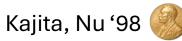


Reactor Neutrino Oscillations

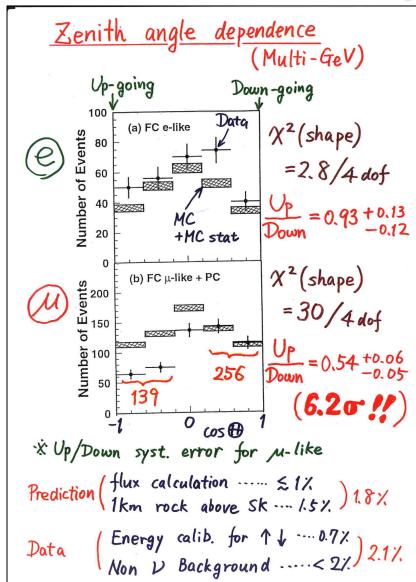
Discovery with KamLAND, Status, and Future Outlook

Jason Detwiler, University of Washington LearnedFest '25, Honolulu HI, May 2, 2025



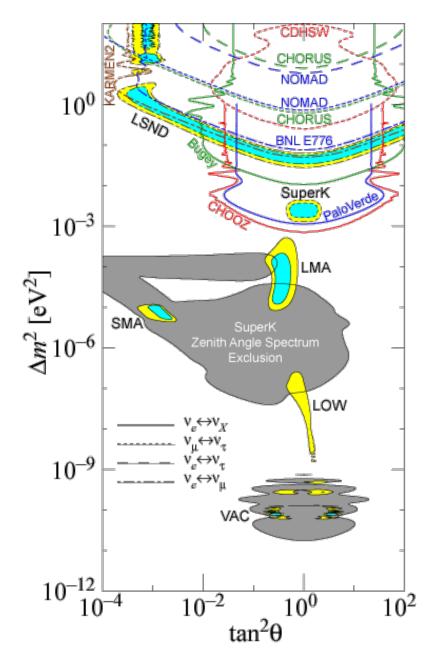
Neutrino '98: SuperKamiokande!

- Kajita-sensei announced the first evidence of neutrino oscillation on June 5 at the Neutrino conference in Takayama.
- I was still an undergrad and was working at LANL that summer in condensed matter experiment. I remember this being the most exciting science news of that summer.
- I believe it was in 1999 I was presenting at the APS meeting in Atlanta and attended a session on neutrino oscillation where a young Giorgio Gratta mentioned this new reactor neutrino experiment in Japan during his presentation...



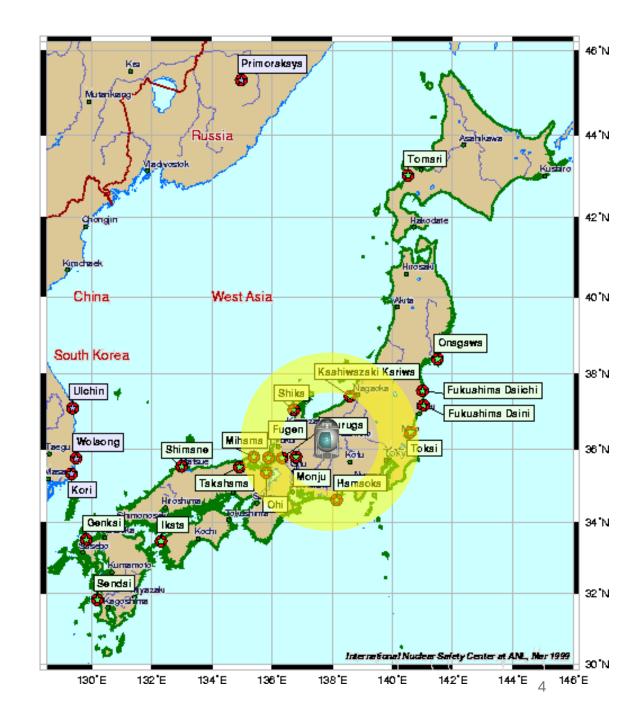
Neutrino Oscillation Pre-KamLAND

- SuperK oscillations at $\Delta m^2 \sim 10^{-3}$
- LSND hints at higher Δm^2
- Solar neutrino problem solutions:
 - MSW flavor transformation
 - Small mixing angle
 - Large mixing angle
 - Low Δm^2
 - Vacuum oscillations



KamLAND Strategy

- Atsuto Suzuki: "let's build a 1 kton LS detector in the old Kamiokande cavern"
- "Free beam" from the Japanese nuclear power industry, with convenient distribution
- Sensitive to LMA solution
 - Early '90's theorists: "Don't waste your time, it's obviously SMA like the quarks..."

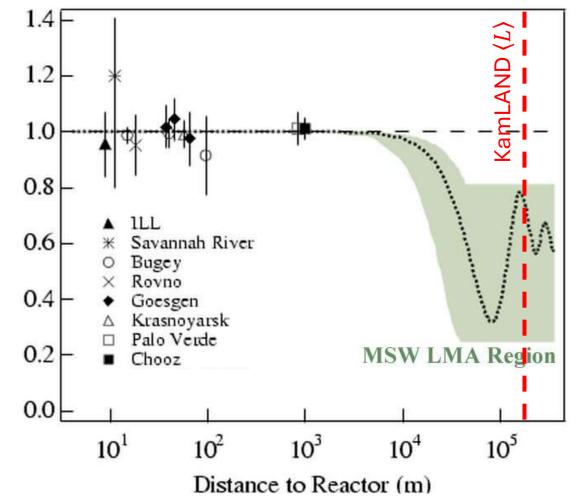


KamLAND Strategy

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Nobs/Nexp

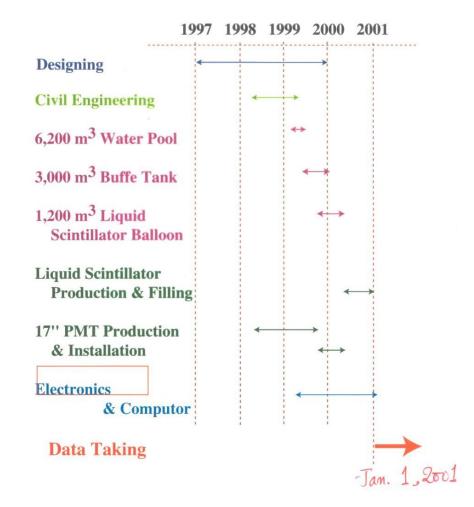
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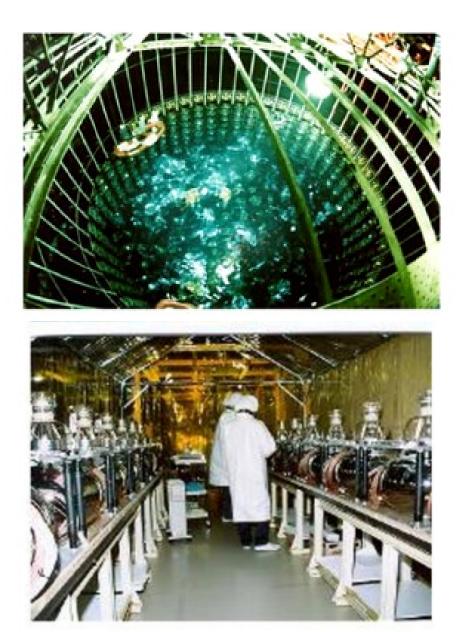
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- '99: major infrastructure construction DOE funding approval
- '00: PMT / balloon installation
- '01: filling, commissioning
- '02: start of data taking

