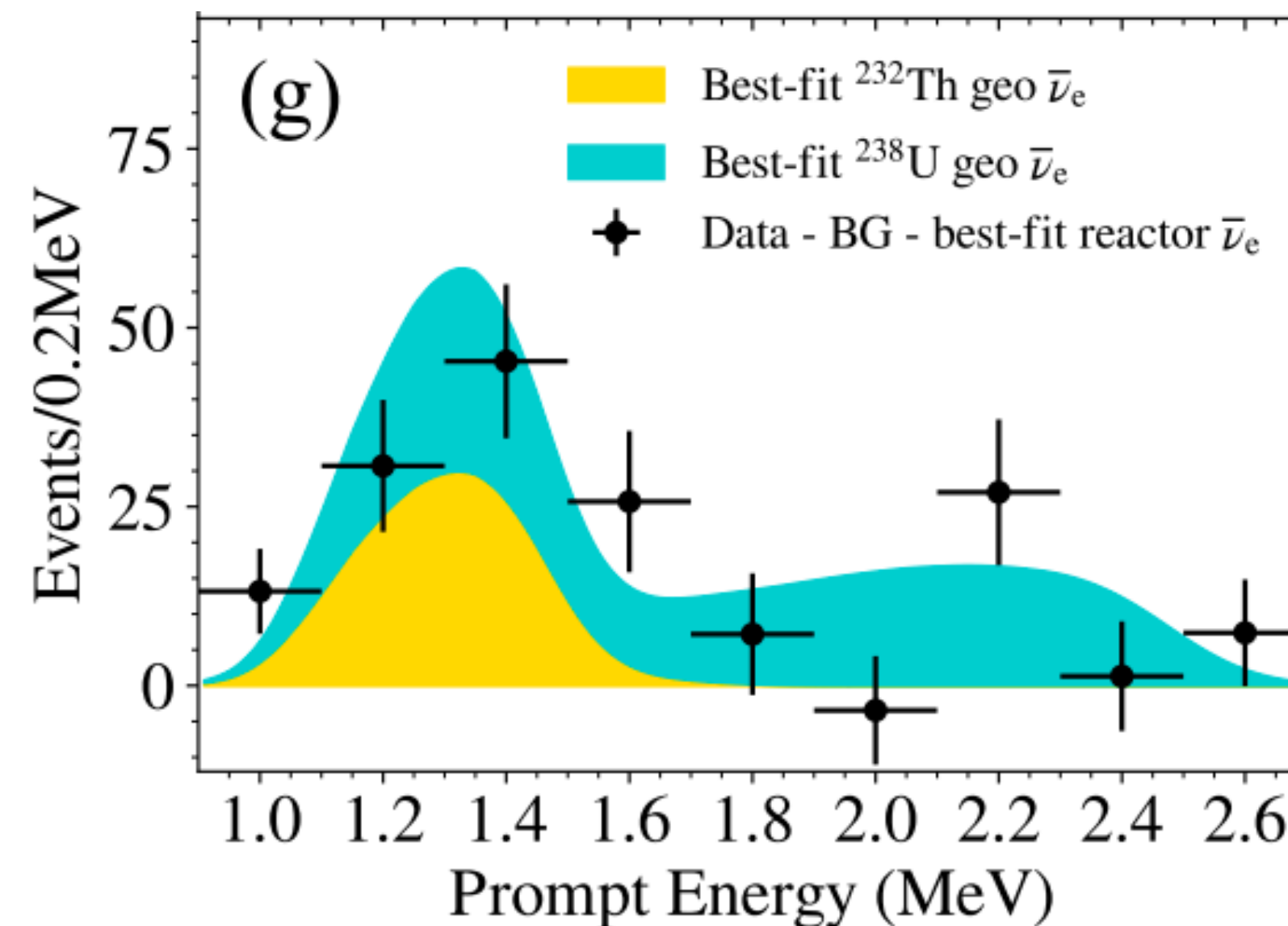


Dataset : Mar, 2002-Dec, 2021

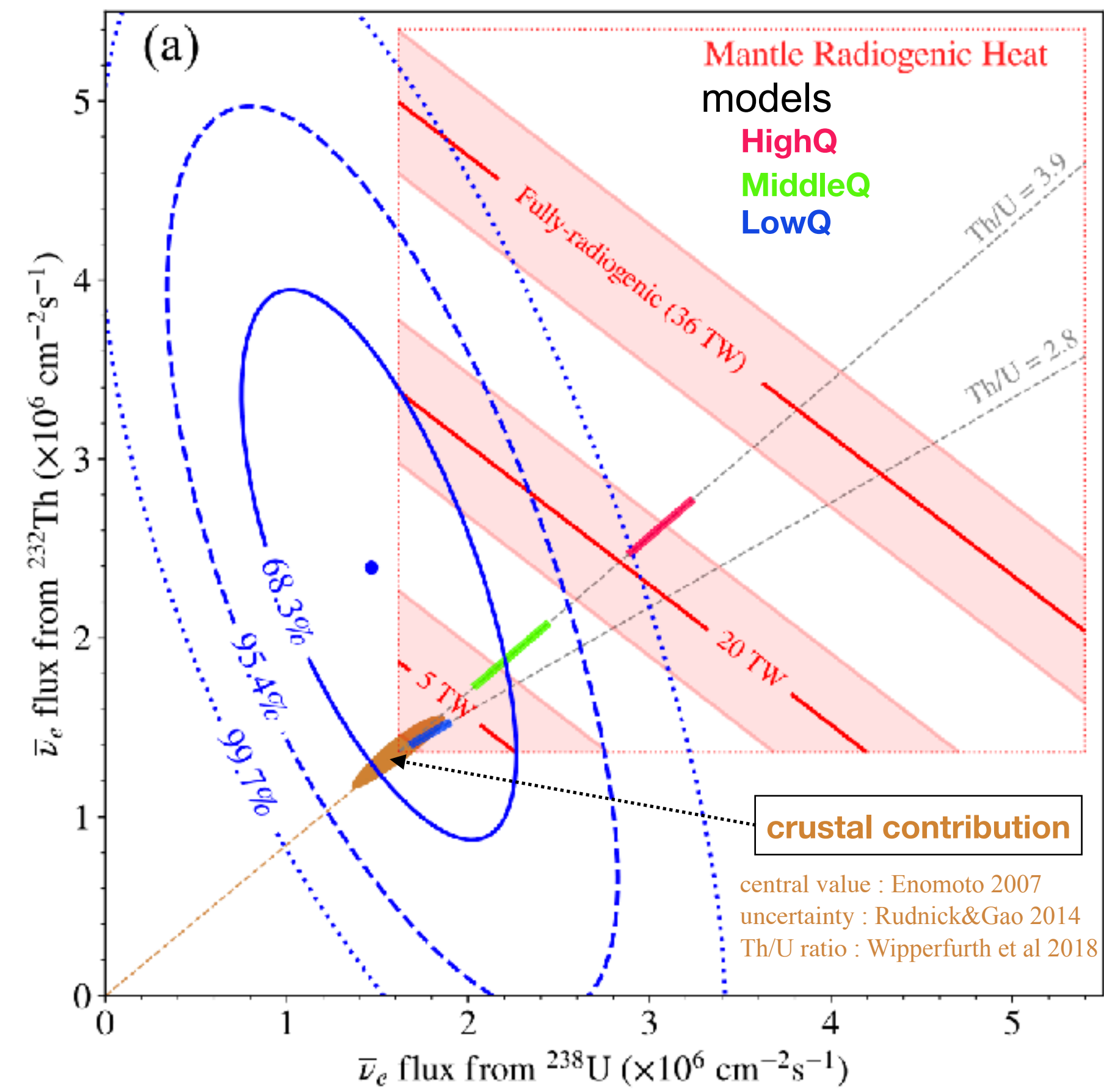
Livetime : 5227 days

(low-reactor phase : 2590 days)

✓ massive dataset of low-reactor period
→ precise measurement of U and Th contributions

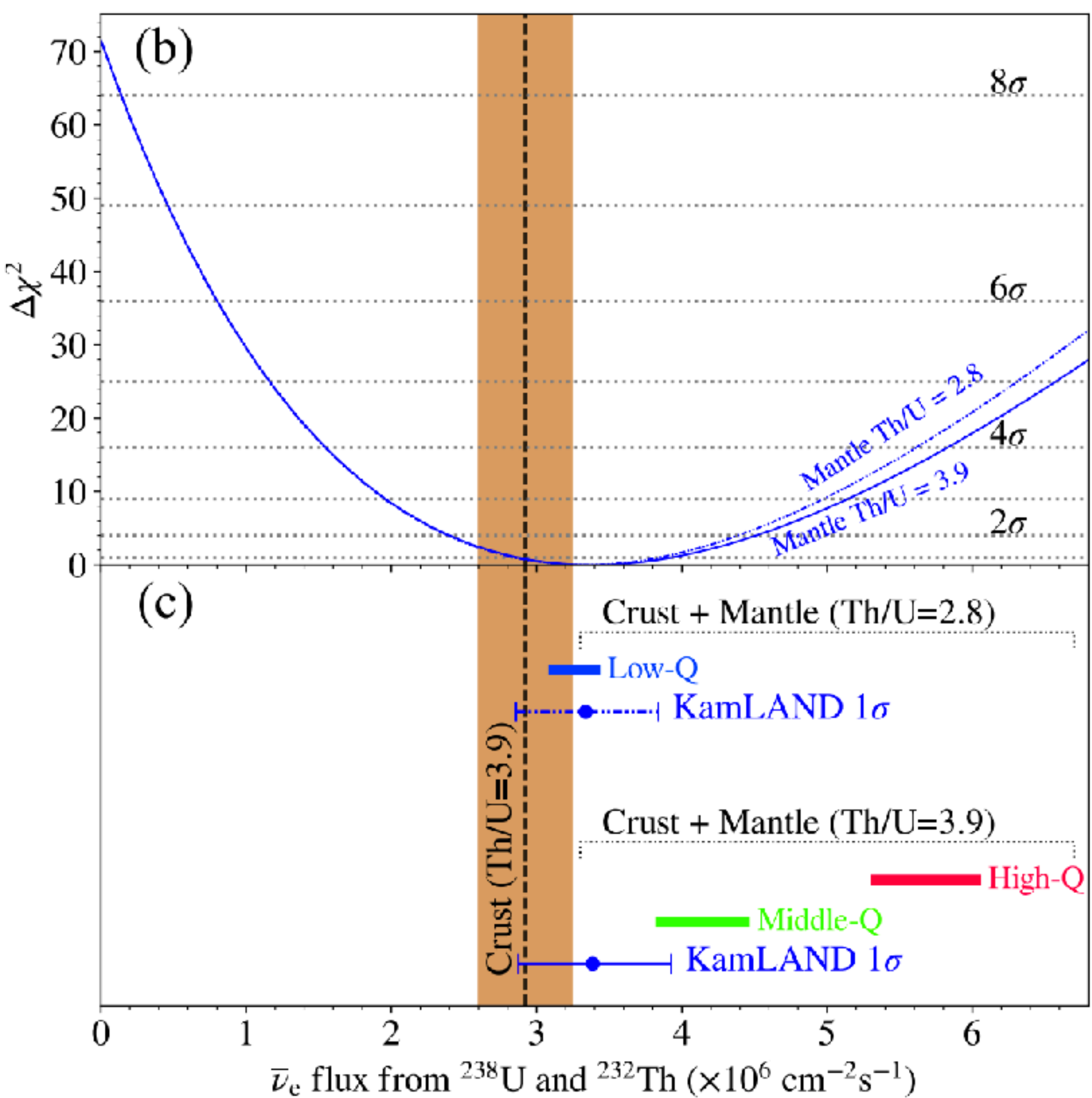


S. Abe et al, “Abundances of uranium and thorium elements in Earth estimated by geoneutrino spectroscopy”, GRL, 49, e2022GL099566



☒ **best-fit**
Th/U free

	N of event	0signal rejection
U	117 ⁺⁴¹ ₋₃₉	3.3σ
Th	58 ⁺²⁵ ₋₂₄	2.4σ
U+Th	174 ⁺²⁹ ₋₂₈	8.3σ



☒ **Radiogenic Heat**
Th/U free

Adding heat estimate from crust,
²³⁸U : 3.4 TW, ²³²Th : 3.6 TW

$$Q^U = 3.3^{+3.2}_{-0.8} \text{ TW}$$
$$Q^{Th} = 12.1^{+8.3}_{-8.6} \text{ TW}$$
$$Q^U + Q^{Th} = 15.4^{+8.3}_{-7.9} \text{ TW}$$

☒ **Model Rejection**
HighQ model is rejected at
99.76 % C.L. (homogeneous mantle)
97.9% C.L. (concentrated at CMB)

Achieved the accuracy level can further geoscientific discussion
Improve the distinct spectroscopic contributions of U and Th