Suzuki, Nu '98

KamLAND Construction Schedule



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KamLAND Collaboration Agreement Between the Collaborating Groups

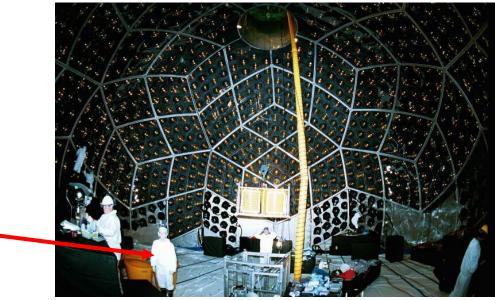
Purpose

The purpose of this document is to define the terms and conditions under which the collaborating groups (at present Japanese and American) agree to work together in building and operating a liquid scintillator anti-neutrino detector at the Kamioka mine. The experiment shall be known as KamLAND. The goals of the experiment include a search for reactor neutrino oscillations, solar neutrinos, geoneutrinos, and studies of/searches for other astrophysical and particle physics phenomena. It is agreed that all collaborating groups are free to participate in all aspects of the experiment.

```
Research Center for Neutrino Science, Tohoku University+
        K. Furuno, J. Goldman, H. Hanada, K. Inoue, K. Ishihara, M. Kishimoto, M.
        Koga, T. Mitsui, M. Motoki, M. Nakajima, T. Nakajima, J. Shirai, F. Suekane,
        A. Suzuki, T. Takavama, K. Tamaee
University of Alabama 🏼
        J. Busenitz, A. Piepkee
Lawrence Berkelev National Laboratory. University of California. Berkelev-
        B. Berger, R. N. Cahn, Y. D. Chan, X. Chen, S. J. Freedman, B. K. Fujikawa, L.
        Greiner, K. T. Lesko, K.-B. Luk, H. Murayama, D. R. Nygren, C. E. Okada, A.