Final Result of KamLAND1

10/23

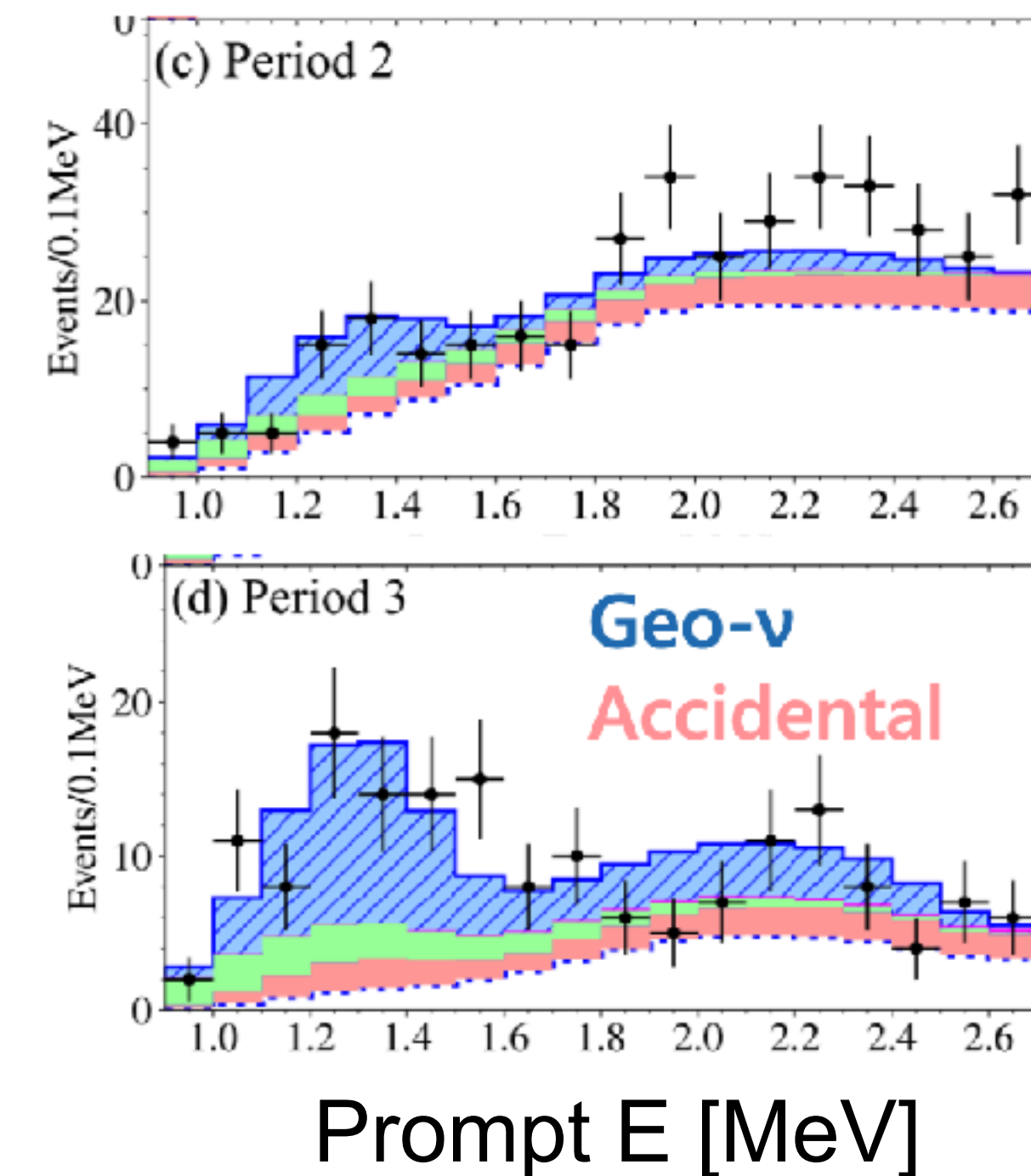
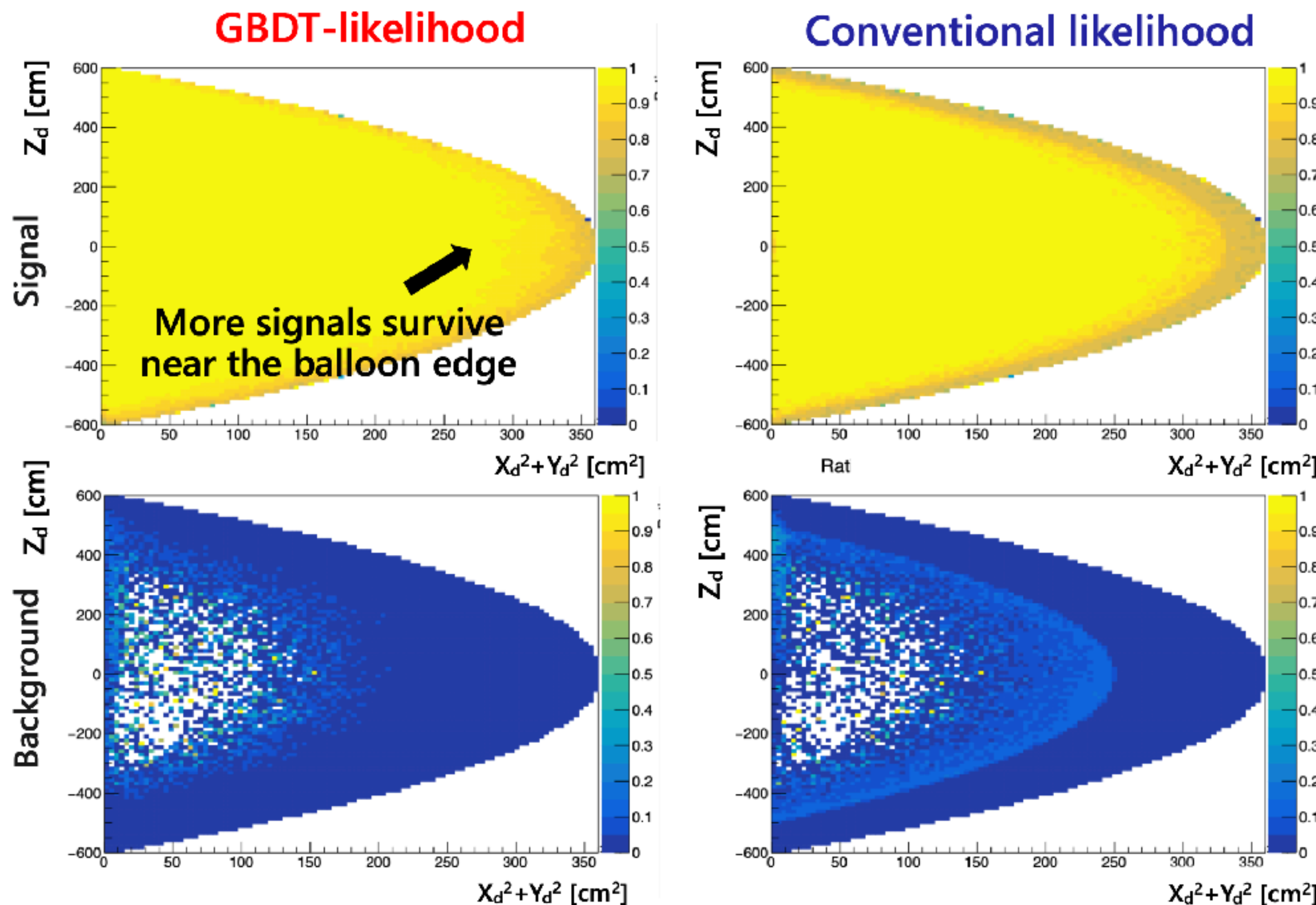
Dataset : Mar, 2002-Aug, 2024 (+2.5 years from last result)

+ **BG reduction with machine learning**

Gradient boosting decision tree is powerful algorithm for **accidental BG** reduction.

Distribution of
remaining ratio
in KamLAND

Z axis
Surviving events
All event

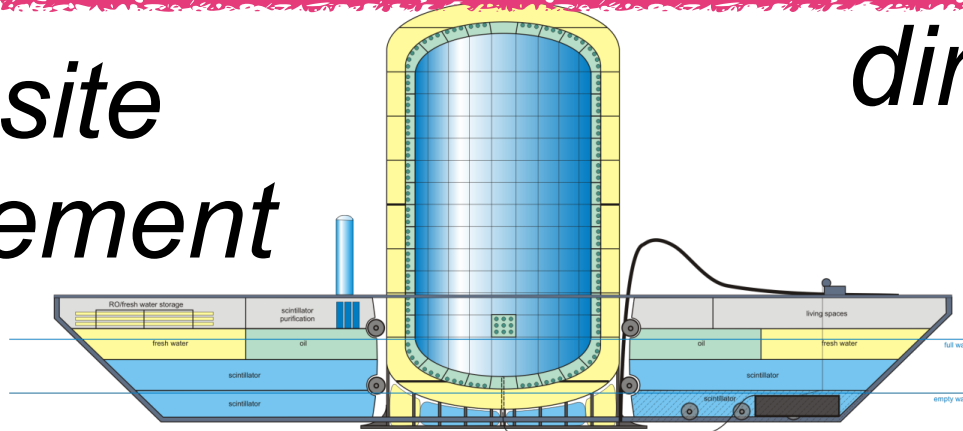


KamLAND full data analysis
with GBDT is ongoing !

what we need

improved accuracy of
measurement
&
modelling

multi-site
measurement



directional sensitive detector

new type detector

detector in the Ocean

current generation

next generation

total radiogenic heat
in the Earth

☒ Measuring

resolving vertical and
horizontal **flux differences**

Th/U ratio

☒ Measuring

Next Target!

distinguishing **mantle**
contribution

☒ Measuring

detecting
K geo-neutrino

what we learn



first measurement in **2005**

**OBD: breakthrough
beyond modern land-based detector
transforming our vision of deep Earth**

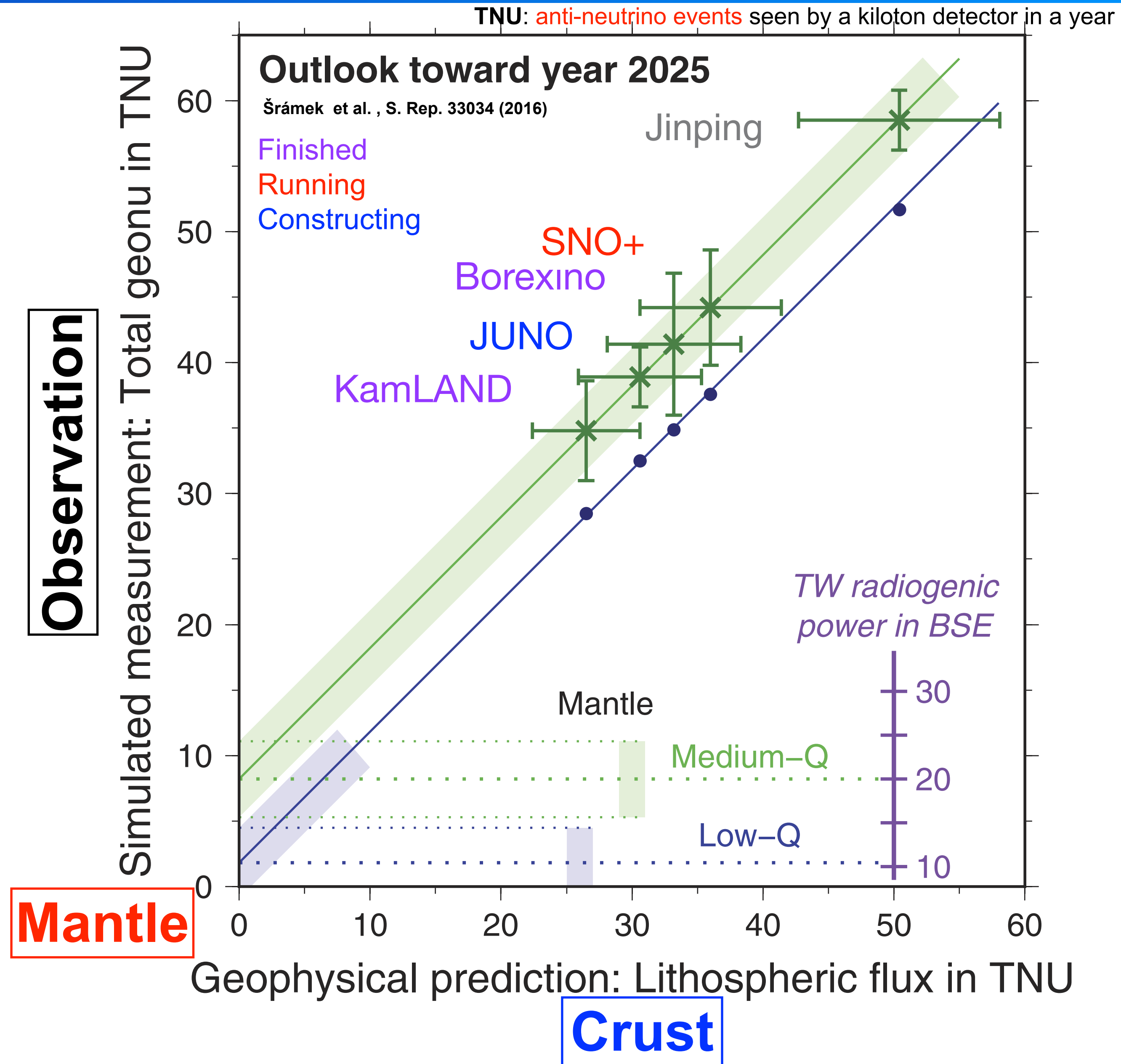
$$\text{Observation} = \text{Crust} + \text{Mantle}$$

$$(y = x + b)$$

Near Future...

4 multi-site measurements can constrain mantle contribution.

- * KamLAND, Borexino, SNO+, JUNO
- * Crust estimation needs to be accurate.



$$\text{Observation} = \text{Crust} + \text{Mantle}$$

$$(y = x + b)$$

Near Future...

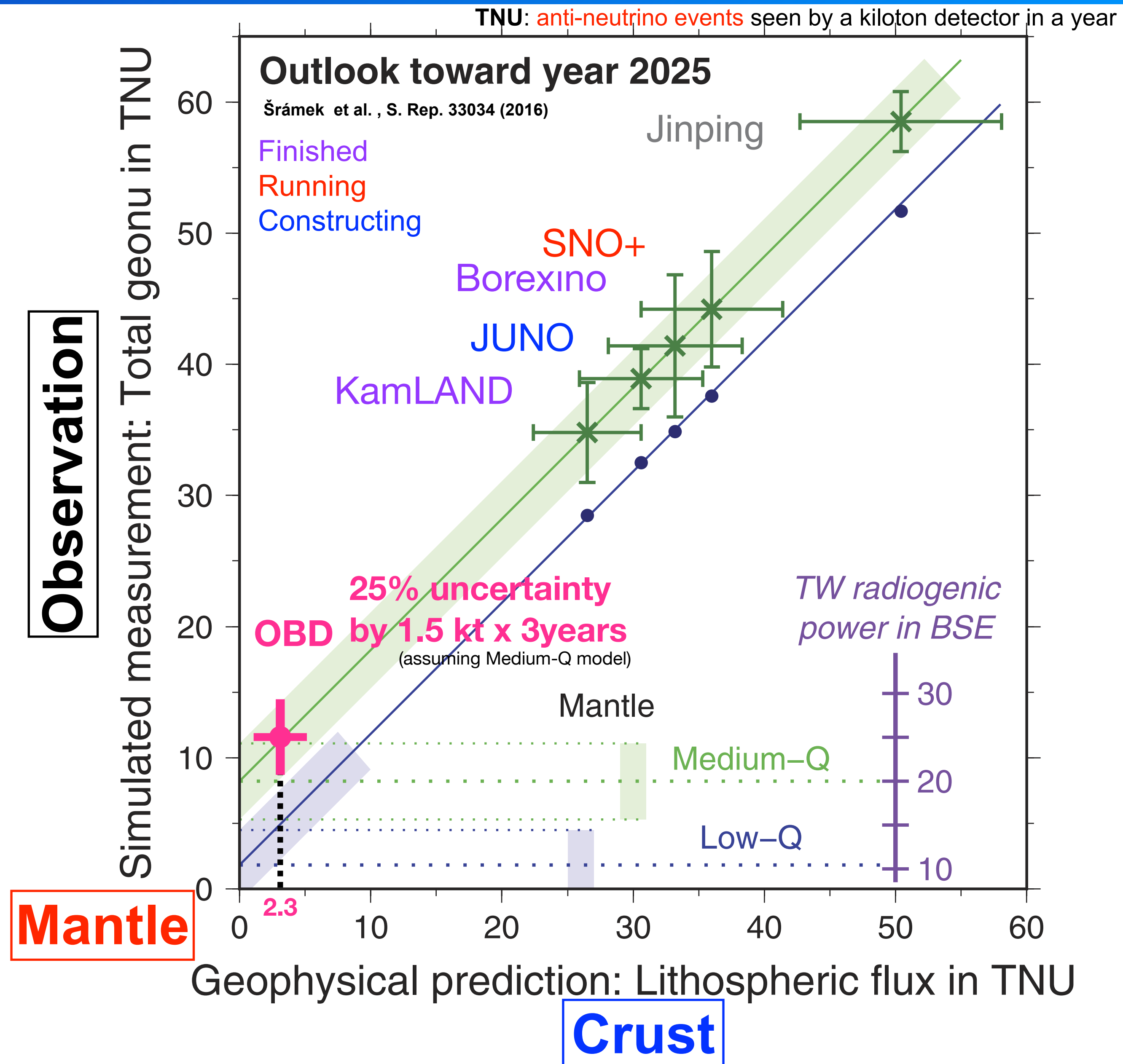
4 multi-site measurements can constrain mantle contribution.

* KamLAND, Borexino, SNO+, JUNO

* Crust estimation needs to be accurate.

+ Ocean Bottom Detector (Hanohano)

directly measure mantle contribution.



• Direct Measurement of Mantle

need to be far from crust
can be far from reactors

• Multi-site Measurements

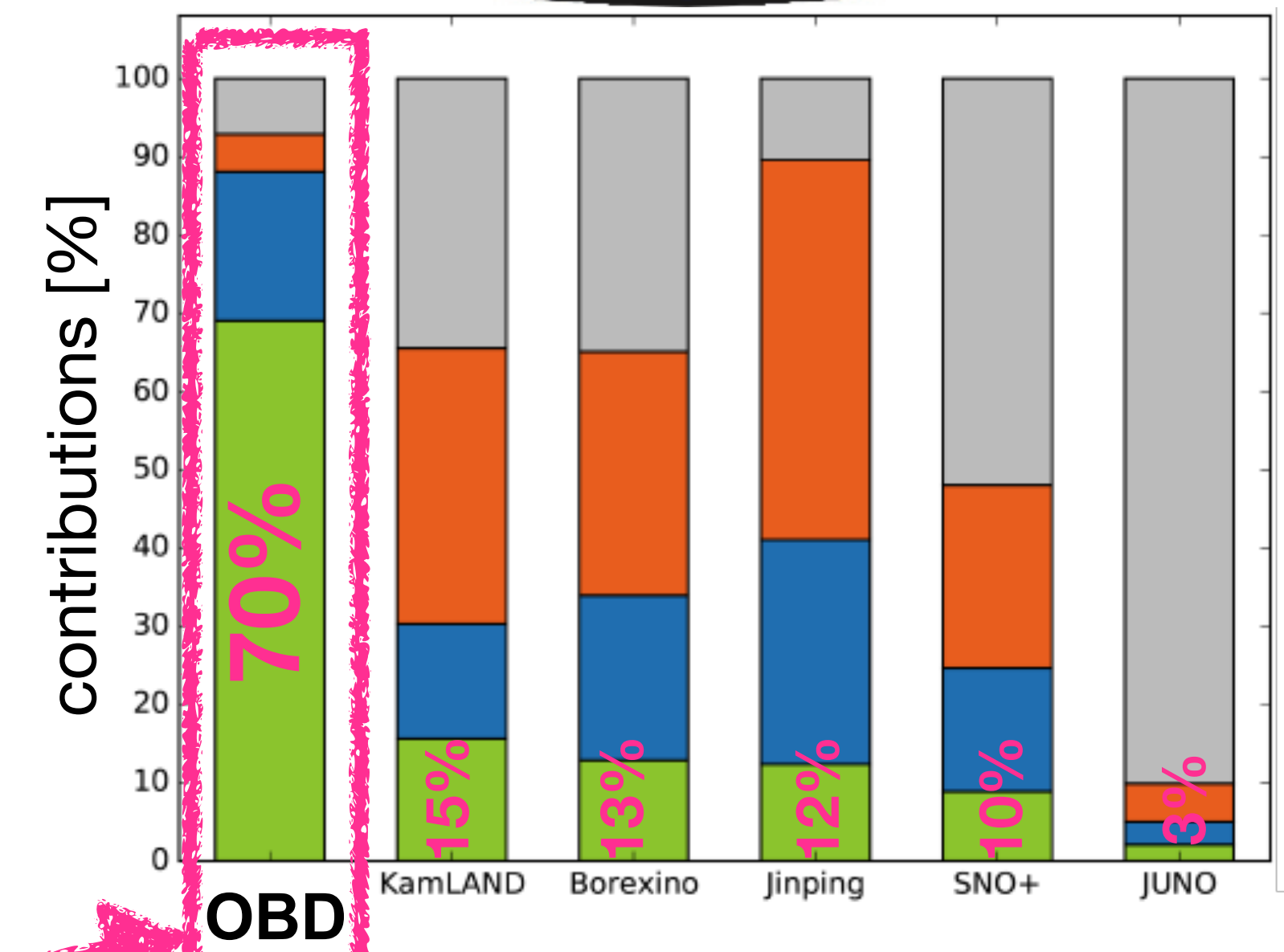
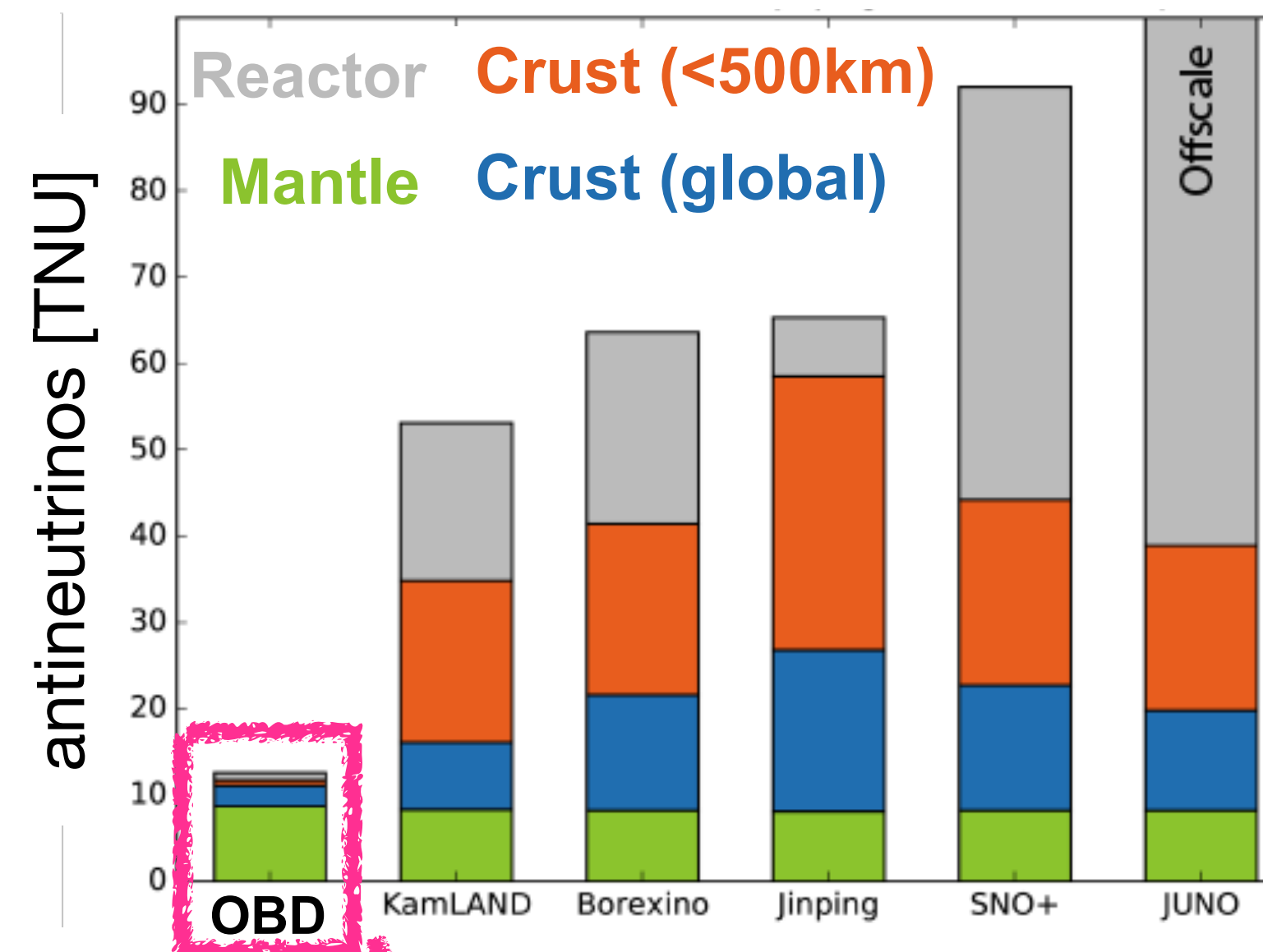
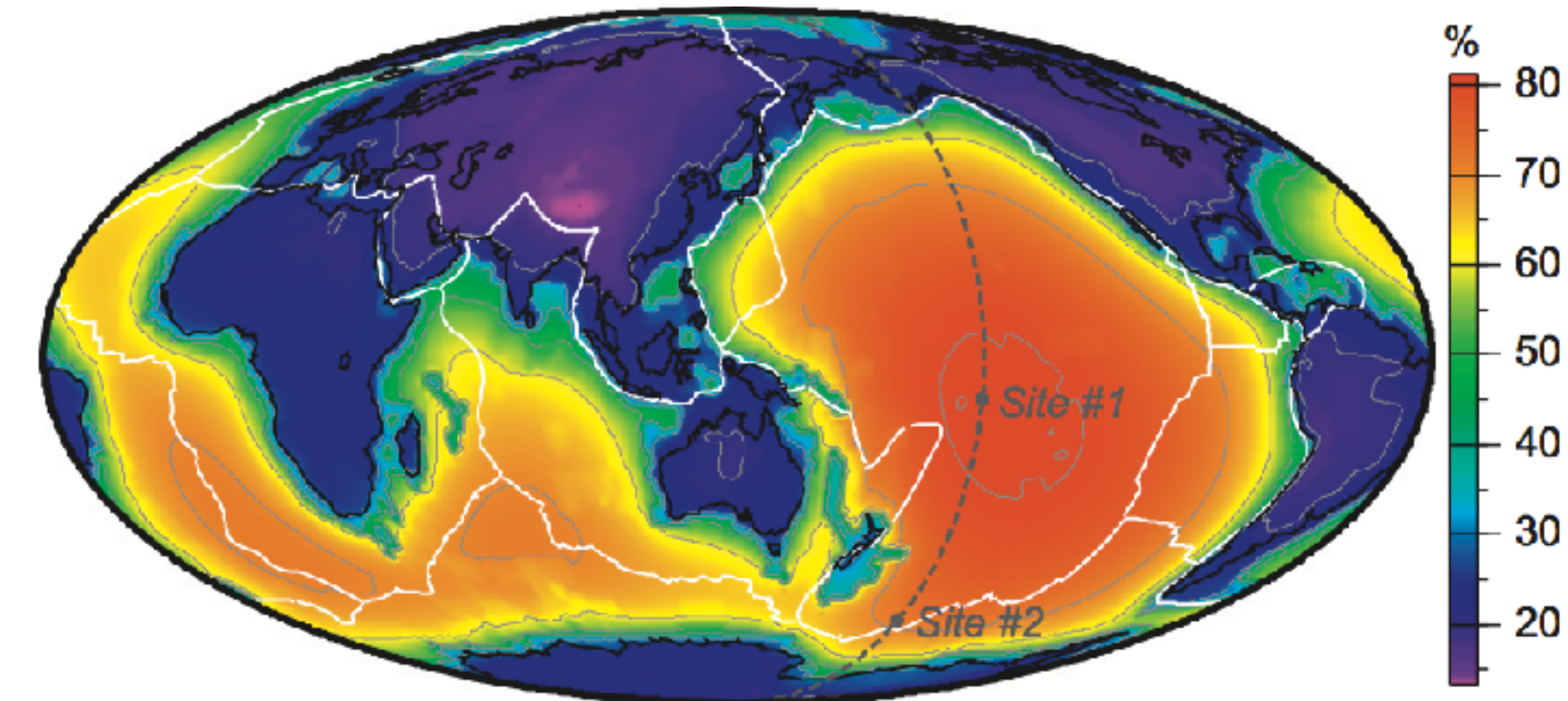
Solve the mystery of deep Earth!