        W. Poon, H. M. Steiner+
California Institute of Technology+
        G. Horton-Smith, R. D. McKeown, V. Novikov, B. Tipton, P. Vogel+
Drexel University &
        C. E. Lane↔
Duke University, North Carolina State University, Triangle Universities Nuclear
Laboratorv₽
        L. DeBraeckeleer, C. R. Gould, H. Karwowski, D. Markoff, K. Nakamura, R.
        Rohm, W. Tornow, A. Younge
University of Hawaii₽
        P. Gorham, J. Learned, S. Matsuno, S. Pakvasa &
Louisiana State University+
        S. Dazeley, S. Hatakeyama, R. C. Svoboda+
University of New Mexico↔
        B. Dieterle, S. Rileve
Stanford University
        G Gratta, Y. Uchidae
University of Tennessee
        S. Berridge, W. Bugg, H. Cohn, Yu. Efremenko, Yu. Kamvshkov+
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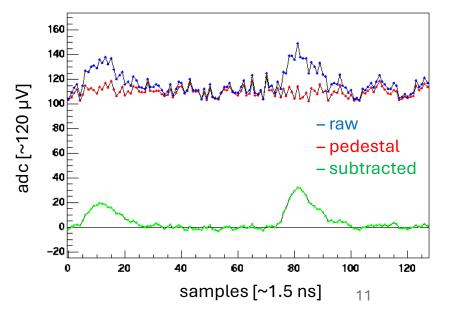
yours truly

NASA Astro Pic of the Day, June 23 2023 $^{\,\rm 10}$

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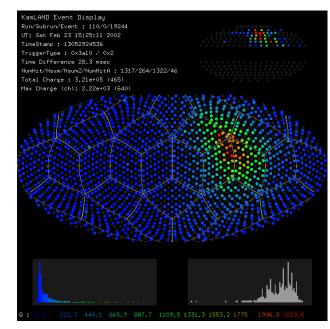




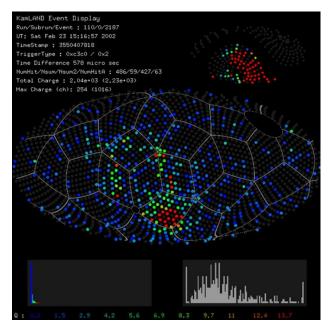
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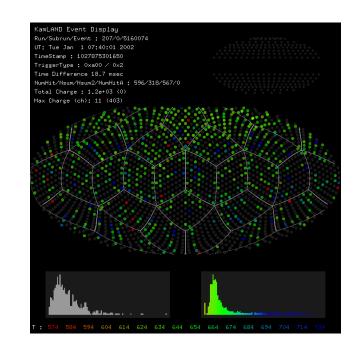




stopping $\boldsymbol{\mu}$



buffer oil muon with Cherenkov ring



first antineutrino candidate

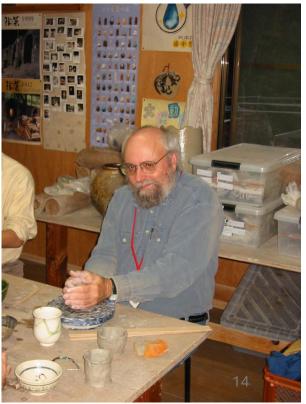
Working in Japan





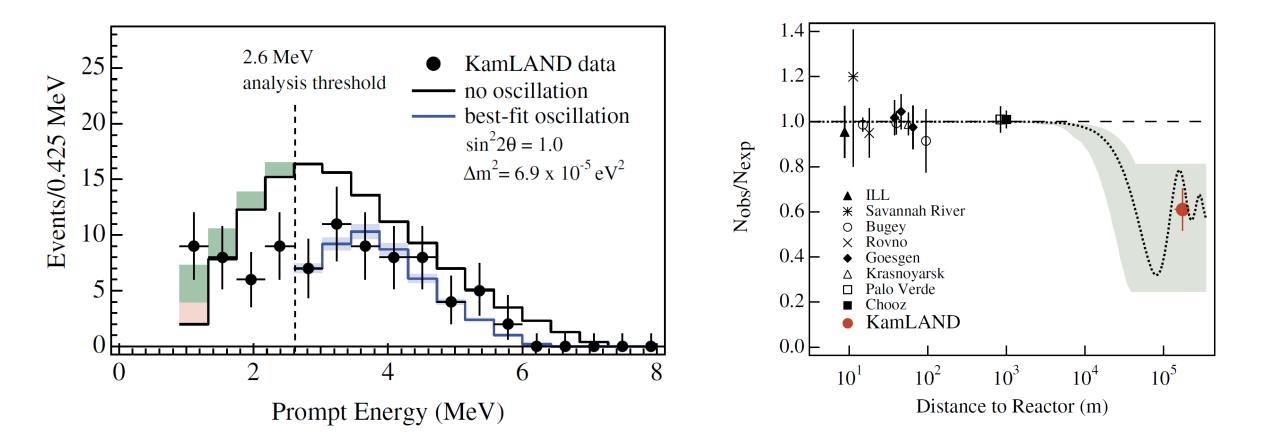






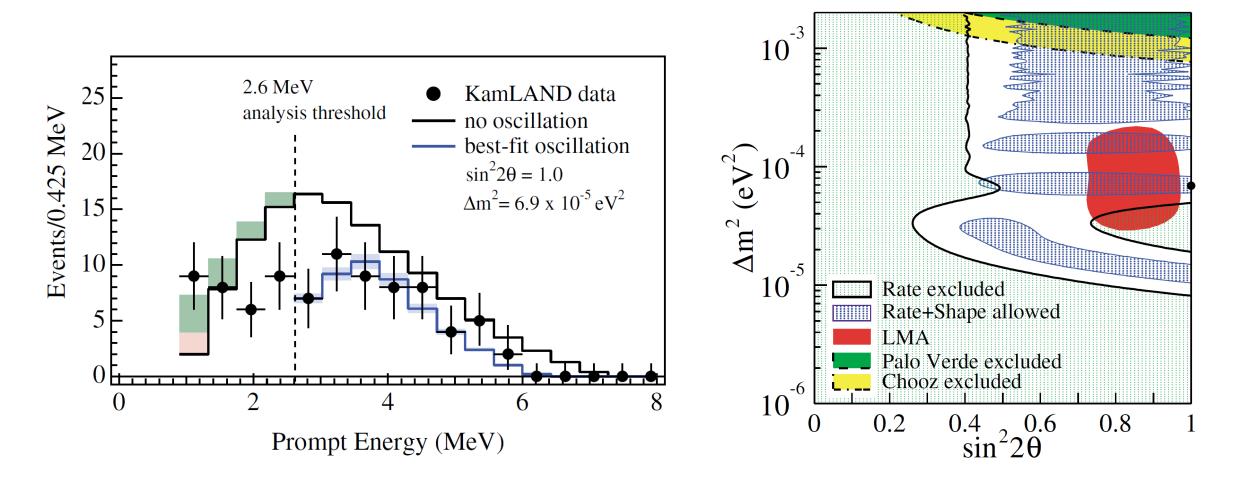