First detector for mapping **the inhomogeneous mantle**

• Multidisciplinary Detector

Šrámek et al (2013) EPS, [10.1016/j.epsl.2012.11.001](https://doi.org/10.1016/j.epsl.2012.11.001)

Mantle/Total



• Direct Measurement of Mantle

need to be far from crust
can be far from reactors

• Multi-site Measurements

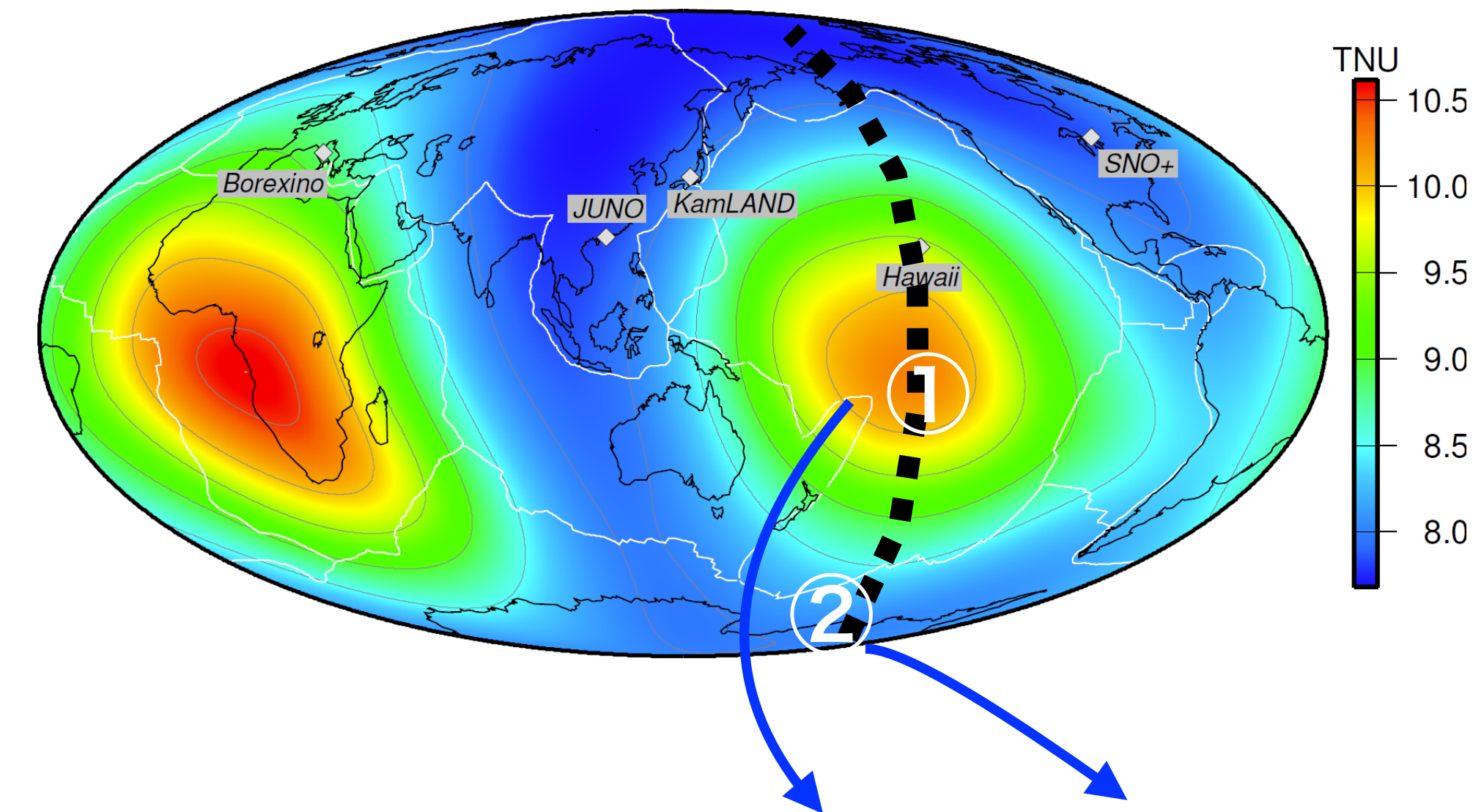
Solve the mystery of deep Earth!

First detector for mapping **the inhomogeneous mantle**

• Multidisciplinary Detector

Šrámek et al (2013) EPS, 10.1016/j.epsl.2012.11.001

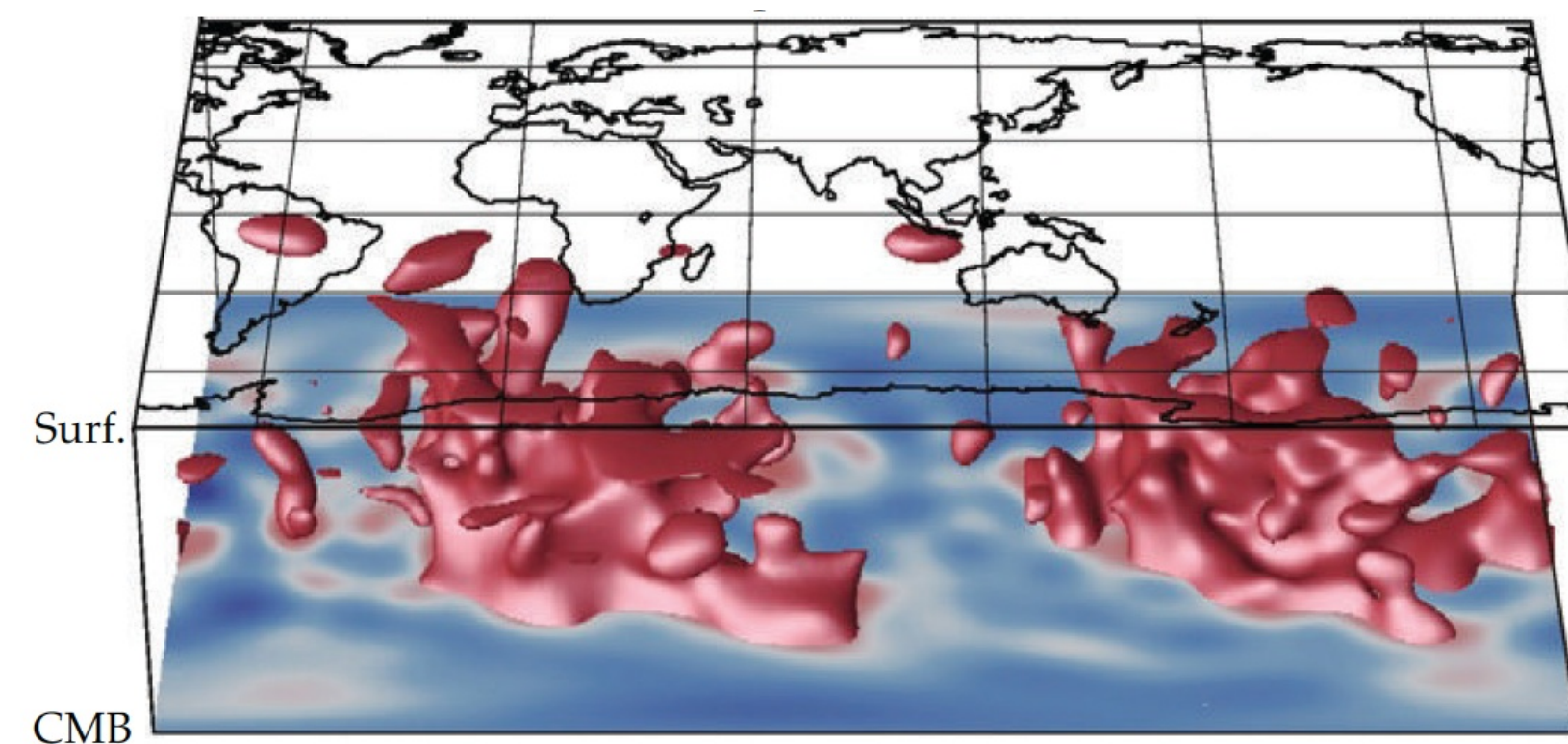
Mantle Geoneutrino Flux



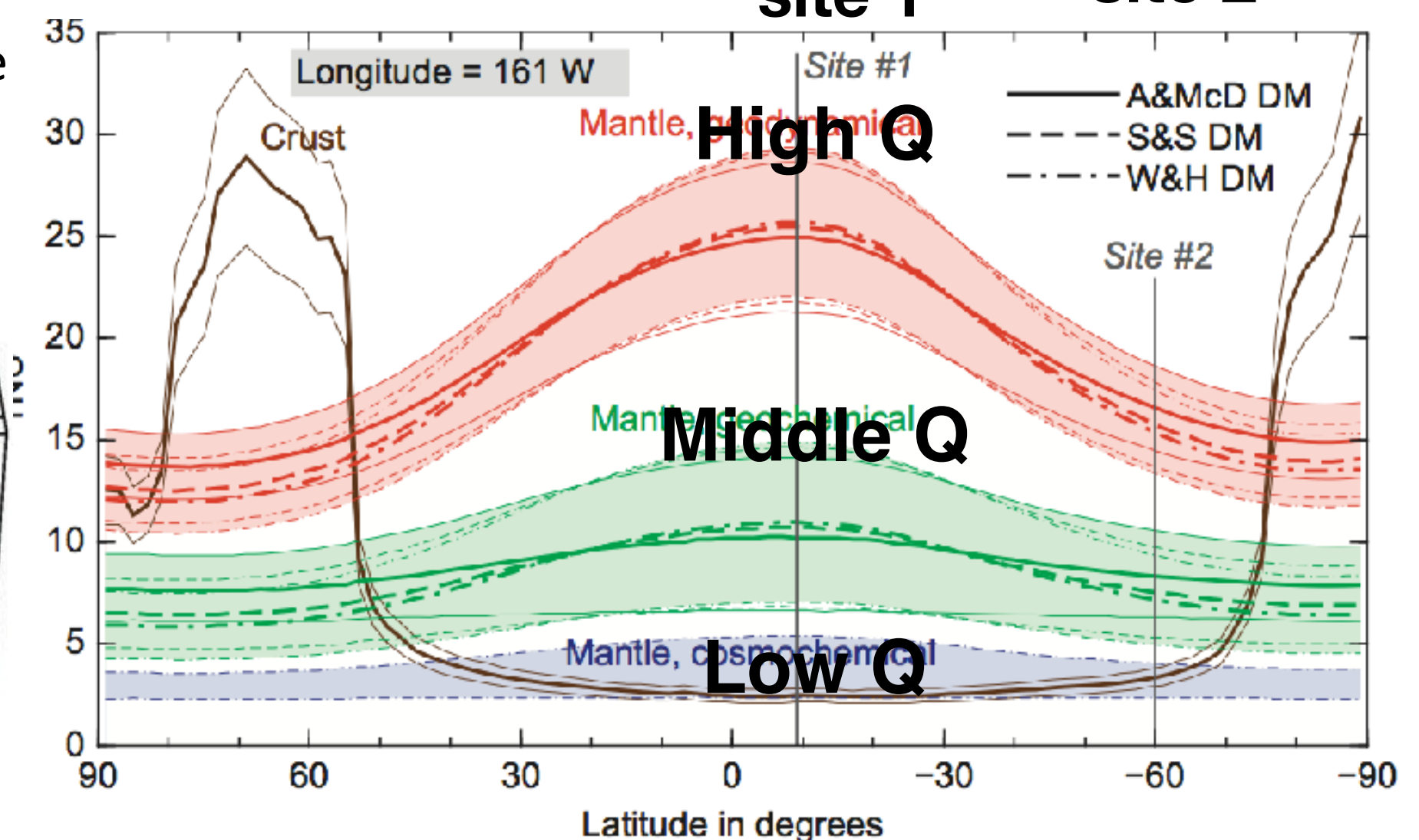
site 1

site 2

Seismically slow “red” regions in the deep mantle



From Alan McNamara after Ritsema et al (Science, 1999)



- **Direct Measurement of Mantle**

need to be far from crust
can be far from reactors

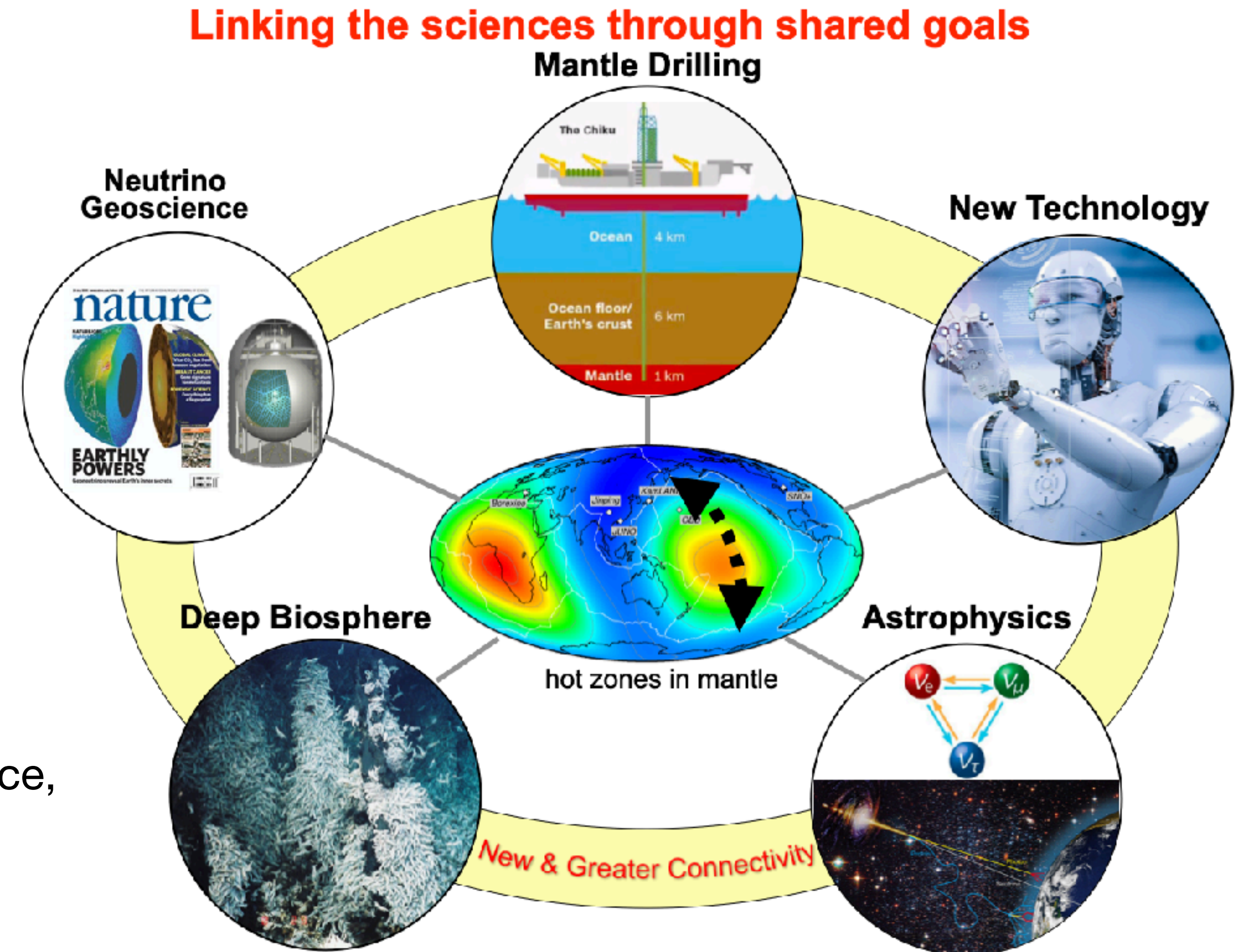
- **Multi-site Measurements**

Solve the mystery of deep Earth!

First detector for mapping **the inhomogeneous mantle**

- **Multidisciplinary Detector**

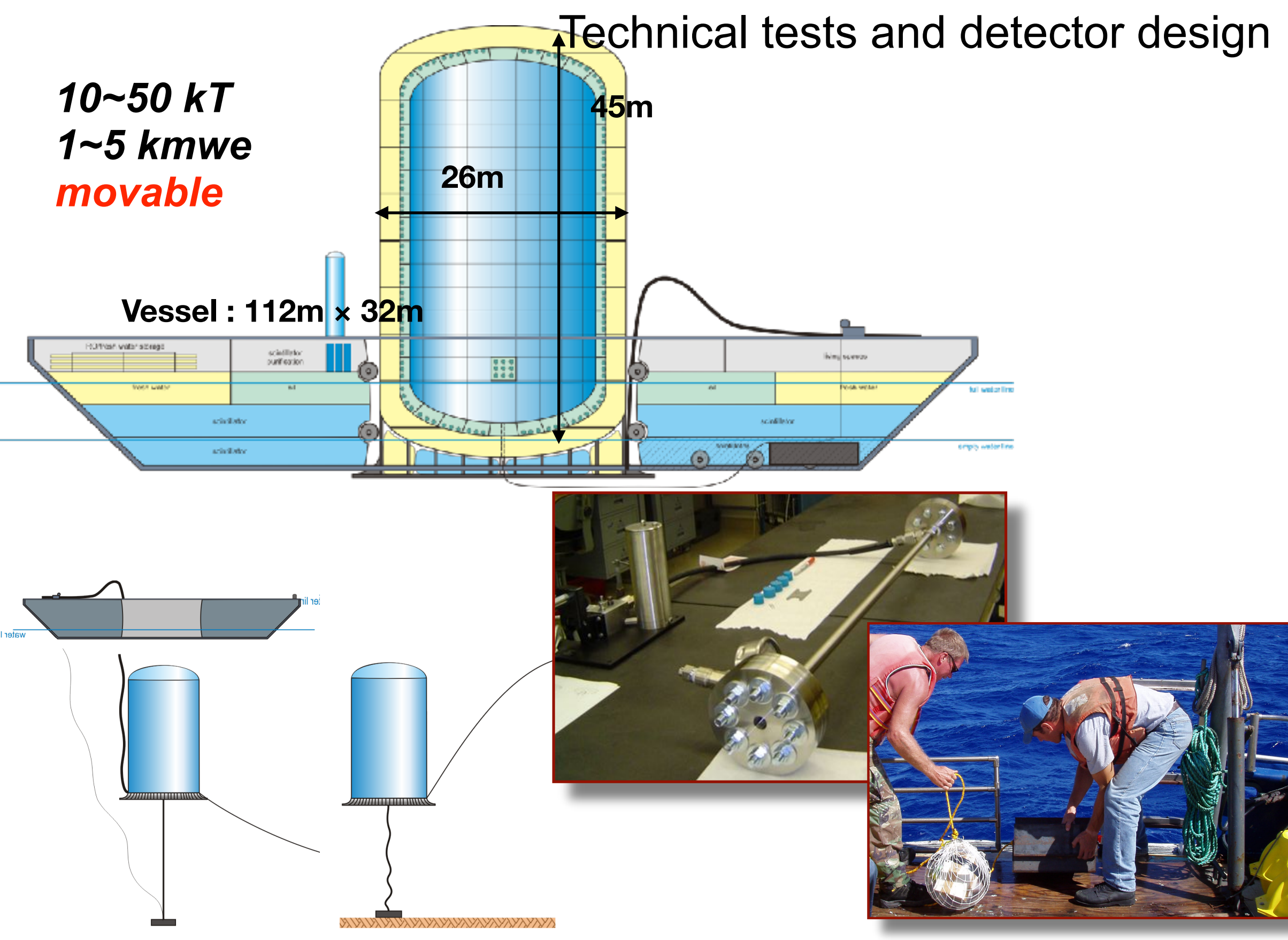
Physics, Geoscience, Mantle drilling, Environmental science,
Biology, New technology,...



OBD Present & Future



“Hanohano” 2005 U. Hawaii & Makai Ocean Engineering



A Deep Ocean Anti-Neutrino Detector near Hawaii - Hanohano

Final Report

Prepared for:

The National Defense Center of Excellence for Research in Ocean Sciences (CEROS)
73-4460 Queen Kaahumanu Highway, Suite 111
Kailua-Kona, Hawaii 96740

Prepared by:

MAKAI OCEAN ENGINEERING, INC.
Waimanalo, Hawaii 96734
<http://www.makai.com>

and

DEPARTMENT OF PHYSICS AND ASTRONOMY
University of Hawaii, Manoa

Approved for Public Release, Distribution Unlimited

arXiv > hep-ex > arXiv:0810.4975v1

High Energy Physics - Experiment

[Submitted on 28 Oct 2008]

Hanohano: A Deep Ocean Anti-Neutrino Detector for Unique Neutrino Physics and Geophysics Studies

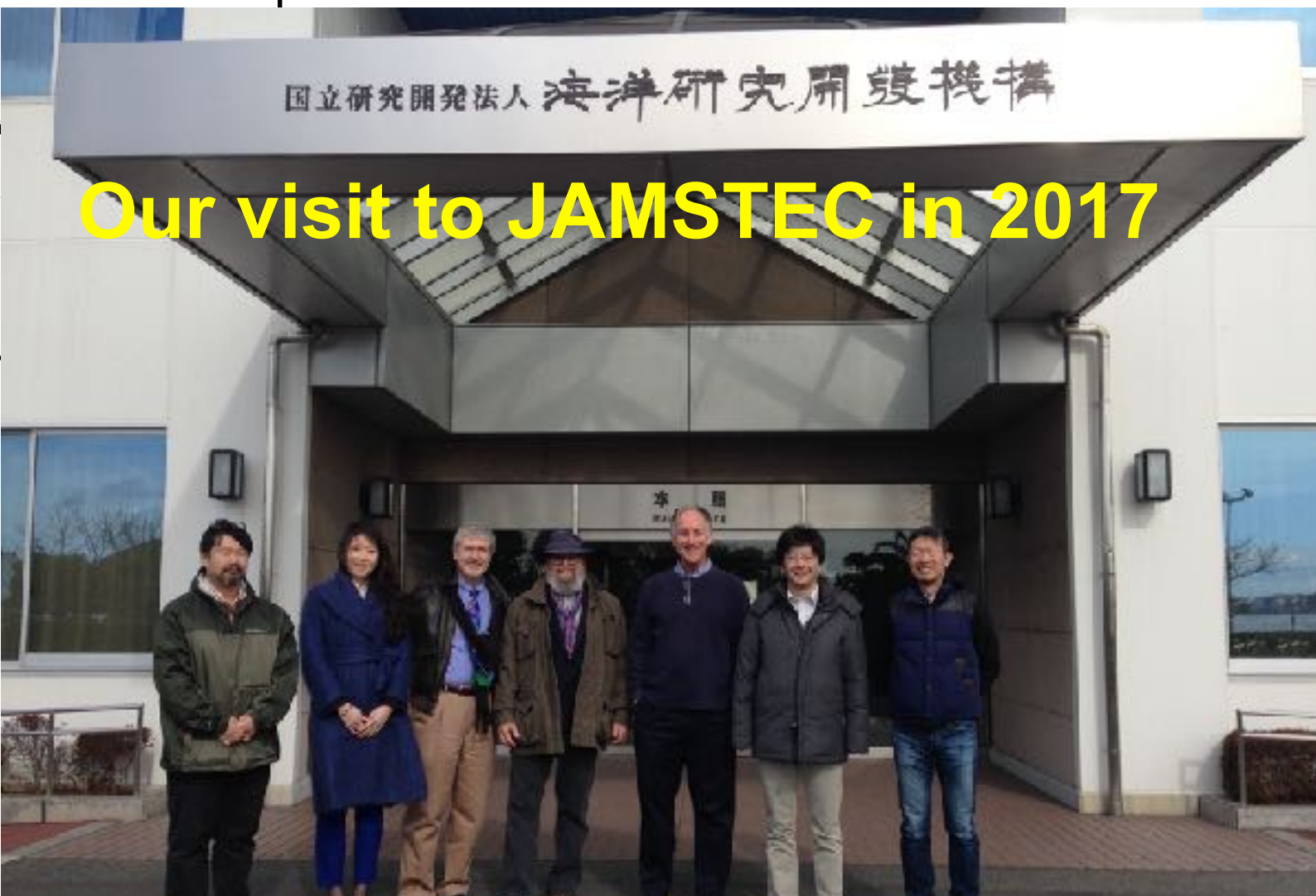
John G. Learned, Stephen T. Dye, Sandip Pakvasa

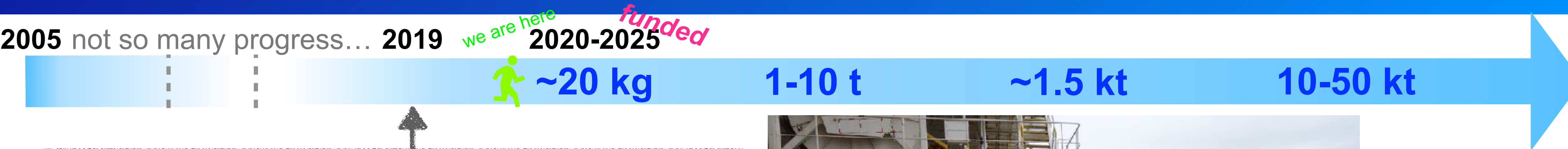
The science potential of a 10 kiloton deep-ocean liquid scintillation detector for ~1 MeV energy scale electron anti-neutrinos has been studied. Such an instrument, designed to be portable and function in the deep ocean (3-5 km) can make unique measurements of the anti-neutrinos from radioactive decays in the Earth's mantle. This information speaks to some of the most fundamental questions in geology about the origin of the Earth, plate tectonics, the geomagnetic field and even somewhat indirectly to global warming. Measurements in multiple locations will strengthen the potential insights. On the particle physics side, we have identified a unique role in the study of anti-neutrinos from a nuclear power complex, at a range of 55-60 km off shore. Not only can precision measurements be made of most neutrino mixing parameters, including θ_{13} (depending on magnitude), but the neutrino mass hierarchy can be determined in a method not heretofore discussed, and one which does not rely upon matter effects. This detector is under active study on paper, in the laboratory, and at sea. An interdisciplinary and international collaboration is in formation, and plans are in motion for a major proposal, to be followed by construction over several years.