First results: reactor neutrino disappearance



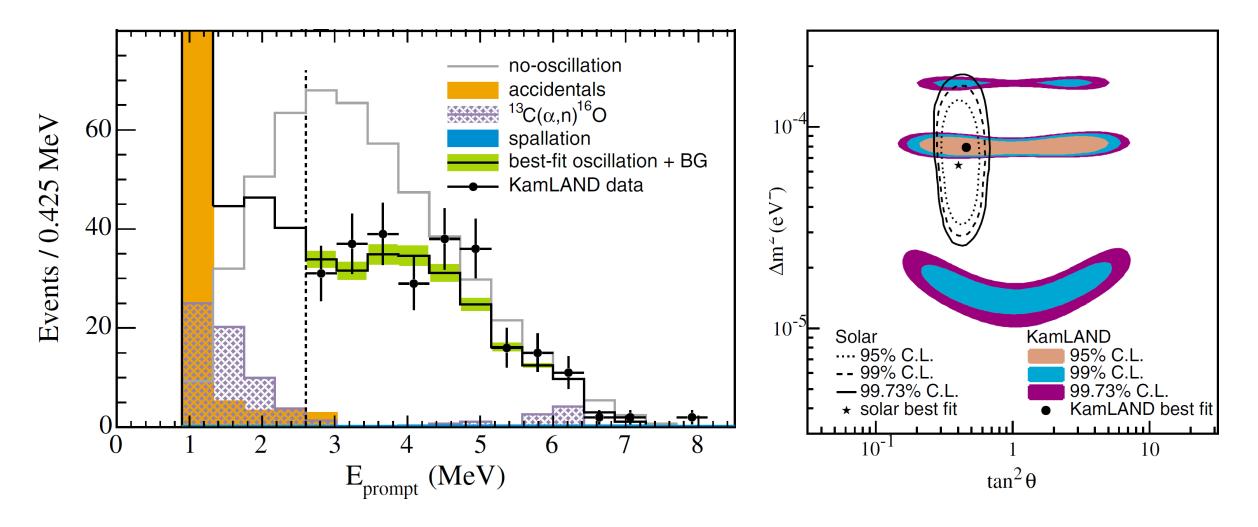
PRL 90, 021802 (2003)

First results: reactor neutrino disappearance

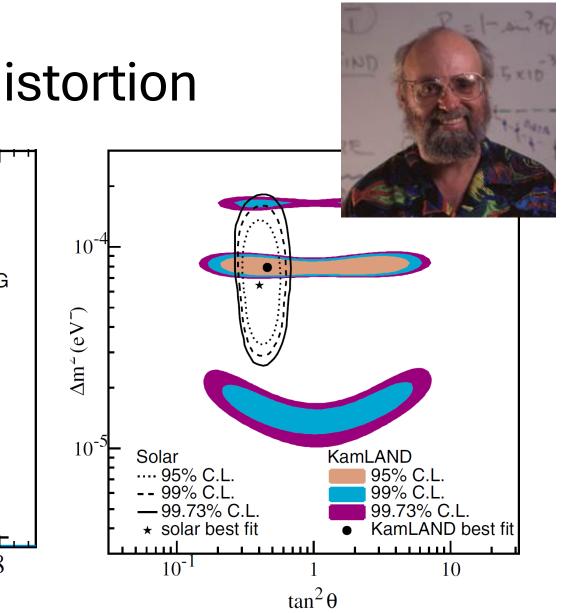


PRL 90, 021802 (2003)

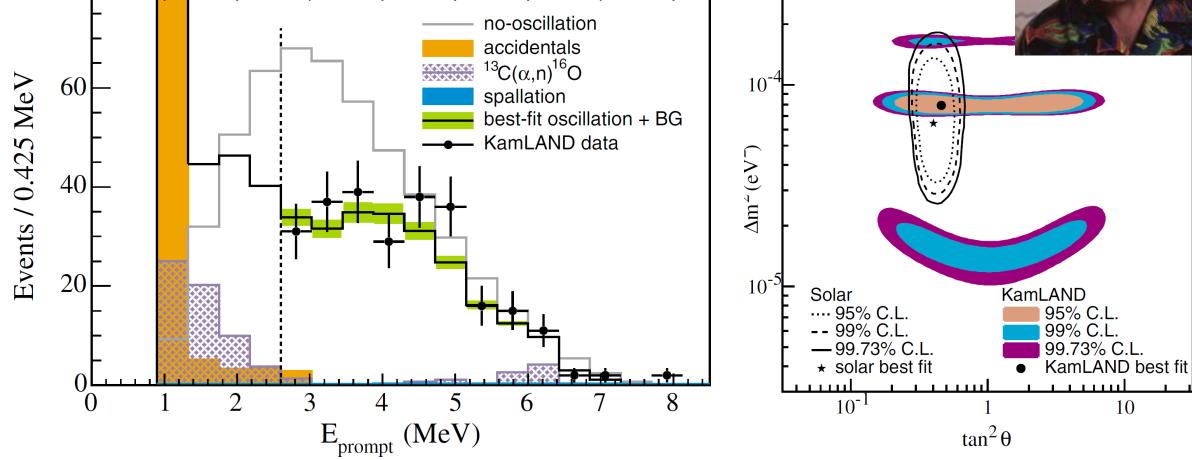
Second Result: spectral distortion

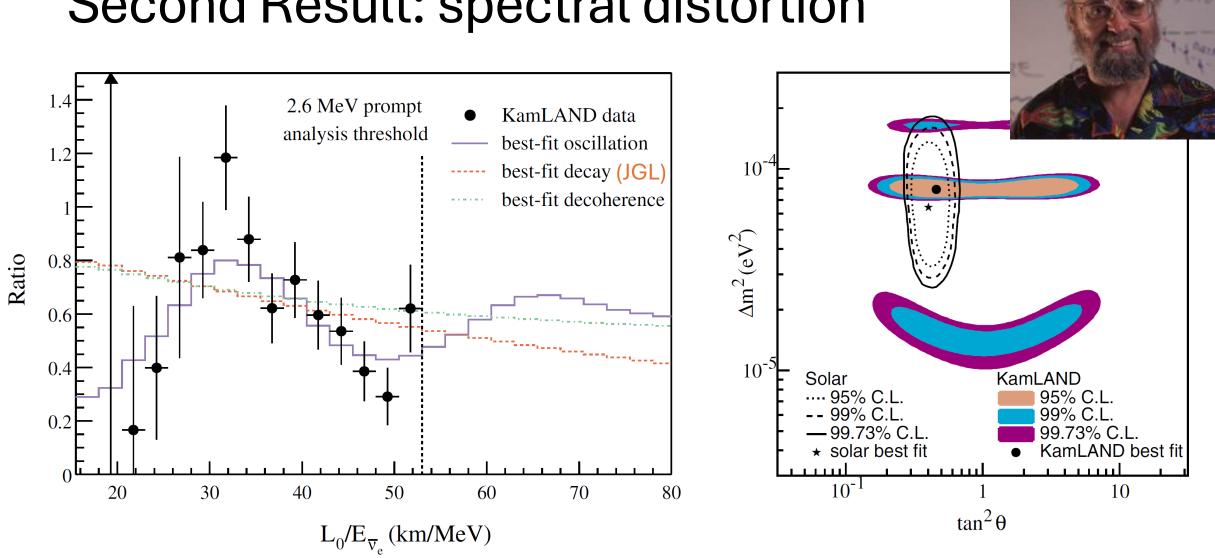


PRL 94, 081801 (2005)



Second Result: spectral distortion



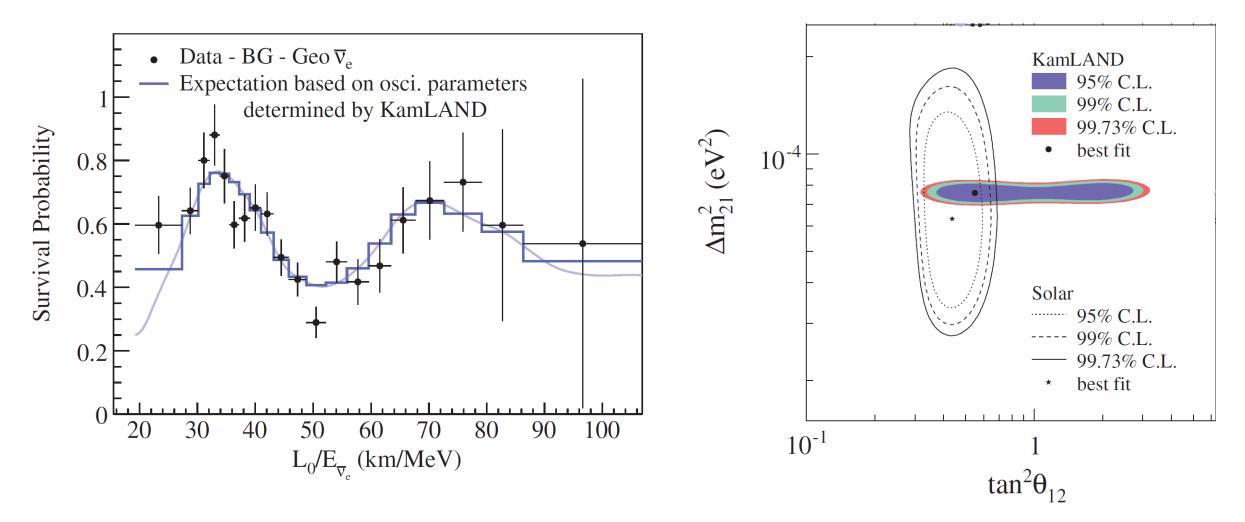


Second Result: spectral distortion

PRL 94, 081801 (2005)

NE

Third Result: precision oscillation



PRL 100, 221803 (2008)

Key Contributions

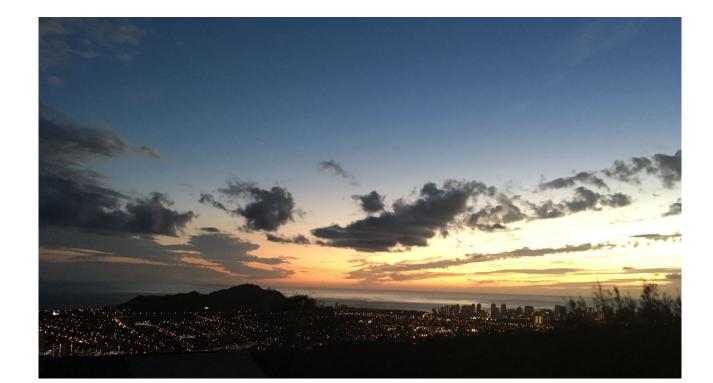






Key Contributions









Key Contributions















Disaster Strikes...

From: kuni0_in0ue@yahoo.co.jp
Sent: Saturday, March 12, 2011 3:07 AM
Subject: terrible earthquake

Dear colleagues,

We had really terrible earthquake yesterday. Electricity is not delivered yet. And we cannot send/receive e-mail with our usual account.

I'm sending this with a battery and connection is very unstable.

However, as far as I know, No one was injured in RCNS. We are just waiting for a recovery of electricity. I think RCNS people in Kamioka cannot receive e-mail, but their telephone (81-578-85-0030) may be working. I cannot call them due to regulated use of the telephone line. Please inform them and our colleagues that we are OK, please ask them to maintain KamLAND safely.

Thank you very much for your help.

Best regards, Kunio Inoue



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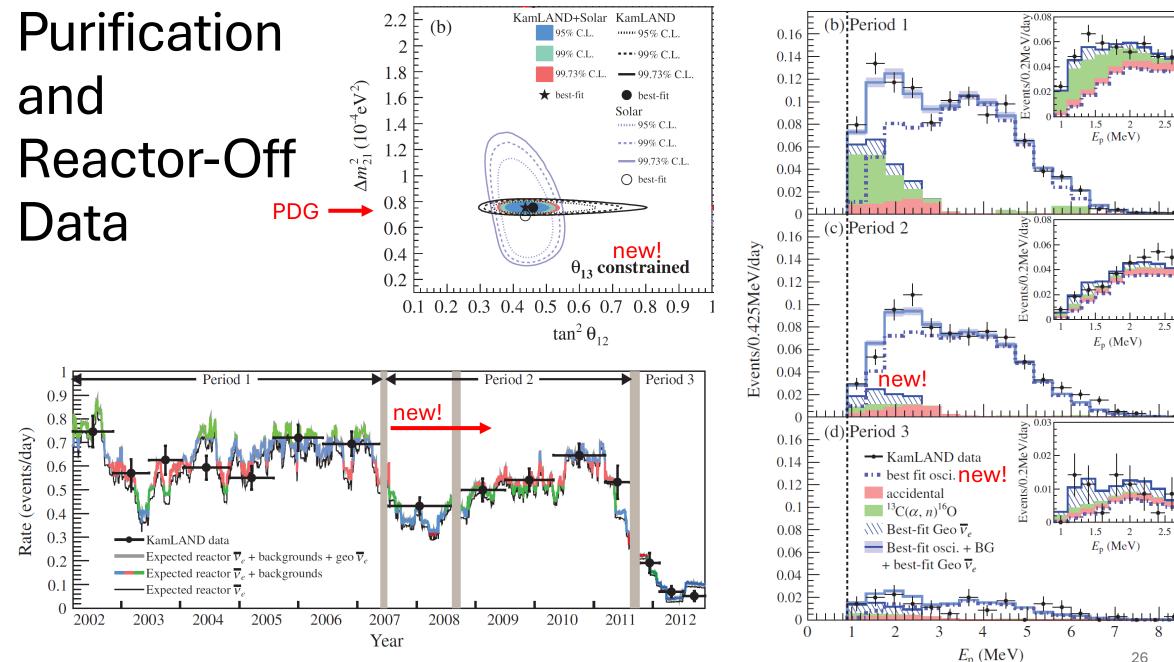
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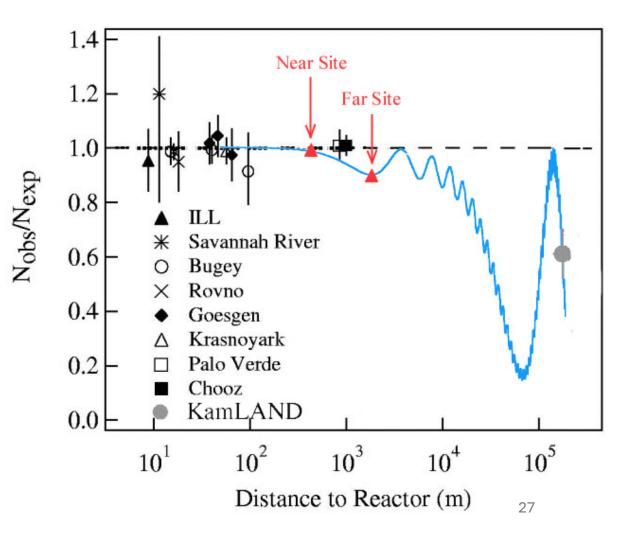
My memory: John was one of the most active on the collaboration in monitoring and communicating the situation, and in working to find ways to help!



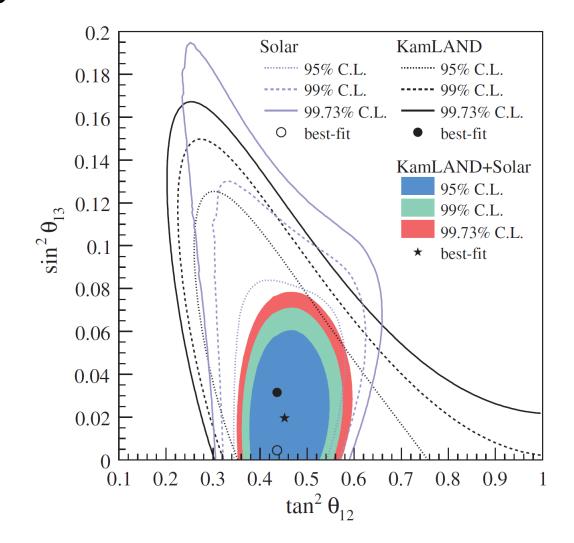
PRD 88, 033001 (2013)

26

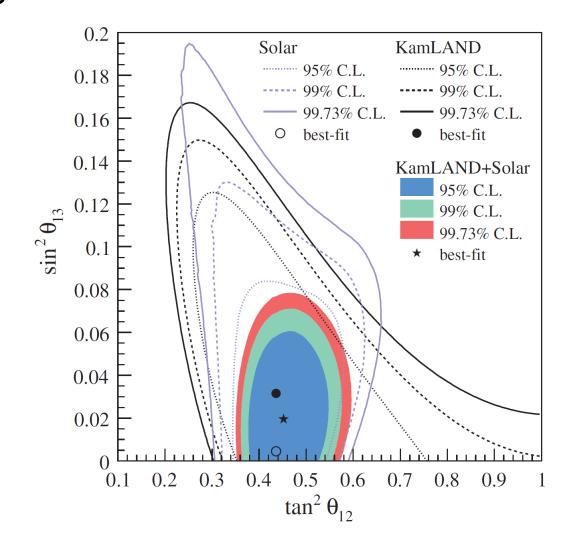
• Stringent CHOOZ limit: is θ_{13} too small for CPV to be measured?

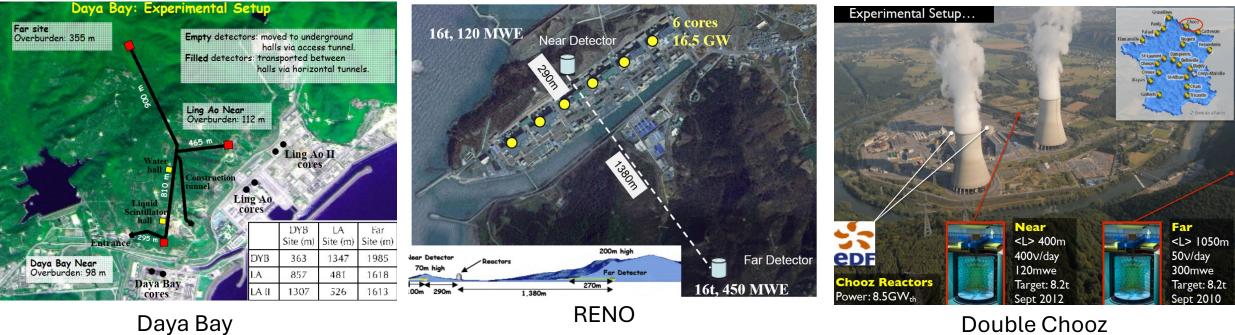


- Stringent CHOOZ limit: is θ_{13} too small for CPV to be measured?
- Non-zero θ_{13} relieved tension between KamLAND & solar θ_{12}

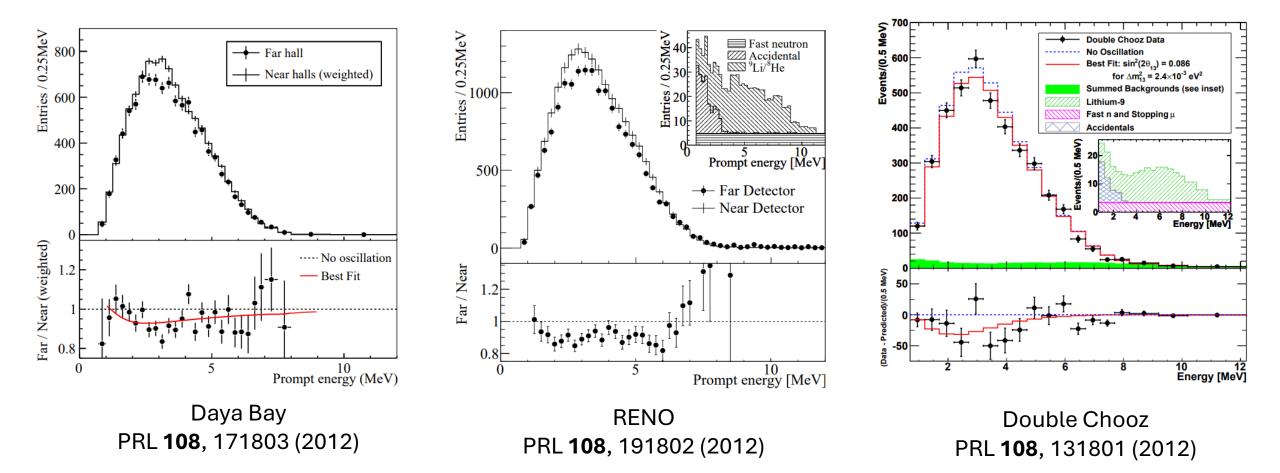


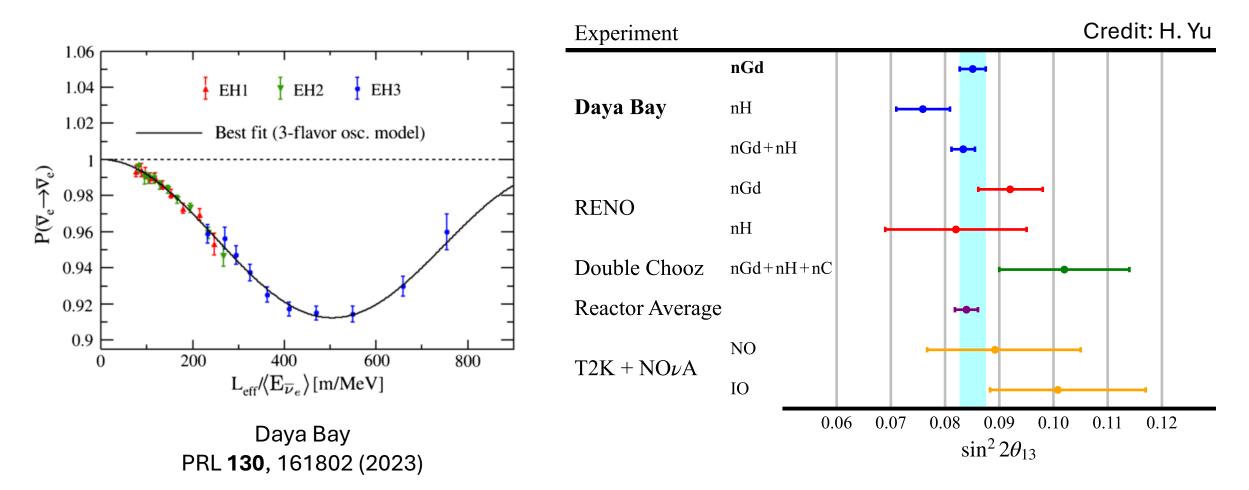
- Stringent CHOOZ limit: is θ_{13} too small for CPV to be measured?
- Non-zero θ_{13} relieved tension between KamLAND & solar θ_{12}
- Daya Bay, RENO, and Double-Chooz performed definitive measurements





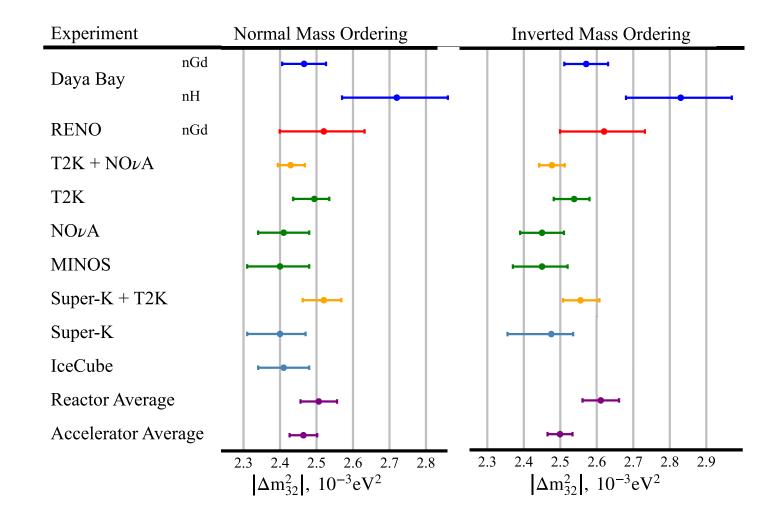
Double Chooz





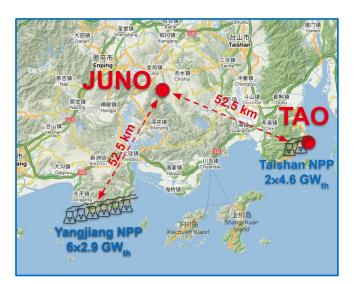
Neutrino Mass Hierarchy with Reactors

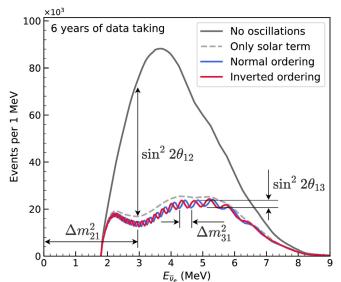
 Daya Bay + RENO gives a weak preference for the normal ordering when combined with accelerator and atmospheric measurements