Comments: 35 pages, 23 figures. Published in the Proceedings of the Twelfth International Workshop on Neutrino Telescopes, Venice, March 2007

Subjects: High Energy Physics - Experiment (hep-ex); High Energy Physics - Phenomenology (hep-ph)

Cite as: [arXiv:0810.4975](https://arxiv.org/abs/0810.4975) [hep-ex]
(or [arXiv:0810.4975v1](https://arxiv.org/abs/0810.4975v1) [hep-ex] for this version)
<https://doi.org/10.48550/arXiv.0810.4975>





OBD project has started with JAMSTEC* & Tohoku U.

* Japan Agency for Marine-Earth Science and Technology

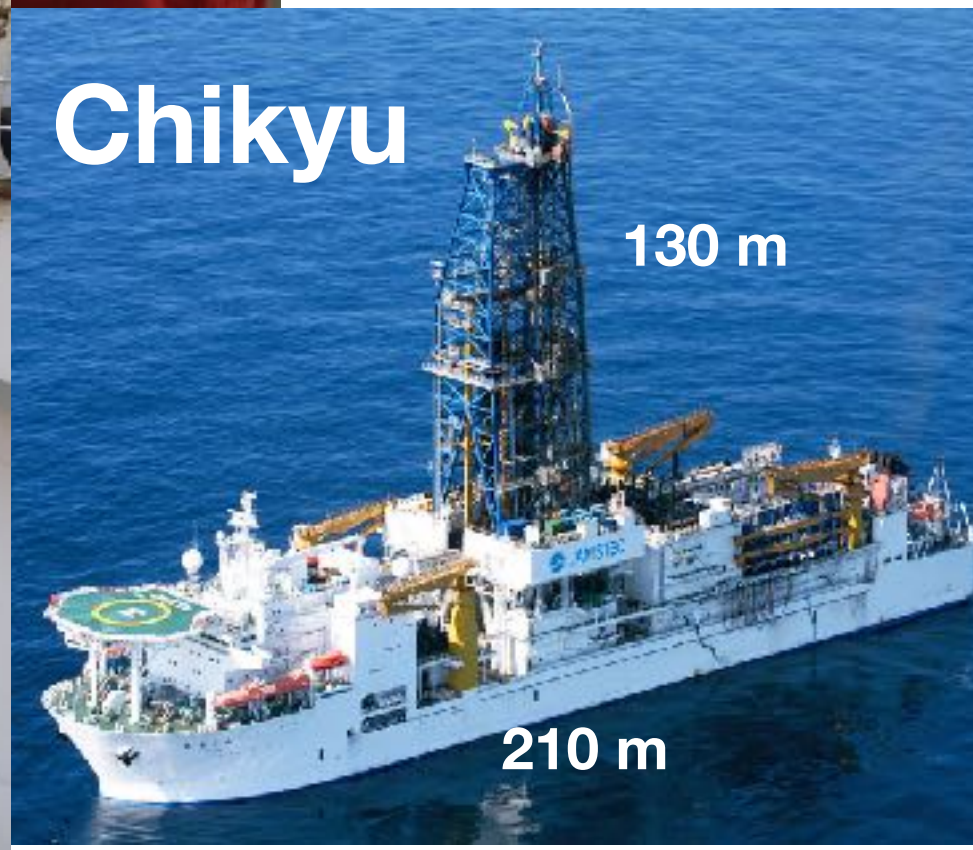


Generator
5,000kW×6 2,500kW×2

Equal to the power needed by a town of 10,000 homes (3 kW per home)



“moon pool”
- 12m x 22m
- cart : < 400t





Technical test & world's first measurement in the ocean with LS detector

* Install detector into ~1km seafloor (JAMSTEC's Hatsushima Observatory) take data for **several months**

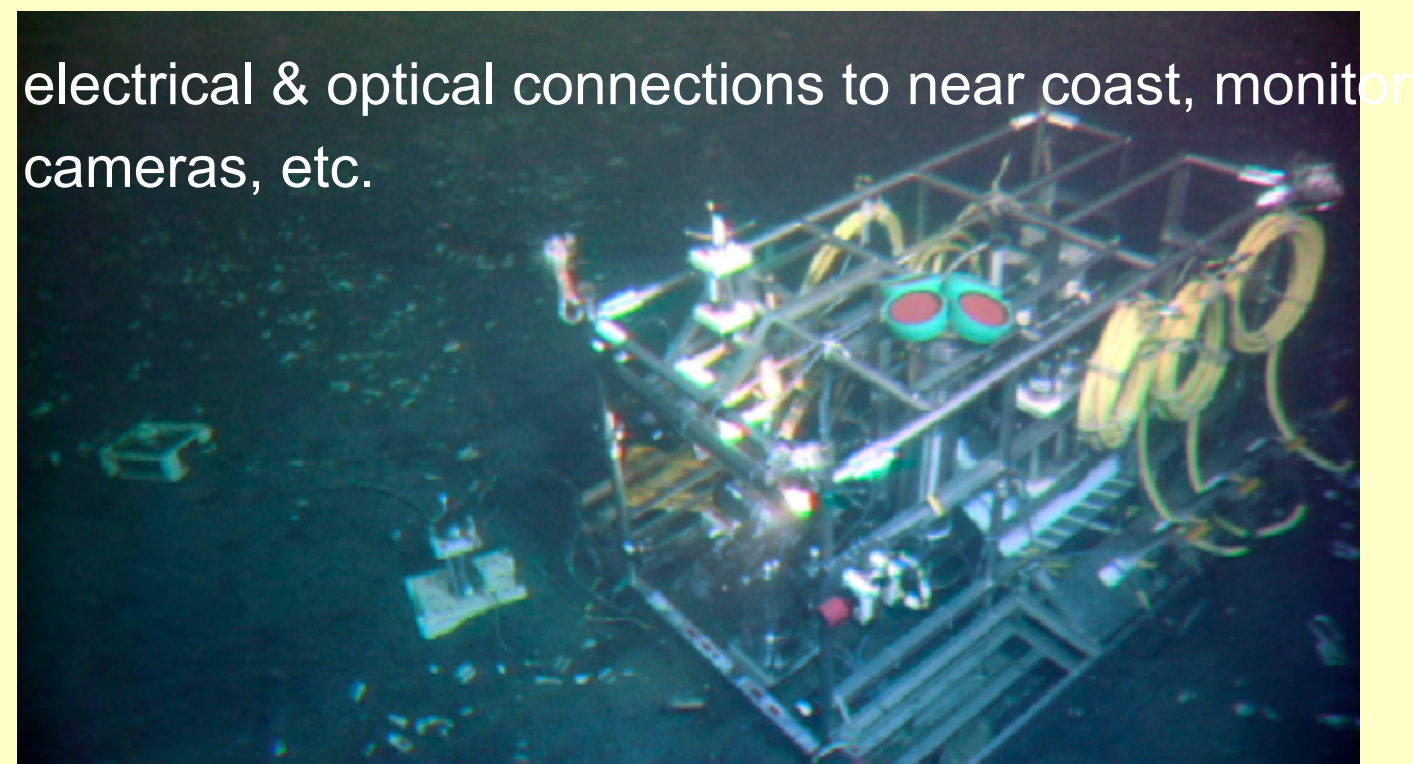
Another option: seafloor around Hawaii with battery operation

* **measure muon late in the sea** → input parameter for future large detector

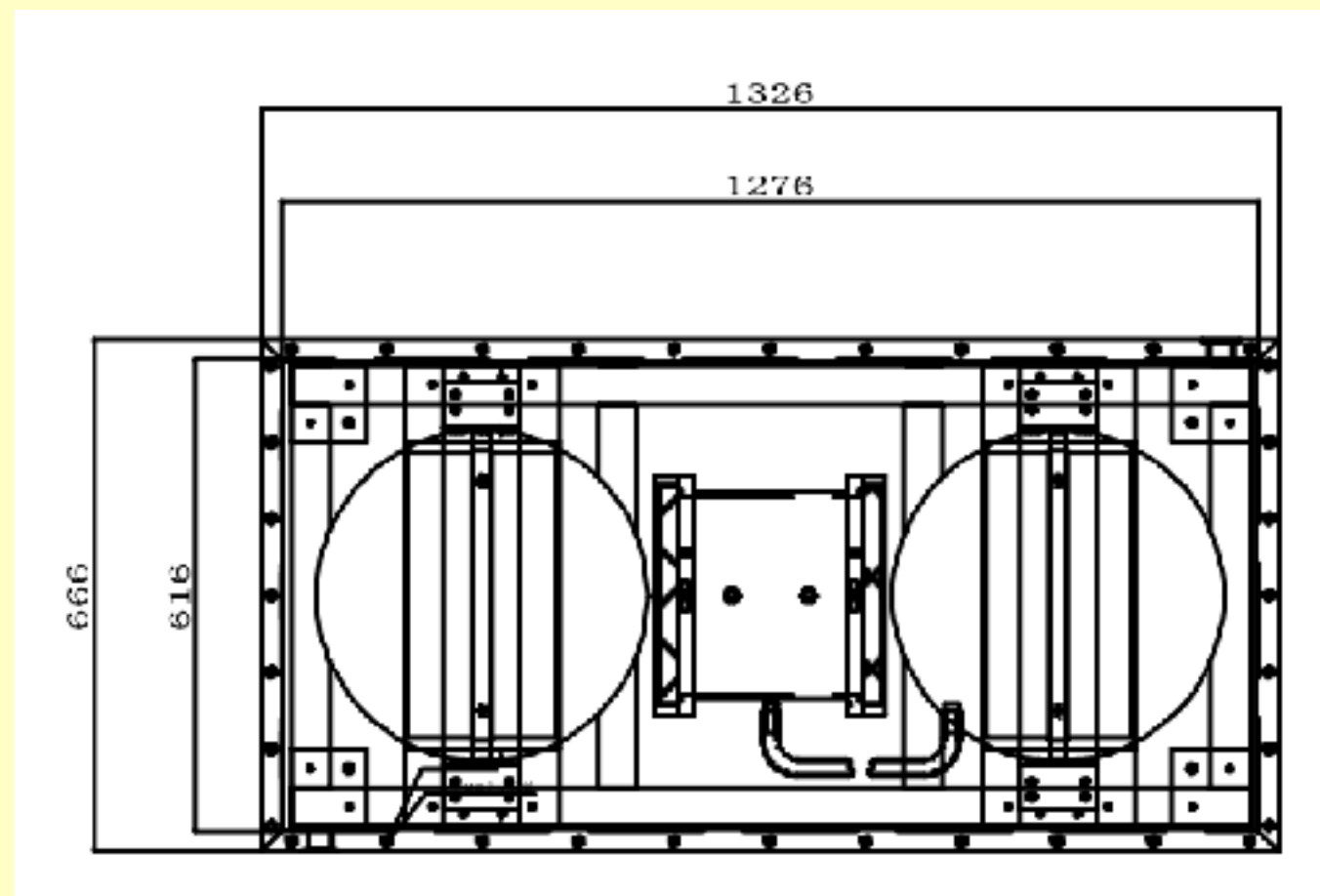
* Technical developments are in progress.

* on-going: Study on glass-shielded PMT+electronics module with IceCube team in Chiba

Hatsushima Observatory



detector design



Construction test at Tohoku U. with glass-shielded PMTs



OBD Present & Future

20/23



Technical demonstration & environment measurement in the sea
deep sea neutrino & muon flux, ocean water density & temperature, radioactivity
→ input parameters for ~1.5 kt detector design