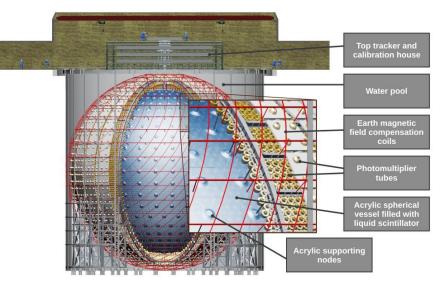


Neutrino Mass Hierarchy with Reactors

- Daya Bay + RENO gives a weak preference for the normal ordering when combined with accelerator and atmospheric measurements
- JUNO is mounting a definitive test with reactor neutrinos using a 20 kton (!) LS detector → requires x2 improvement in E scale systematic uncertainty over KamLAND
- JUNO will also make precision measurements of all oscillation parameters
- Currently in commissioning!





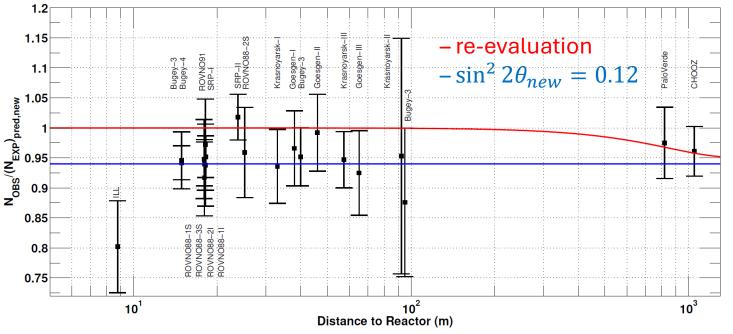


	PDG2023 [4]	JUNO 6 y
Δm ² ₂₁	7.53·10 ⁻⁵ eV ² ± 2.4%	0.3%
Δm ² ₃₁	2.5283·10 ⁻³ eV ² ± 1.3%	0.2%
sin²θ ₁₂	0.307 ± 4.2%	0.5%
$\sin^2 \theta_{13}$	0.022 ± 3.2%	12.1%

Malyshkin, Nu '24

The Reactor Antineutrino Anomaly

• Re-evaluation of reactor antineutrino spectra combining global nuclear data with fission beta spectral measurements revealed a global suppression relative to expectations



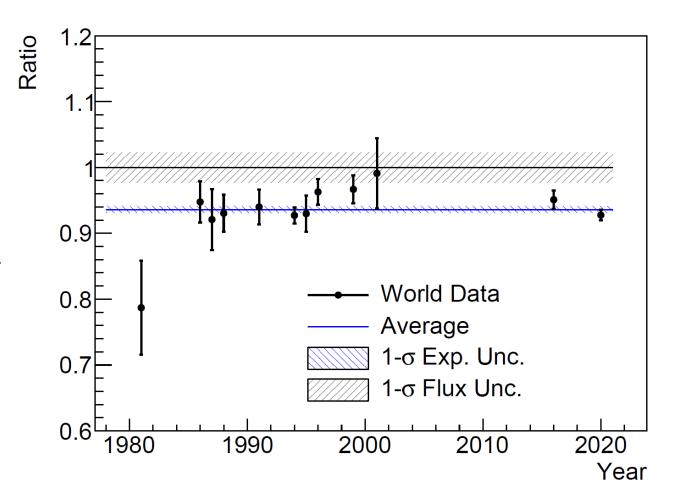
PRD 83, 073006 (2011)

The Reactor Antineutrino Anomaly

- Re-evaluation of reactor antineutrino spectra combining global nuclear data with fission beta spectral measurements revealed a global suppression relative to expectations
- Reactor antineutrino measurements?
- Nuclear data / nuclear theory?
- Beta spectrum measurements?
- Sterile neutrinos??

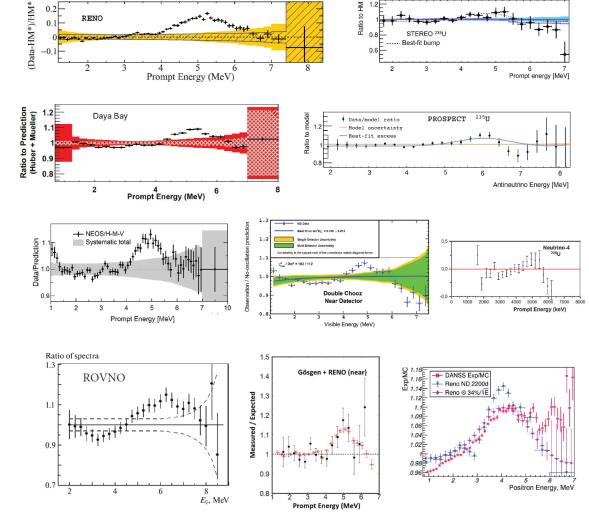
- Reactor antineutrino measurements
- Nuclear data / nuclear theory?
- Beta spectrum measurements?
- Sterile neutrinos??

All reactor data is consistent with a global average with 0.5% uncertainty!



- Reactor antineutrino measurements
- Nuclear data / nuclear theory?
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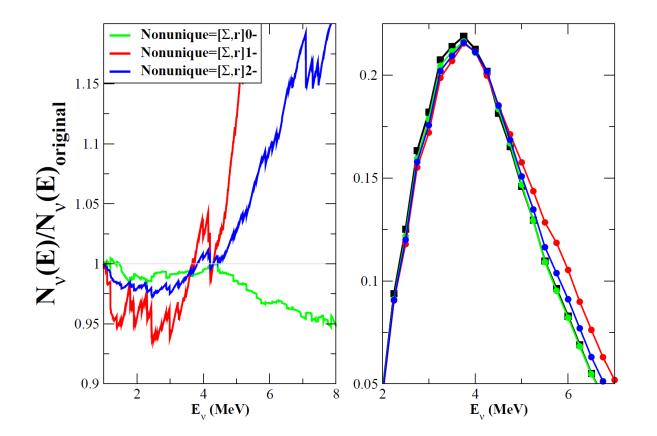
Reactor experiments also consistently show a "6 MeV bump" not in the model, agreeing with each other within uncertainties



Prog Part Nucl Phys 136, 104016 (2024)

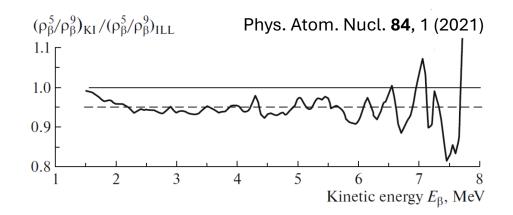
- Reactor antineutrino measurements
- Nuclear data / nuclear theory?
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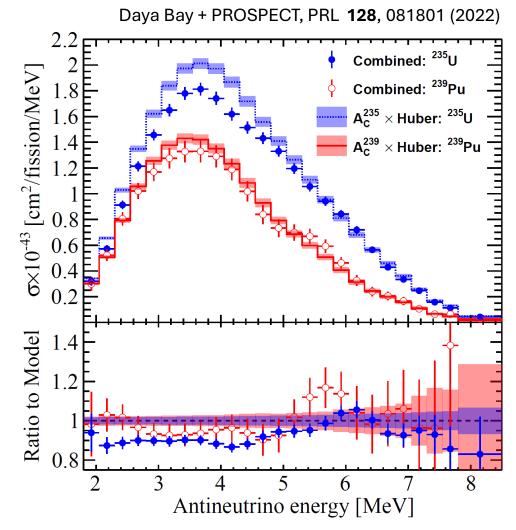
Theory uncertainties also appear to have been underestimated



- Reactor antineutrino measurements
- Nuclear data / nuclear theory?
- Beta spectrum measurements?
- Sterile neutrinos??

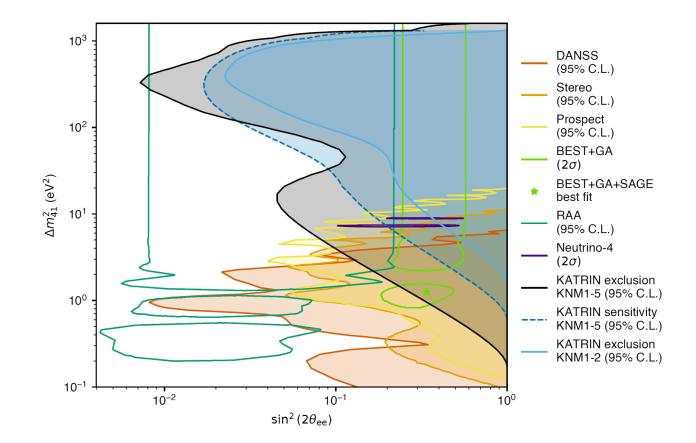
The beta spectra also appear to be questionable, especially for ²³⁵U



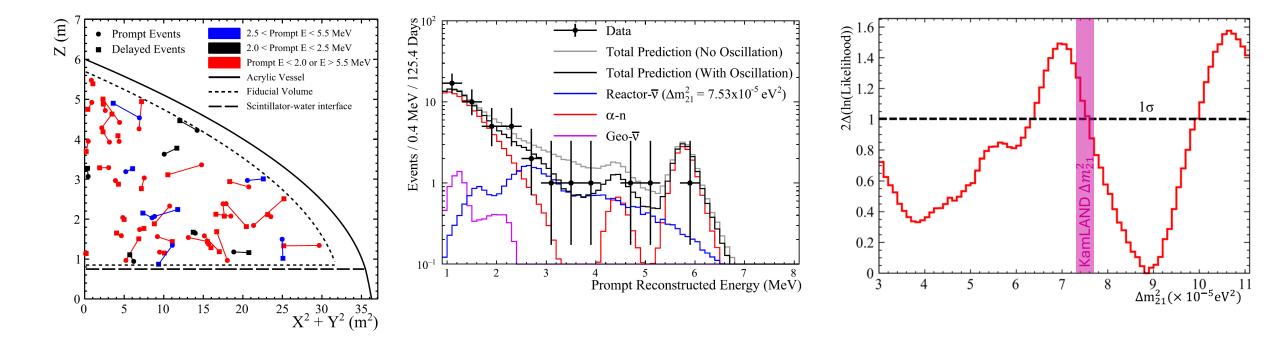


- Reactor antineutrino measurements
- Nuclear data / nuclear theory?
- Beta spectrum measurements?
- Sterile neutrinos??

Vigorous efforts to search for a neutrinobased solution remain inconclusive

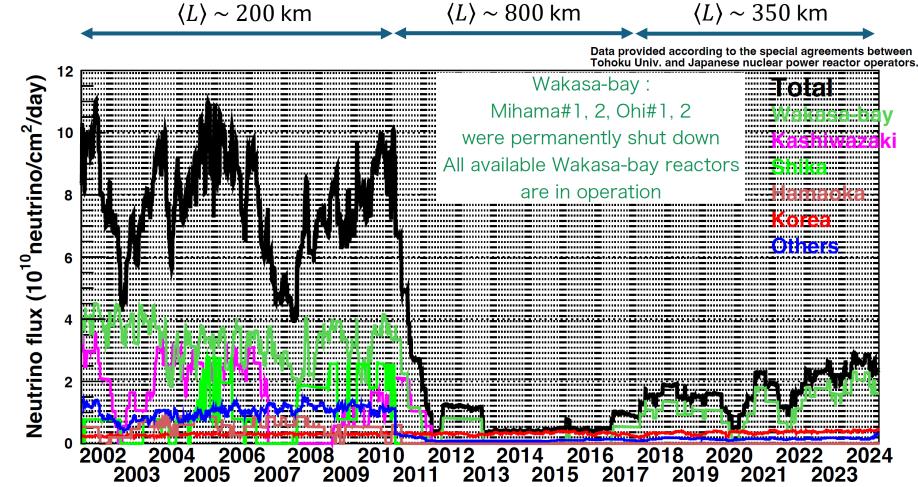


• SNO+: first qualitative test of Δm^2_{21}

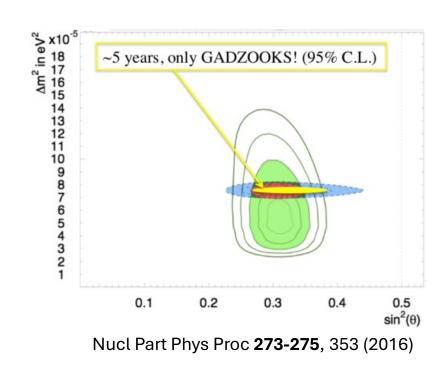