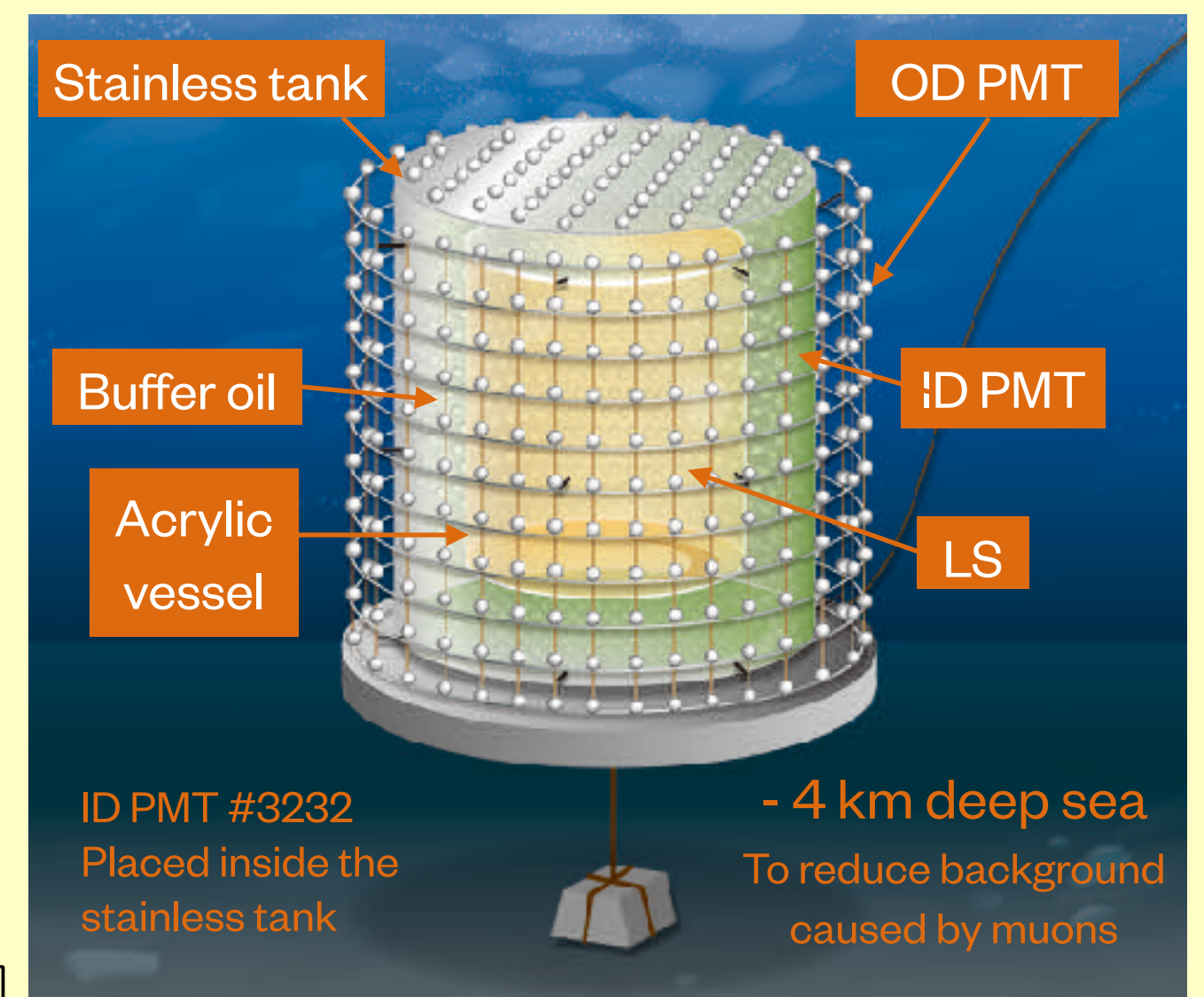
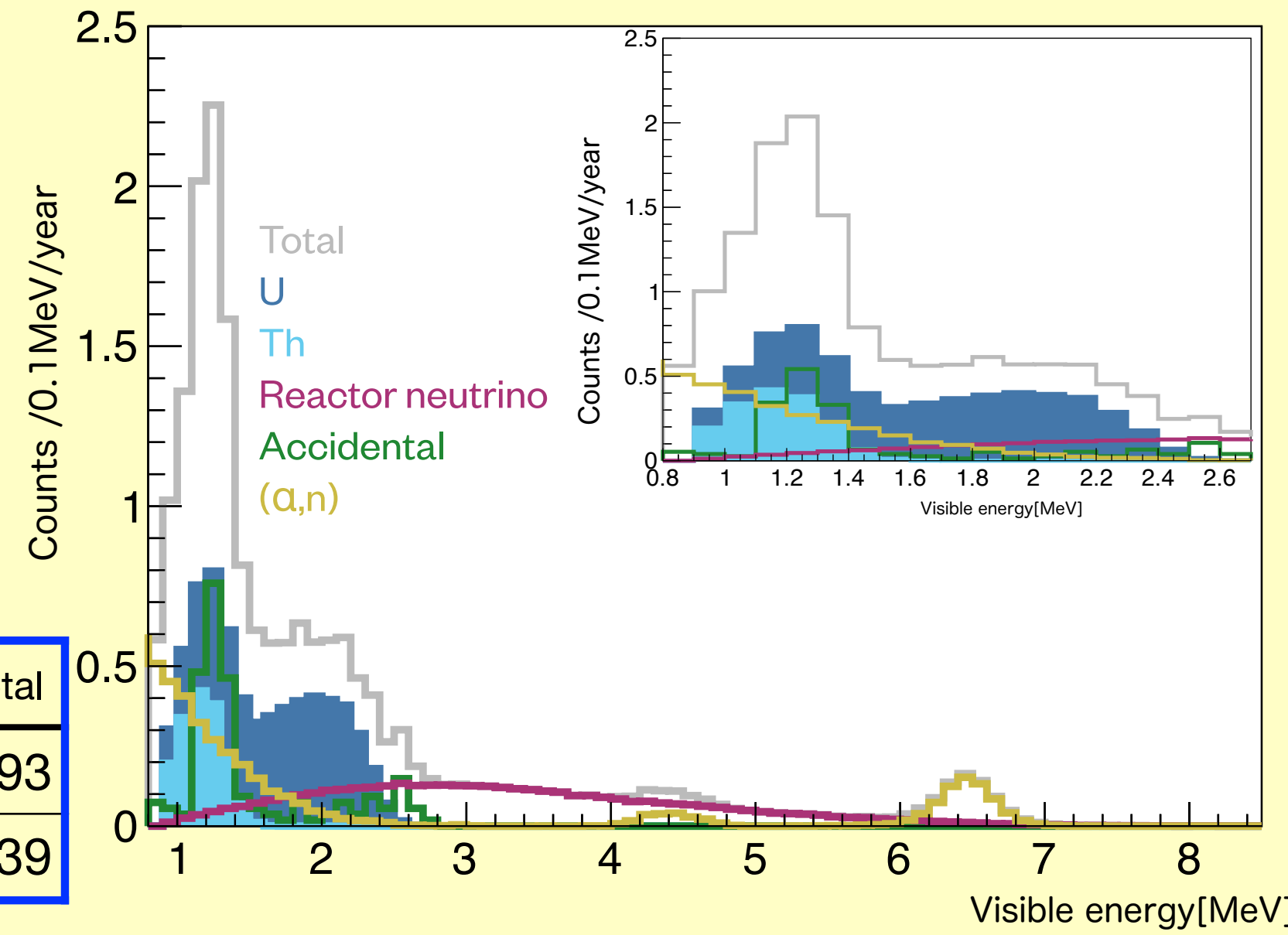
First clear mantle signal

- * Detector simulation study is in progress.
- * Hawaii is possible position.
- * Detector should be installed at ~4km deep sea to shield muons

Low temperature (2-4°C)
high pressure (40MPa)

E reagon	Signal			Backgrounds						Total
	U	Th	Total	Reactor	Acci.	(α,n)	He-Li	Fast-neutron		
All	6.59	1.64	8.23	4.13	1.92	3.88	0	<2.42	9.93	
Geo-nu	(4.61)	(1.15)	(5.76)	1.53	1.90	2.96	0	<0.58	6.39	

(mantle) [Events/year]



*** Mantle geoneutrino sensitivity**

highQ model:	1year → 3.7σ	* 17 kt detector case	1month → 4.1σ
middleQ model:	3year → 3.5σ		3month → 3.9σ
lowQ model:	10year → 2.5σ		1year → 3.0σ

Technical Developments

◆ Liquid scintillator

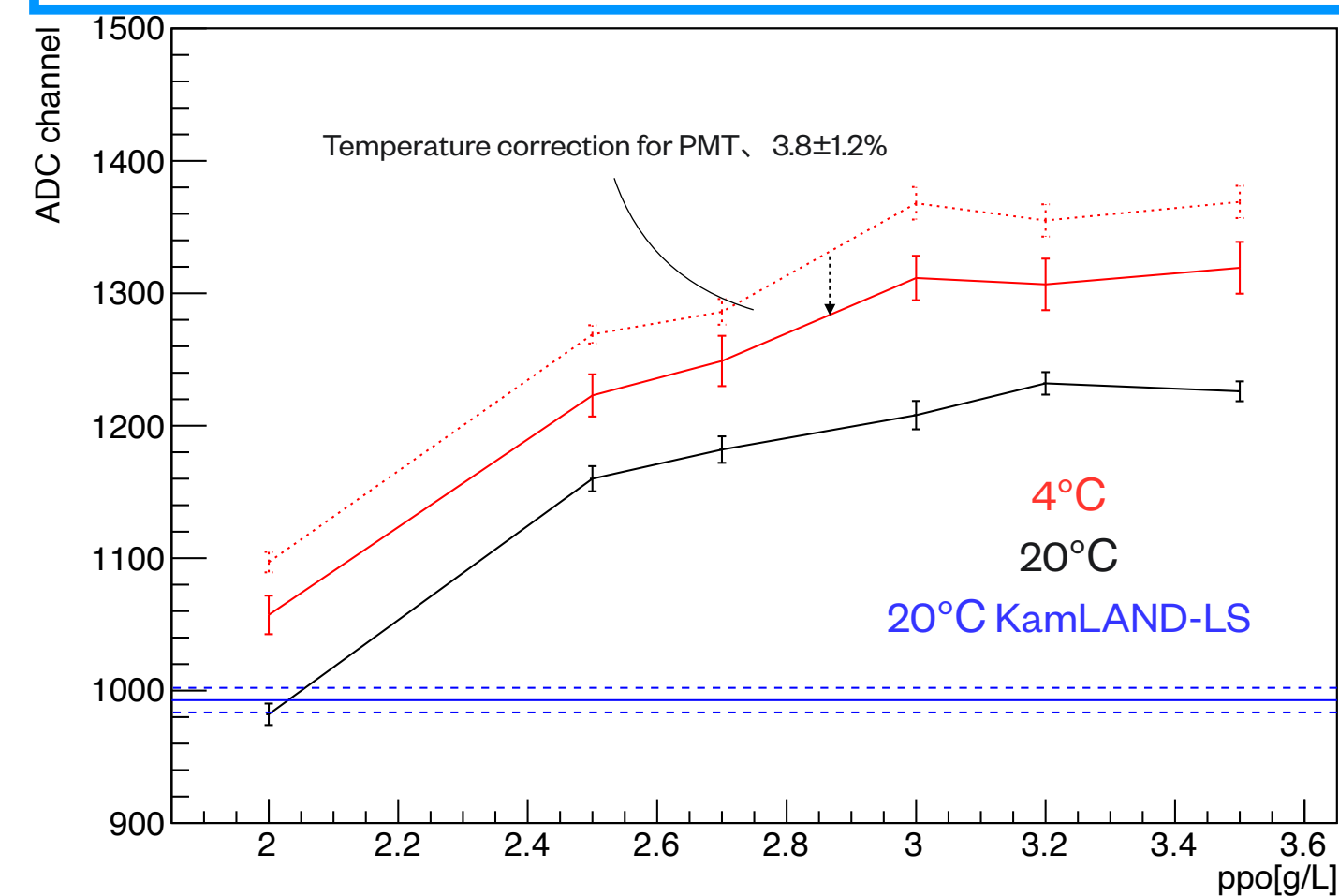
LAB(oil) + PPO(fluorescents)

Low temperature

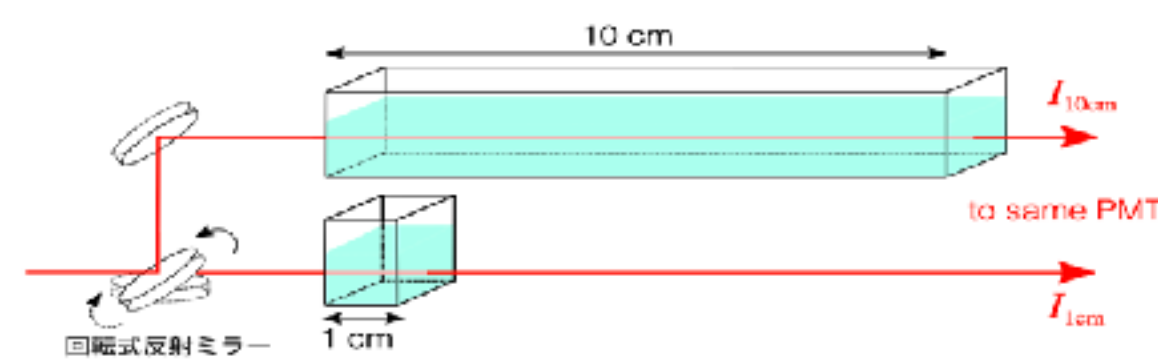
light yield



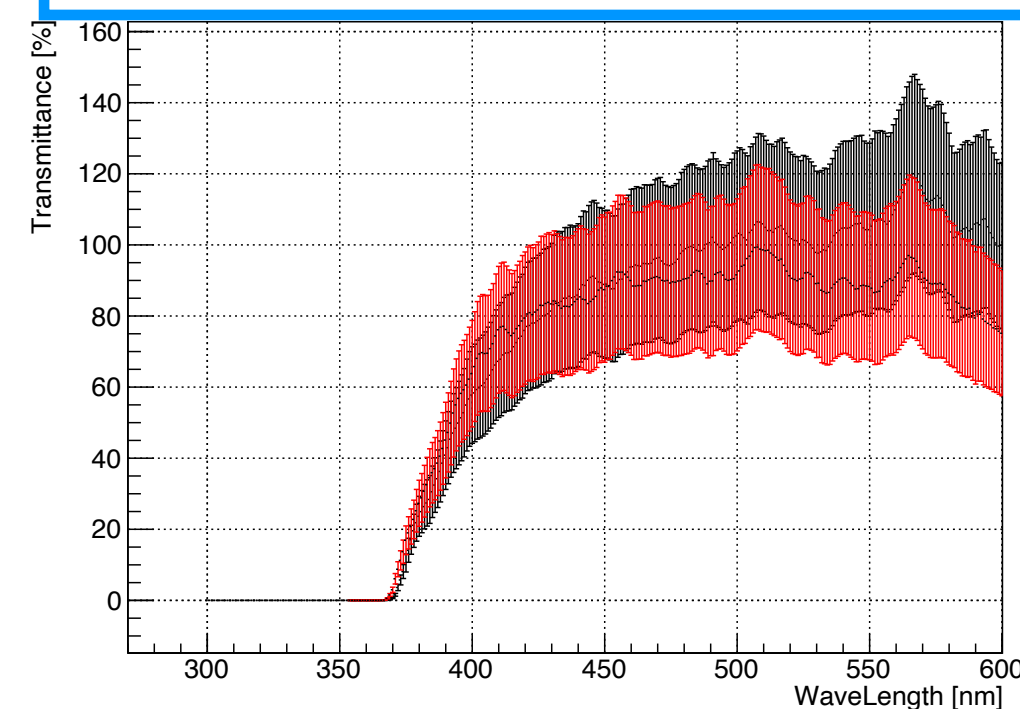
4°C is brighter than 20°C (+9%)