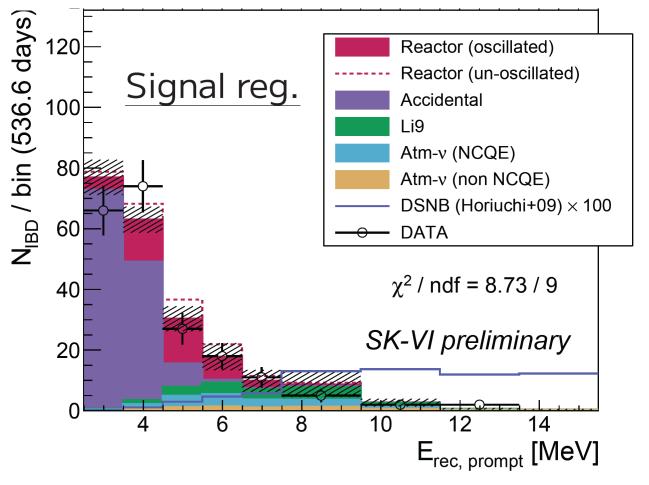
- SNO+
- KamLAND full-dataset:

measure Δm^2_{21} at 3 effective baselines



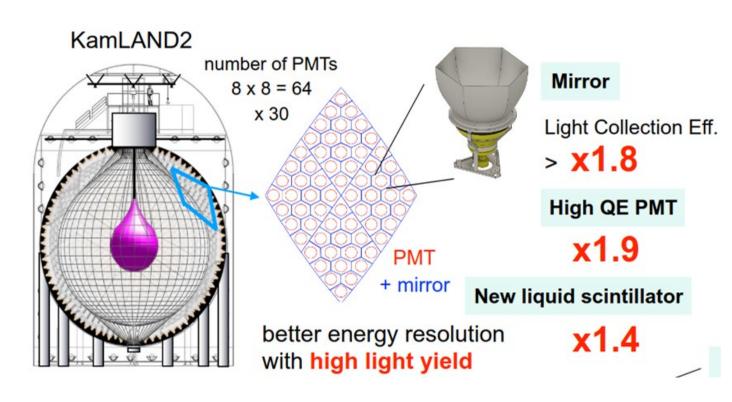
- SNO+
- KamLAND full-dataset
- GADZOOKS! (SK-Gd)





Izumiyama, Nu '24

- SNO+
- KamLAND full-dataset
- GADZOOKS! (SK-Gd)
- KamLAND2
- x2 resolution, lower bg, improved calibration (UH!)
 → qualitatively better Δm² measurement!



• Data taking start in 2027

- SNO+
- KamLAND full-dataset
- GADZOOKS! (SK-Gd)
- KamLAND2



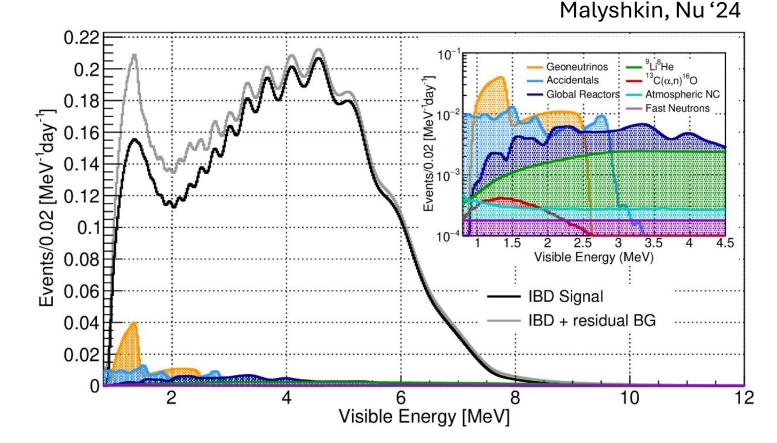


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 → qualitatively better Δm² measurement!
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- SNO+
- KamLAND full-dataset
- GADZOOKS! (SK-Gd)
- KamLAND2
- And of course: JUNO! (see Wei's and Benda's talks!)



Summary

- Reactors continue to reveal to us new secrets about the neutrino!
- A new generation of reactor experiments is poised to determine the mass ordering, measure neutrino oscillation with unprecedented precision, and reveal the source of the RAA
- Reactor neutrino measurement in general, and KamLAND in particular, rely on the techniques pioneered by John, and greatly benefitted from his contributions (including the non-technical ones!)

Happy birthday to John, my respected colleague and friend !



Special message to John from Atsuto Suzuki, speaking for all of us on KamLAND!