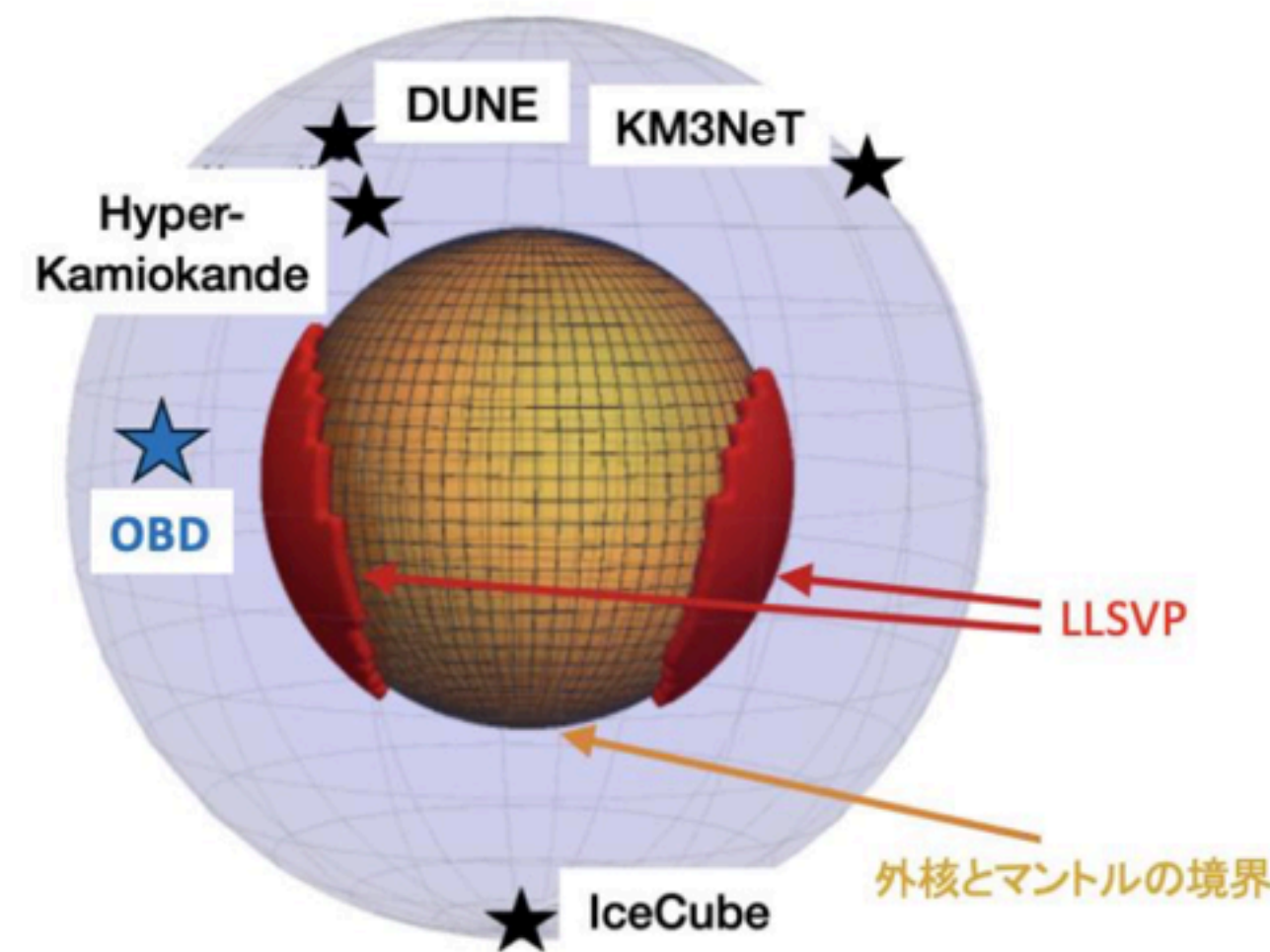
light transition



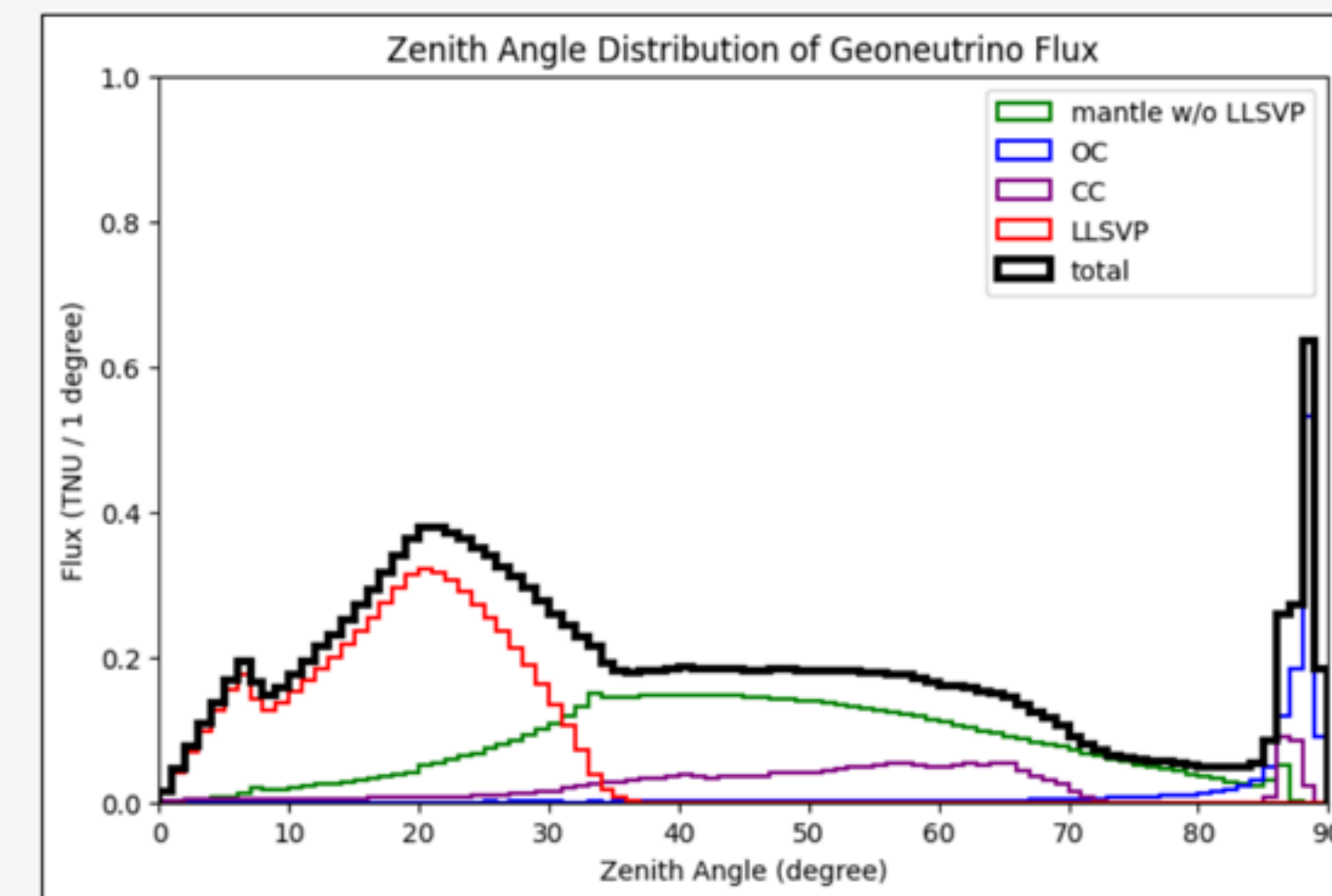
no temperature affect



Simulation Studies: LLSVP



◆ Directional Sensitive Detector



◆ Oscillation Tomography with high energy atmospheric-ν

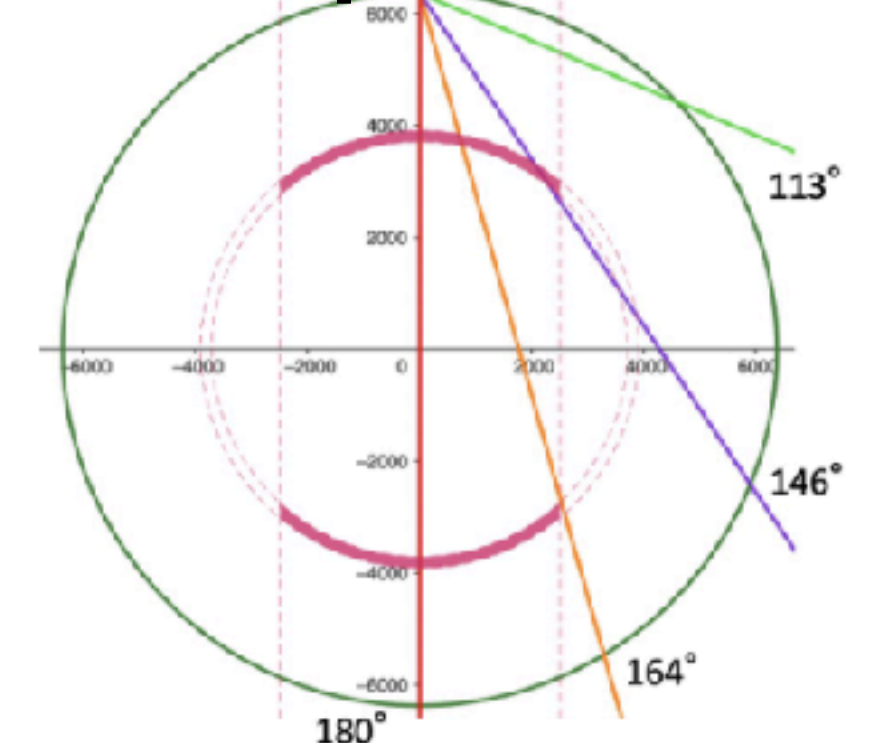
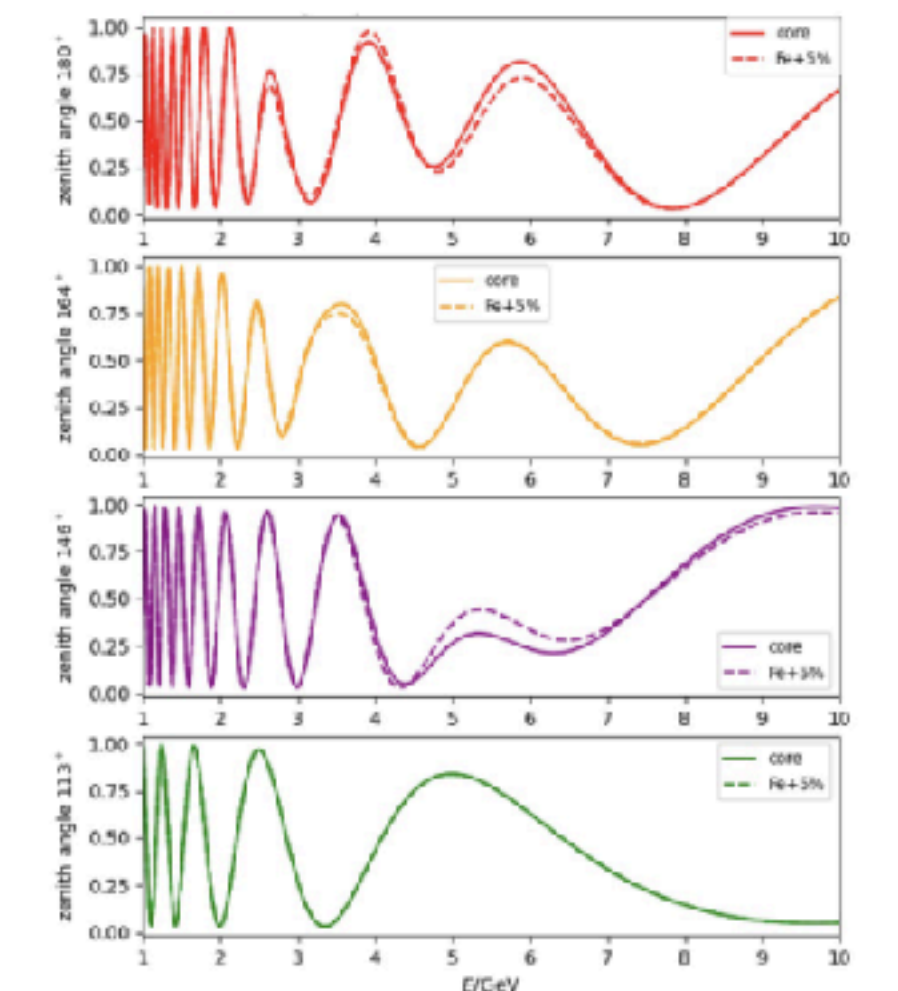


図 6.3: 大気ニュートリノの天頂角ごとの経路



International Collaboration: Tohoku U., U. Hawaii, LLNL



- Started from 2024
- Writing White Paper
- Discussions on:
 - Detector design
 - Detector technology
 - Detector sensitivity
 - Expand collaborative scientific fields
 - Funding strategy etc



- **Geoneutrinos** are unique and new tool to measure directly the Earth's interior.
 - ▶ **Strong way to measure amount of radioactive elements in the Earth**
- To date, physics experiments have shown the usefulness of geoneutrinos.
 - ▶ **Interdisciplinary community has furthered its connection over these past 15 years.**
- "Neutrino Geoscience" : collaborations between geoscience, physics, ocean engineering and beyond
 - ▶ **Ocean Bottom Detector (OBD) = Breakthrough**

<Transformative insights>

OBD's Primary Goal :

- **map the mantle**
- **constrain the planet's cooling history**

- **Neutrino Geoscience 2025, October 27-30, Kingston!**
 - ▶ Hawaii 2005, Hawaii 2007, Sudbury 2008,
Gran Sasso 2010, Japan 2013, Czech 2019



Neutrino Geoscience 2025 – Kingston

October 27-30, 2025